



Jet Propulsion Laboratory
California Institute of Technology

HysplRI Thermal Infrared L1/L2: Algorithm Maturity and Cal-Val

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Theoretical Basis: Planck's law and Temperature

$$B_{\lambda} = \frac{C_1}{\lambda^5 \left[\exp\left(\frac{C_2}{\lambda T_s}\right) - 1 \right]}$$

$$T_s = B_{\lambda}^{-1} \left(\frac{L_{\lambda}(\theta) - \rho_{\lambda} L^{\downarrow}}{\varepsilon_{\lambda}} \right)$$

where :

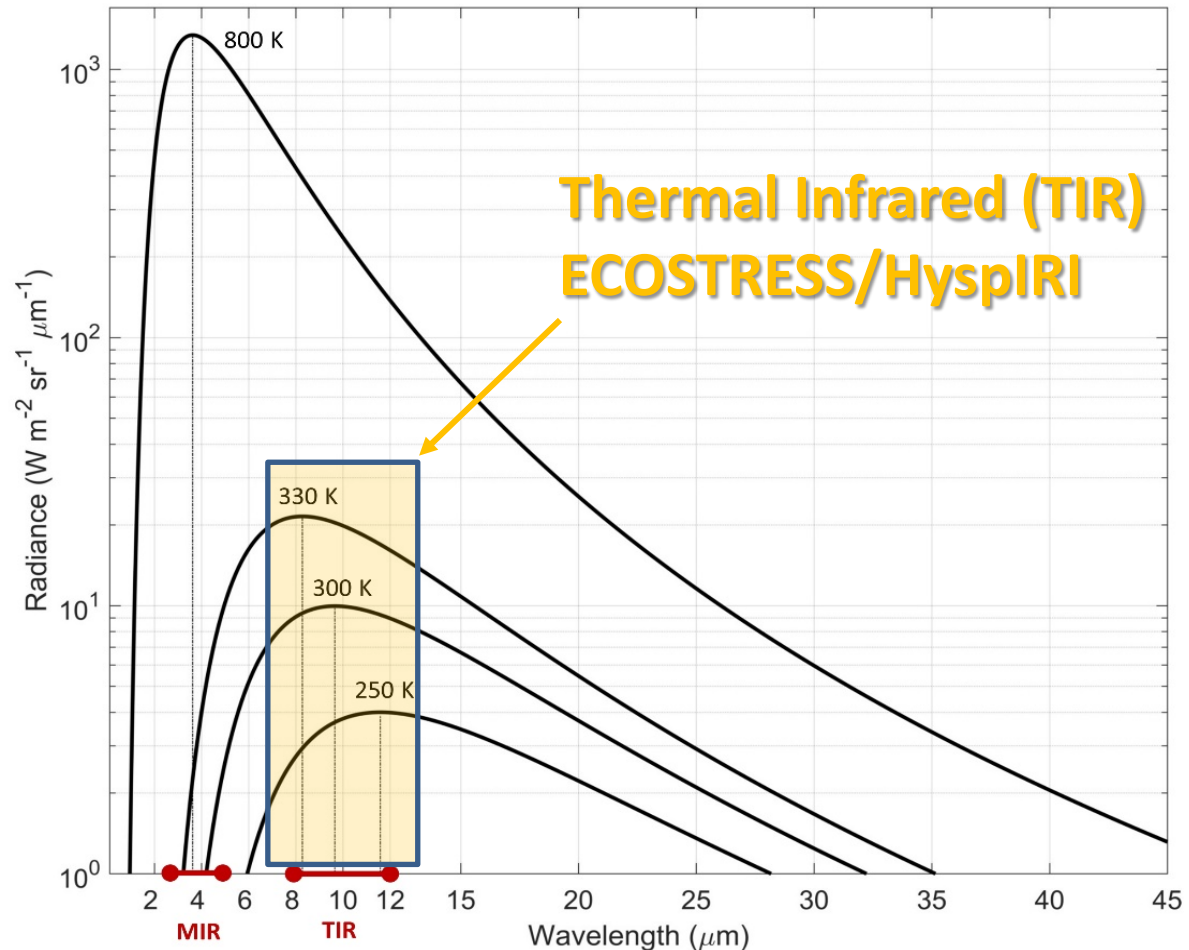
B_{λ} = blackbody spectral radiance

λ = wavelength

T_s = Surface Temperature

C_1 = first radiation constant

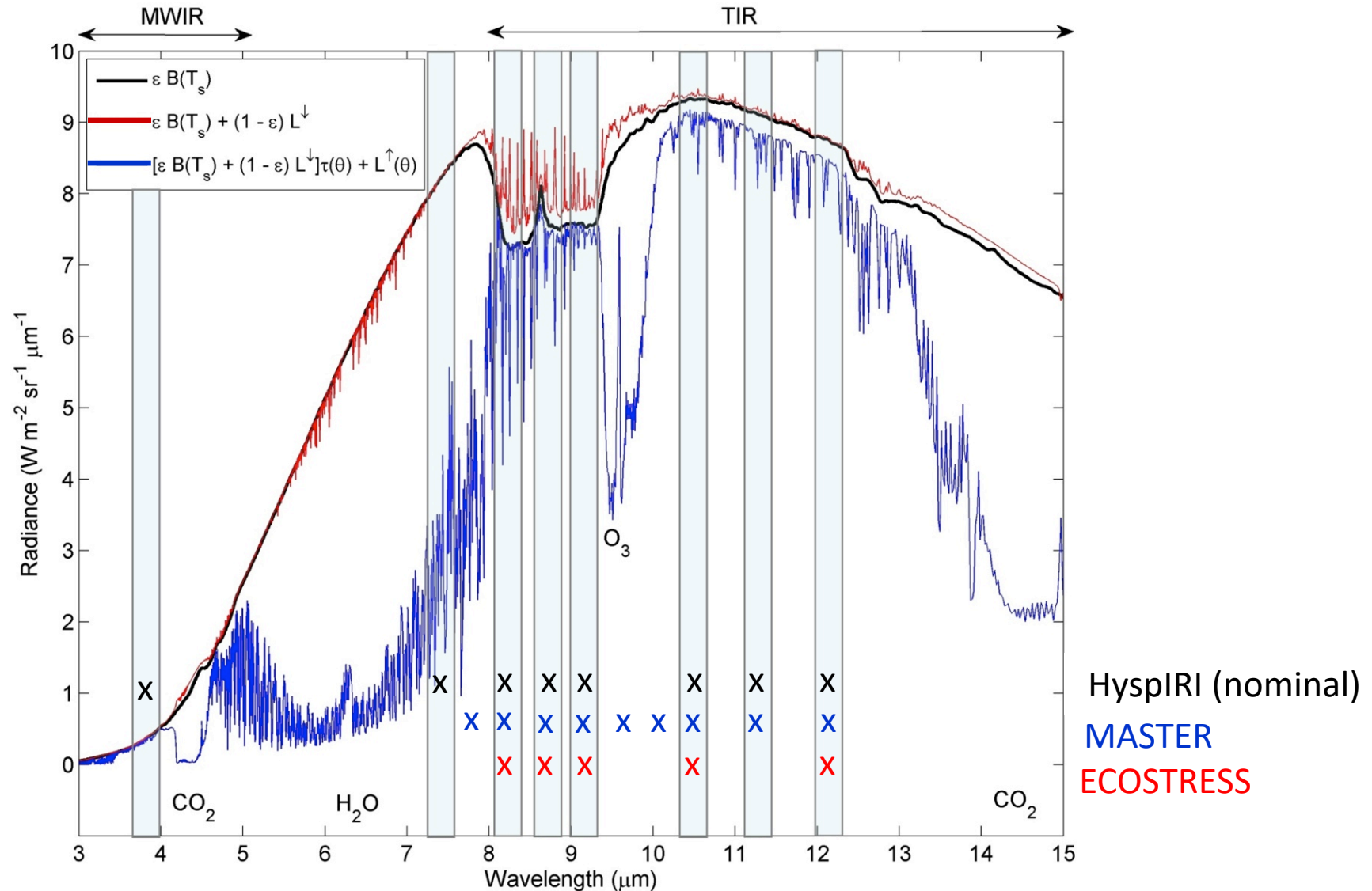
C_2 = second radiation constant



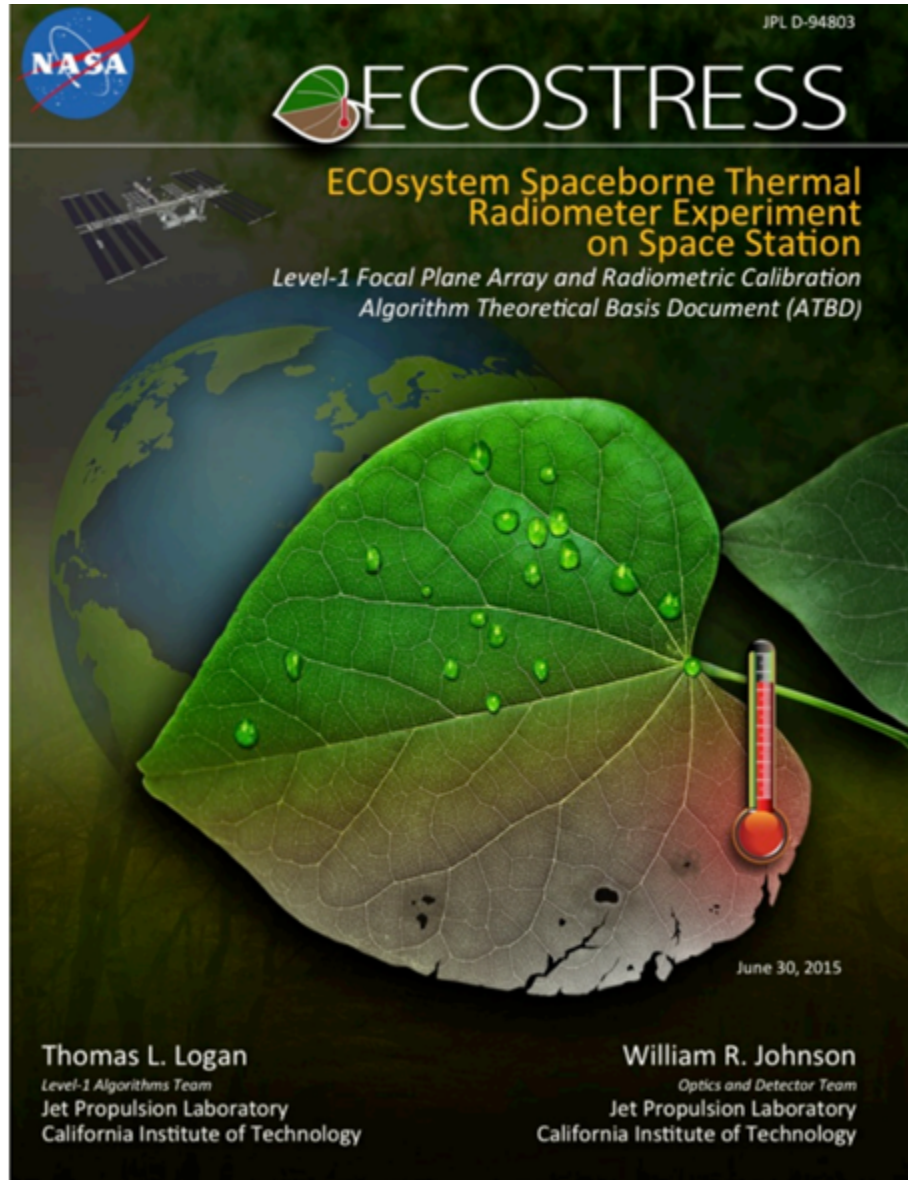
As the temperature increases the peak in the Planck function shifts to shorter and shorter wavelengths

Thermal Infrared Radiative Transfer

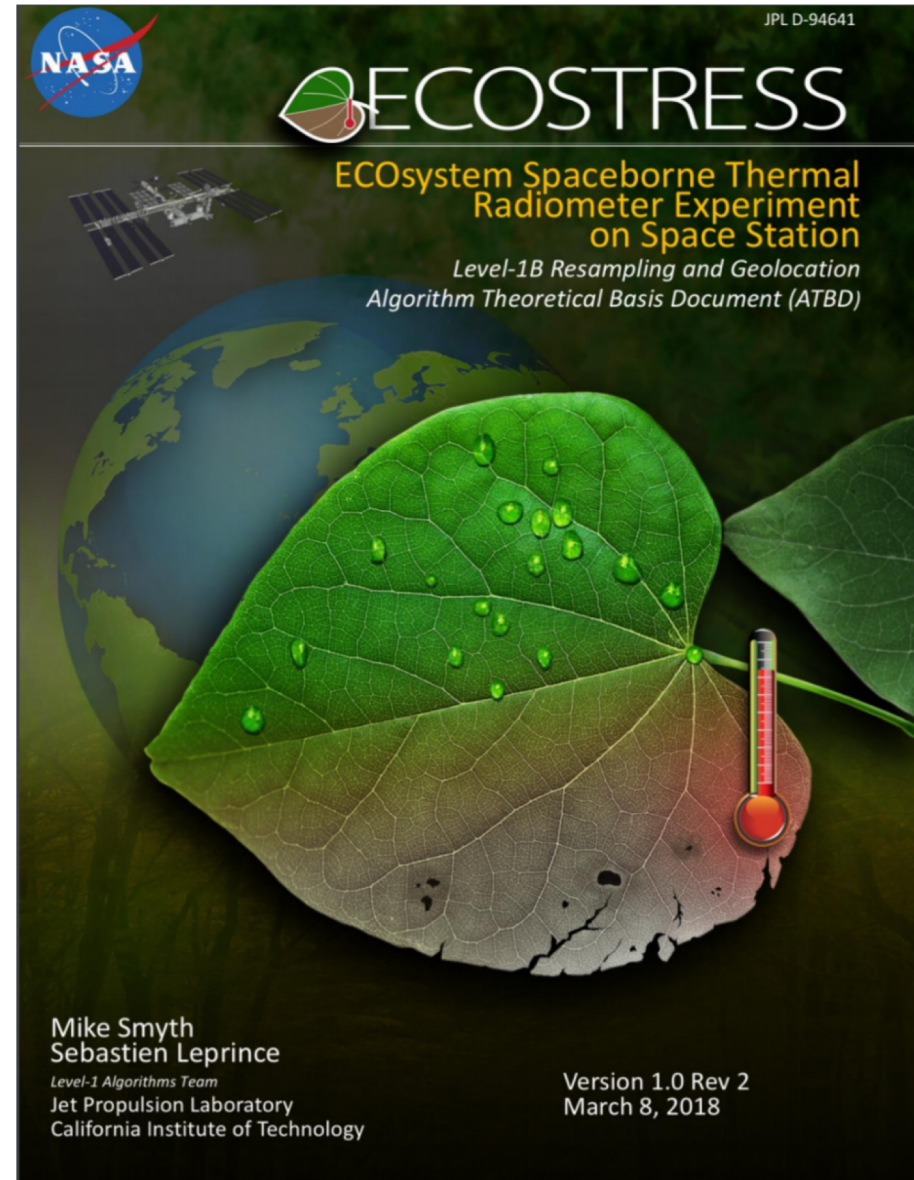
$$L_{sat,\lambda} = [\varepsilon_{\lambda} B_{\lambda}(LST) + (1 - \varepsilon_{\lambda}) L_{sky,\lambda}^{\downarrow}] \tau_{\lambda} + L_{sky,\lambda}^{\uparrow}$$



L1A ATBD



L1B ATBD

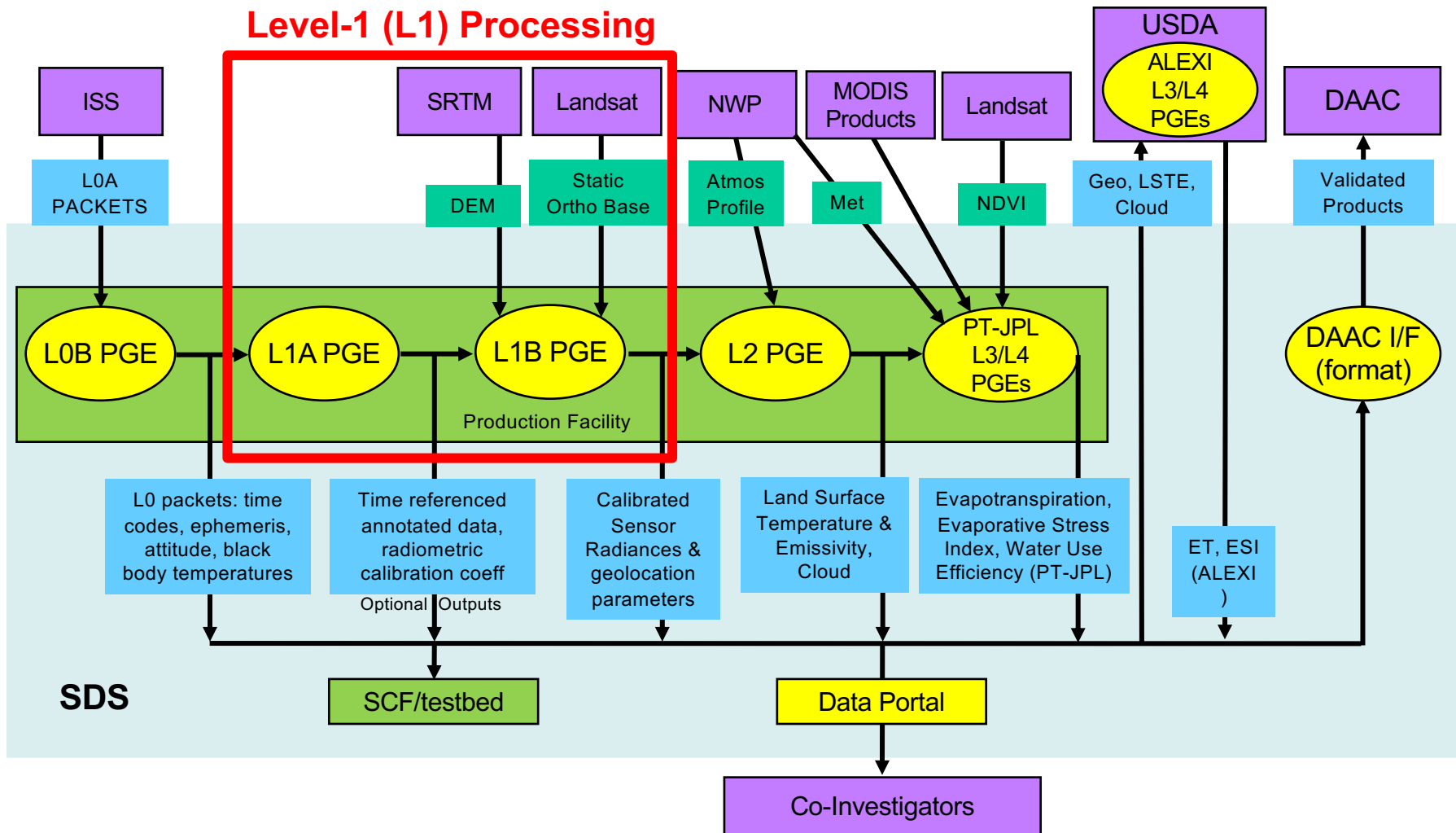


L1 TIR Algorithm Overview

- L1 Processing consists of four PGEs (Product Generation Executable)
 - L1A
 - Raw Data Processing
 - Reformat incoming data packets, metadata, and ancillary data
 - Radiometric Calibration
 - Convert Image Pixel DNs to Radiance Coefficients
 - L1B
 - Resampling and Remapping
 - Merge Focal Plane overlap and average pixels (lines) to improve signal.
 - Geolocation
 - Initial Map Projection from ISS Ephemeris and Pointing data
 - Geolocation Matching (using Landsat orthobase) to correct for Positional Errors

L1 Overview - ECOSTRESS

L1 in the SDS Processing Flow

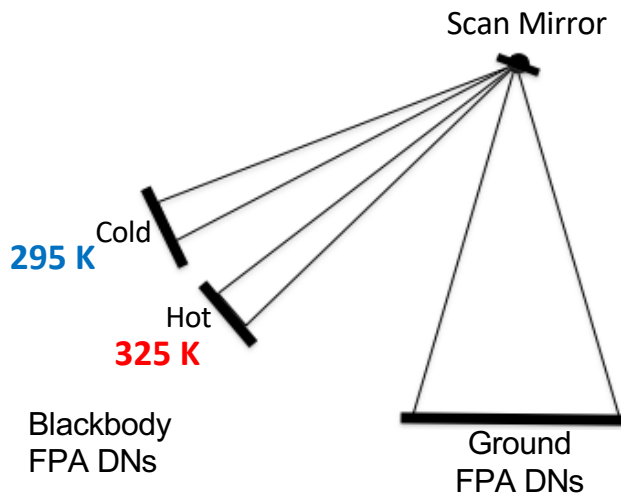


L1A Radiometric Calibration PGE

L1A Radiometric Two-Point Calibration*

Approach

- Read BB Temperatures.
- Create synthetic FPA 256x1 BB Temperature Images.
- Convert FPA BB Images to Radiances
- Collect FPA BB and Ground DNs.
- Apply 2pt Algorithm →



$$R_{\lambda} = a + bD_{\lambda}$$

$$a = \frac{R_h D_c - R_c D_h}{D_c - D_h} \quad b = \frac{R_c - R_h}{D_c - D_h}$$

Where:

R = Calculated Radiance of an input Digital Number (DN)

a = Offset Term

b = Gain Term

D = Input Earth Digital Number (DN)

R_c = Radiance of the Cold Blackbody (Section 3.3.2)

R_h = Radiance of the Hot Blackbody (Section 3.3.2)

D_c = Digital Number (DN) from the Cold Blackbody Calibration File (Section 3.3.3)

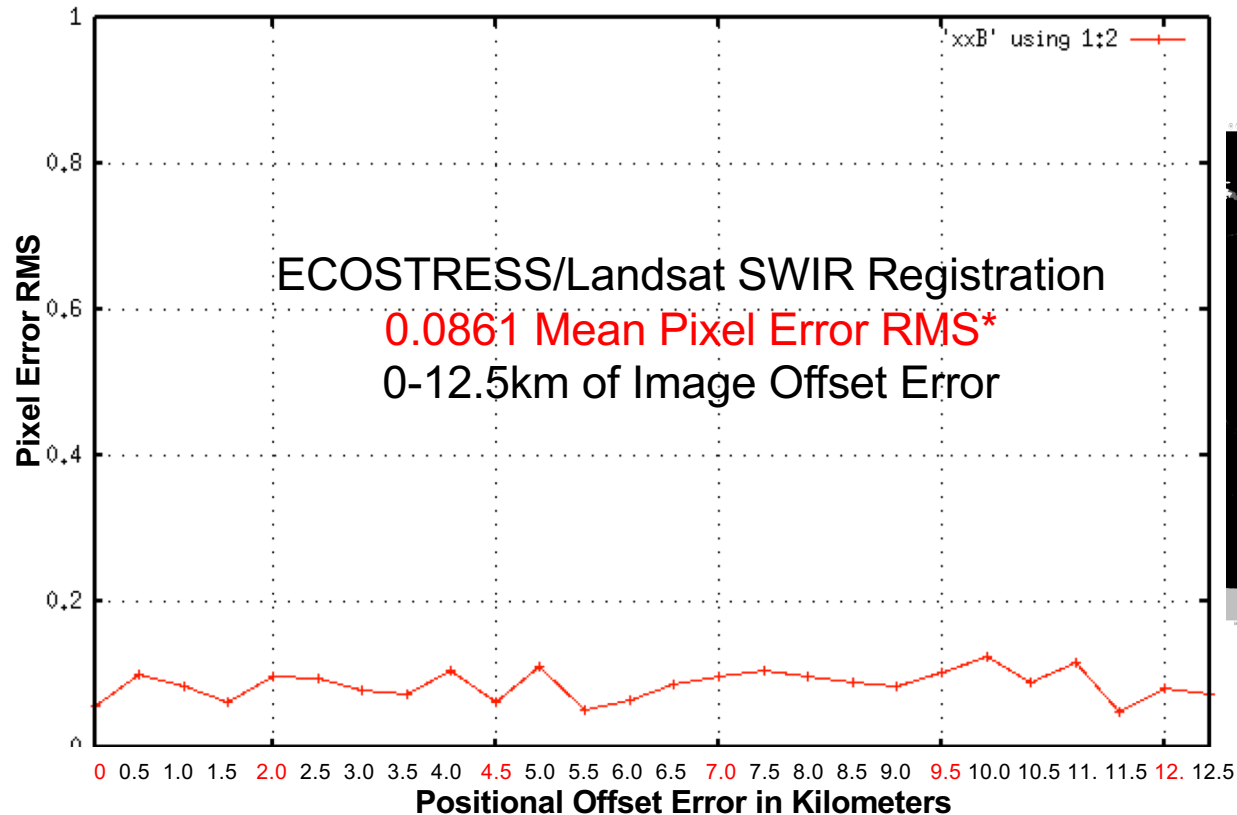
D_h = Digital Number (DN) from the Hot Blackbody Calibration File (Section 3.3.3)

Demonstrated accuracy at <0.5 K

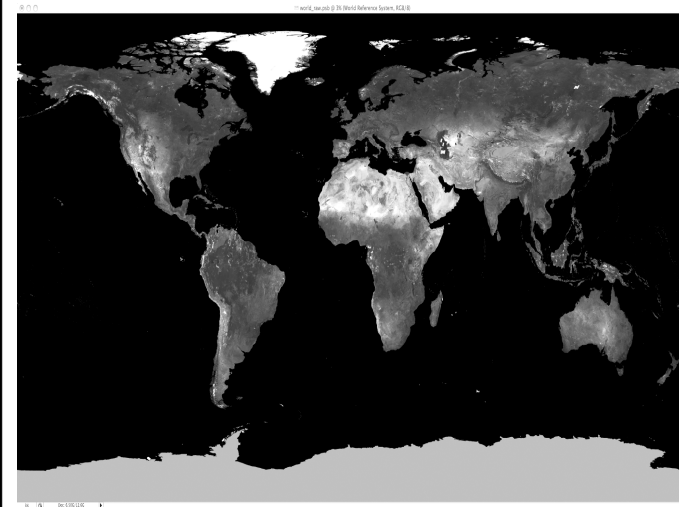
*Documented in: "Level-1 Focal Plane Array and Radiometric Calibration Algorithm Theoretical Basis Document (ATBD)," JPL D-94803.

L1B Geolocation Testbed

SWIR 1.6um Co-Registration Test Results (0-12.5km Offset)

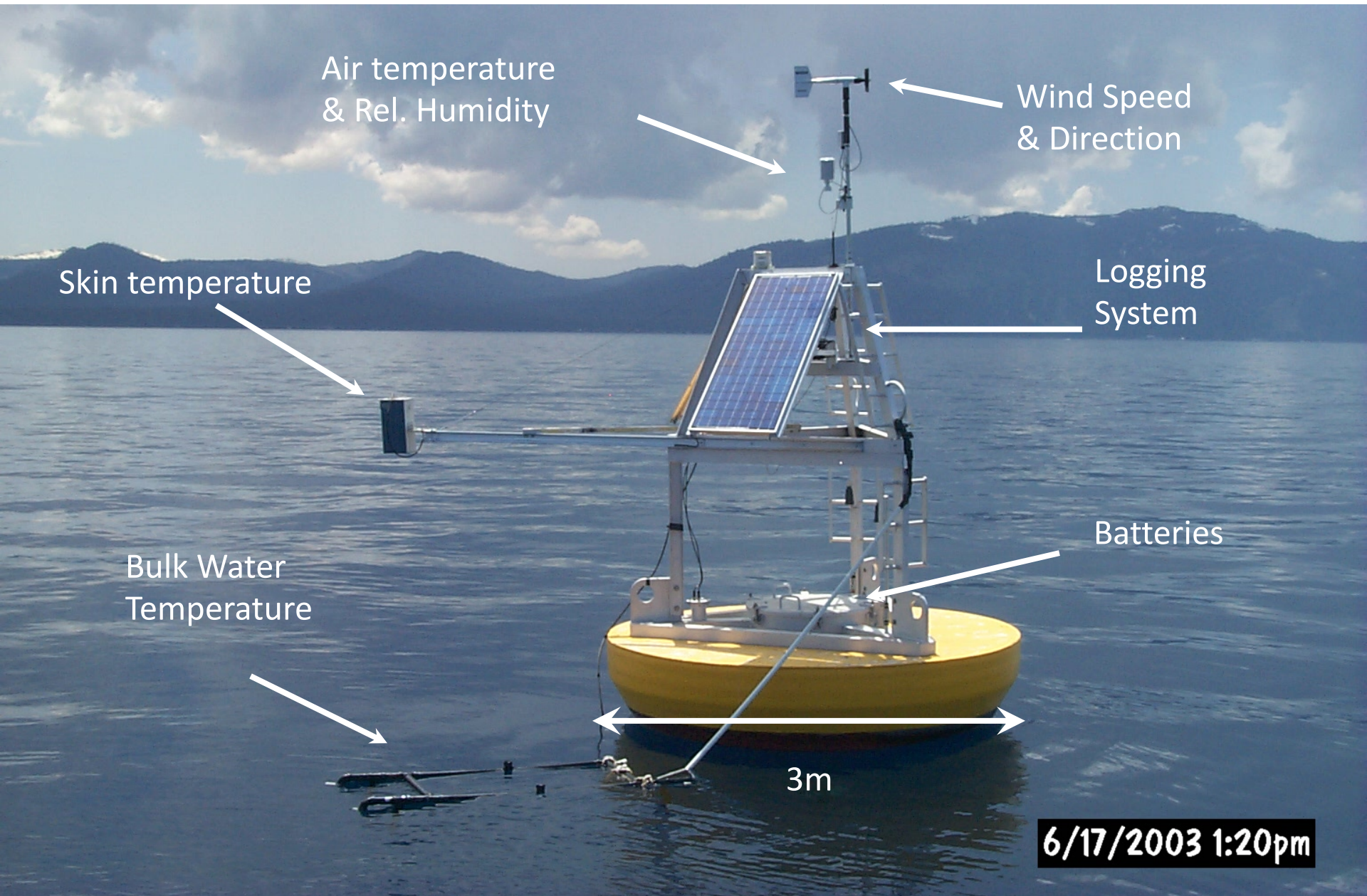


Landsat 7 Global Ortho-Base SWIR/TIR



*Using Pre-Launch AFIDS beta code

Lake Tahoe Cal/Val

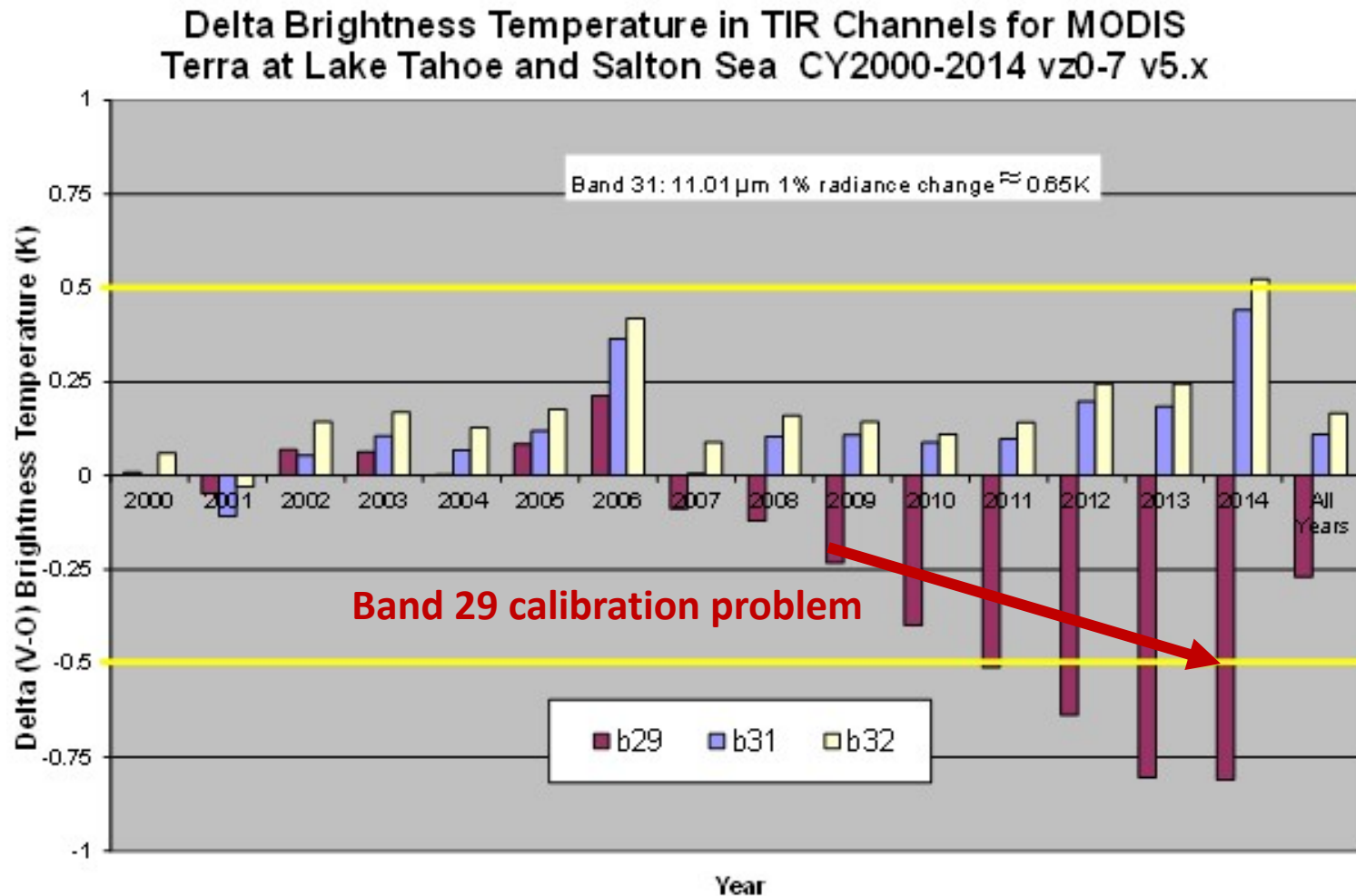


Radiometer Calibration - JPL



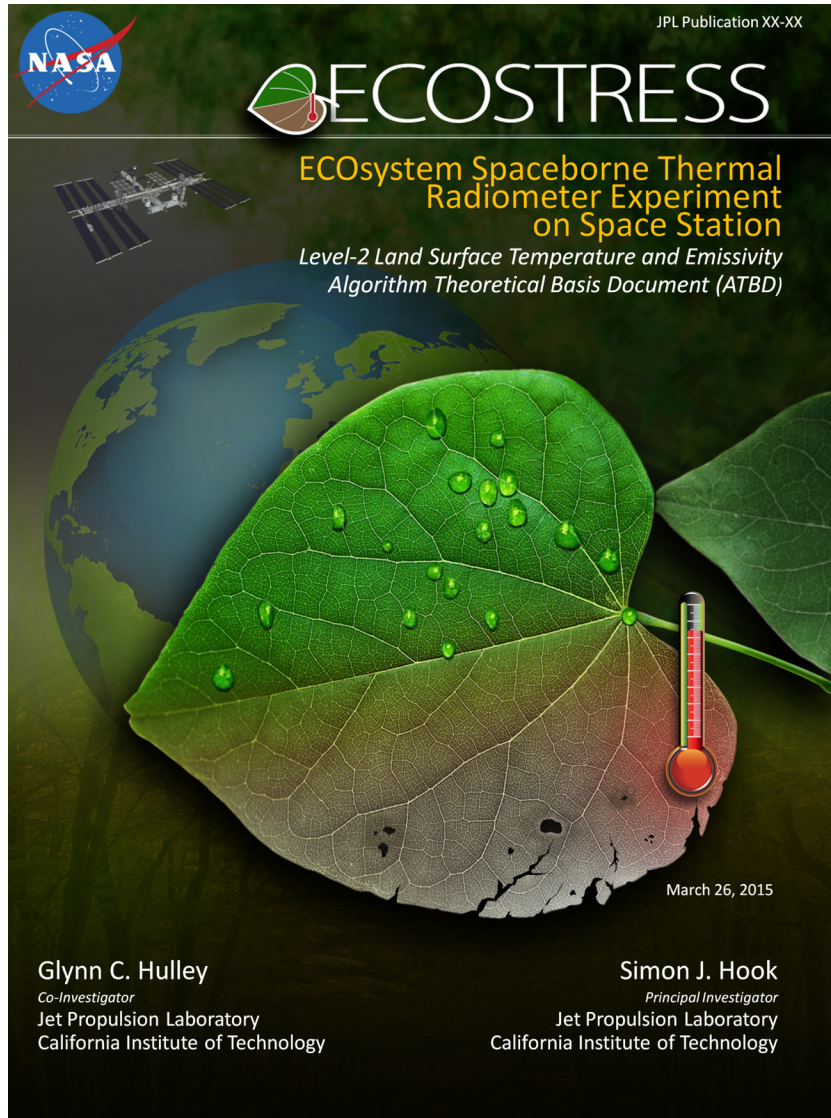
Brightness Temperature at Sensor : On-orbit Validation

Example results from MODIS



ECOSTRESS L2 ATBD

<https://ecostress.jpl.nasa.gov/products/>



HyspIRI L2 ATBD

<https://hyspiri.jpl.nasa.gov/documents>

JPL Publication 11-5



HyspIRI Level-2 Thermal Infrared (TIR) Land Surface Temperature and Emissivity Algorithm Theoretical Basis Document

G. Hulley
S. Hook
Jet Propulsion Laboratory

National Aeronautics and
Space Administration

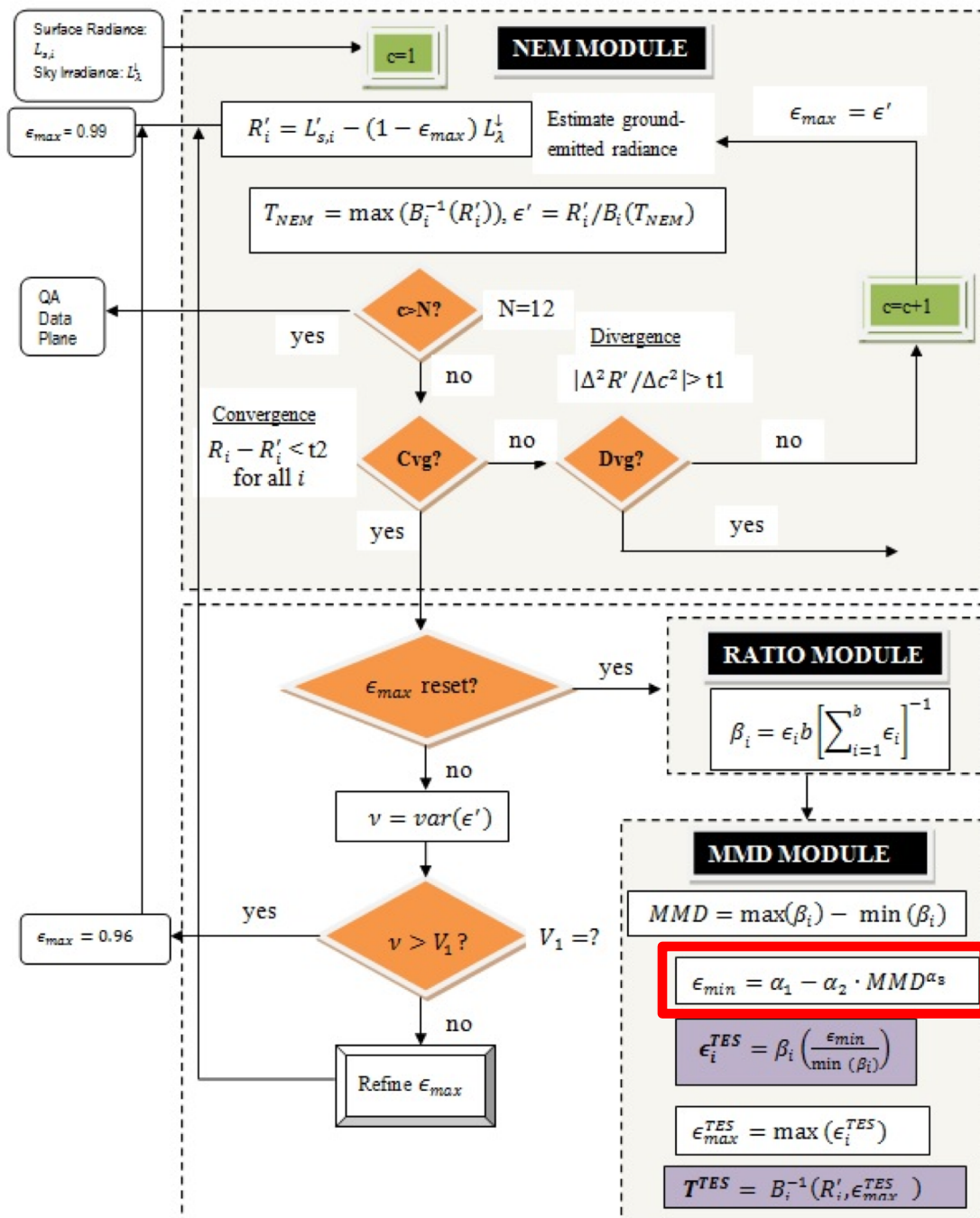
Jet Propulsion Laboratory
California Institute of Technology
Pasadena, California

May 2011

L2 Temperature Emissivity Separation (TES) Heritage

LST&E Product Characteristics	ASTER <i>Terra - 2000</i>	MODIS <i>Terra – 2000 Aqua - 2002</i>	VIIRS <i>NPP - 2011</i>	ECOSTRESS <i>ISS - 2018</i>
Core Algorithm	TES*	TES*	TES*	TES*
Product	AST05/08 - 2000	MOD21	VNP21	LSTE - 2018
Bands used	10 (8.3 μm) 11 (8.6 μm) 12 (9.1 μm) 13 (10.6 μm) 14 (11.3 μm)	29 (8.55 μm) 31 (11 μm) 32 (12 μm)	14 (8.55 μm) 15 (10.76 μm) 16 (12 μm)	1 (8.3 μm) 2 (8.6 μm) 3 (9.1 μm) 4 (11.2 μm) 5 (12 μm)
Radiative Transfer Model	MODTRAN	RTTOV	RTTOV	RTTOV
Atmospheric Profiles	NCEP/MOD07	MERRA-2 WVS model	MERRA-2 WVS model	GEOS5 WVS model
Spatial and Temporal Resolution	90 m 16-day	1-km Twice daily (am/pm)	750 m Twice daily (am/pm)	70 m Variable (CONUS – 5 day)
Science Data Products	- LST - Emissivity (bands 10-14)	- LST - Emissivity (bands 29, 31, 32)	- LST - Emissivity (bands 14, 15, 16)	- LST - Emissivity (bands 1-5)

*TES = Temperature Emissivity Separation Algorithm (Gillespie 1998, Hulley et al. 2012)



TES Algorithm

LST&E Products using TES:

- ASTER 05/08
- MOD21 C6
- ASTER GED
- MASTER
- HyTES
- ECOSTRESS LSTE
- VIIRS VNP21

Temperature Emissivity Separation (TES) Algorithm

T-E separation is under-determined (ill-posed):

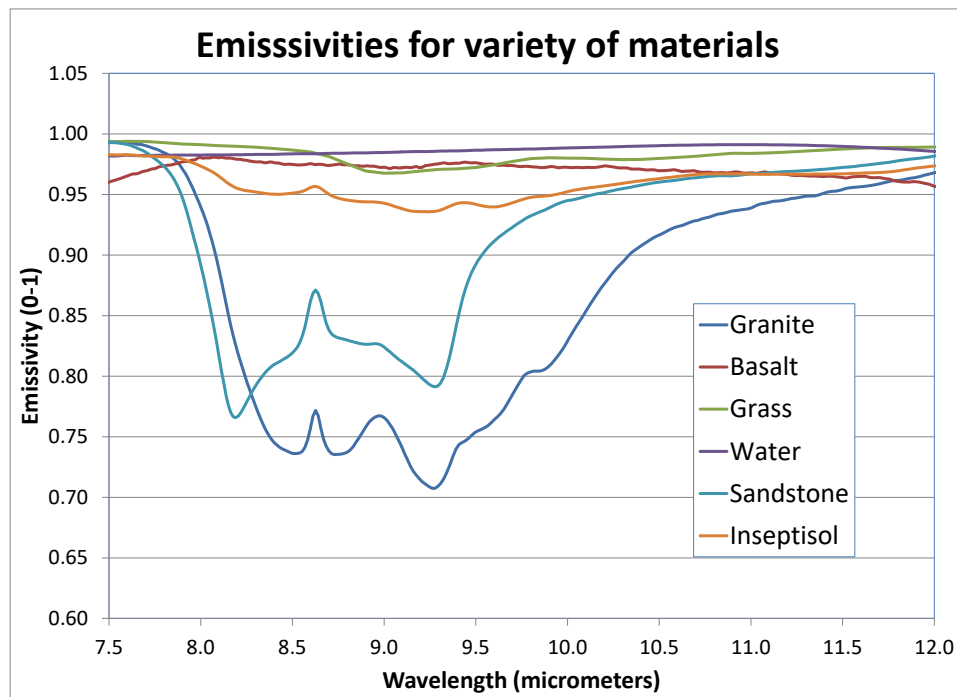
If have N measurements, always have $N+1$ unknowns:

Radiance Band 1 = $T + \text{emissivity}_1$

Radiance Band 2 = $T + \text{emissivity}_2$

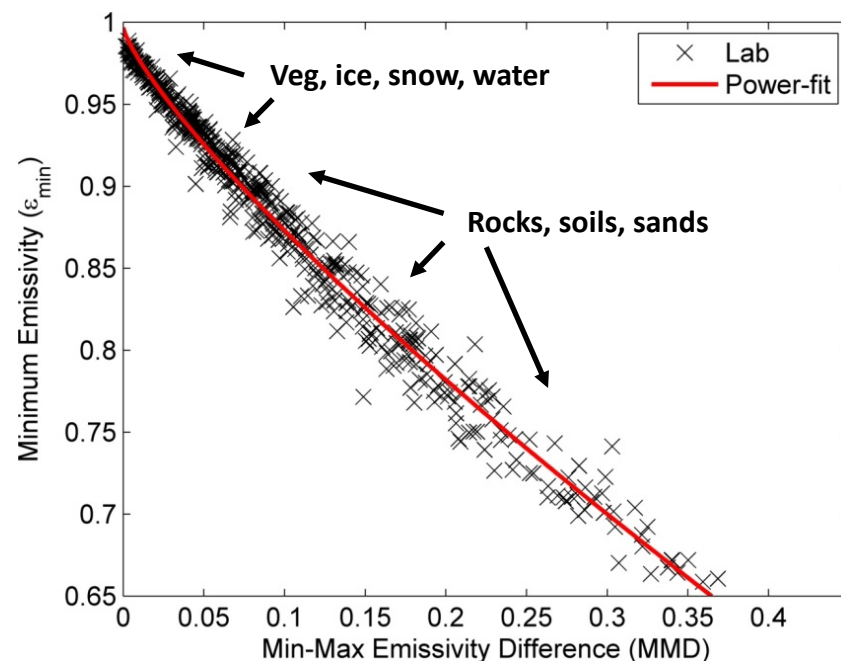
Radiance Band 3 = $T + \text{emissivity}_3$

.....






Emissivity Calibration Curve

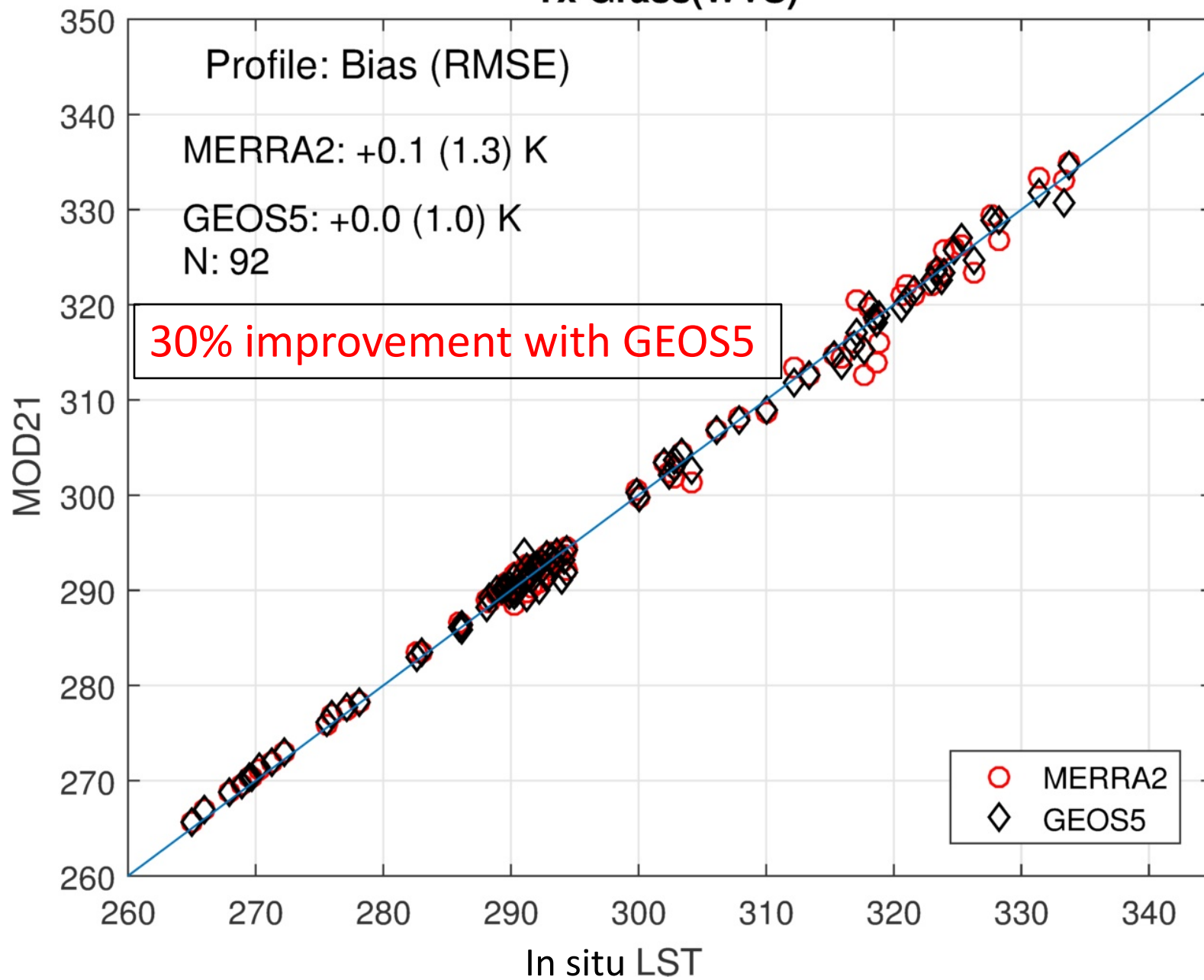
$$\varepsilon_{\min} = 0.994 - 0.687 \cdot MMD^{0.737}$$



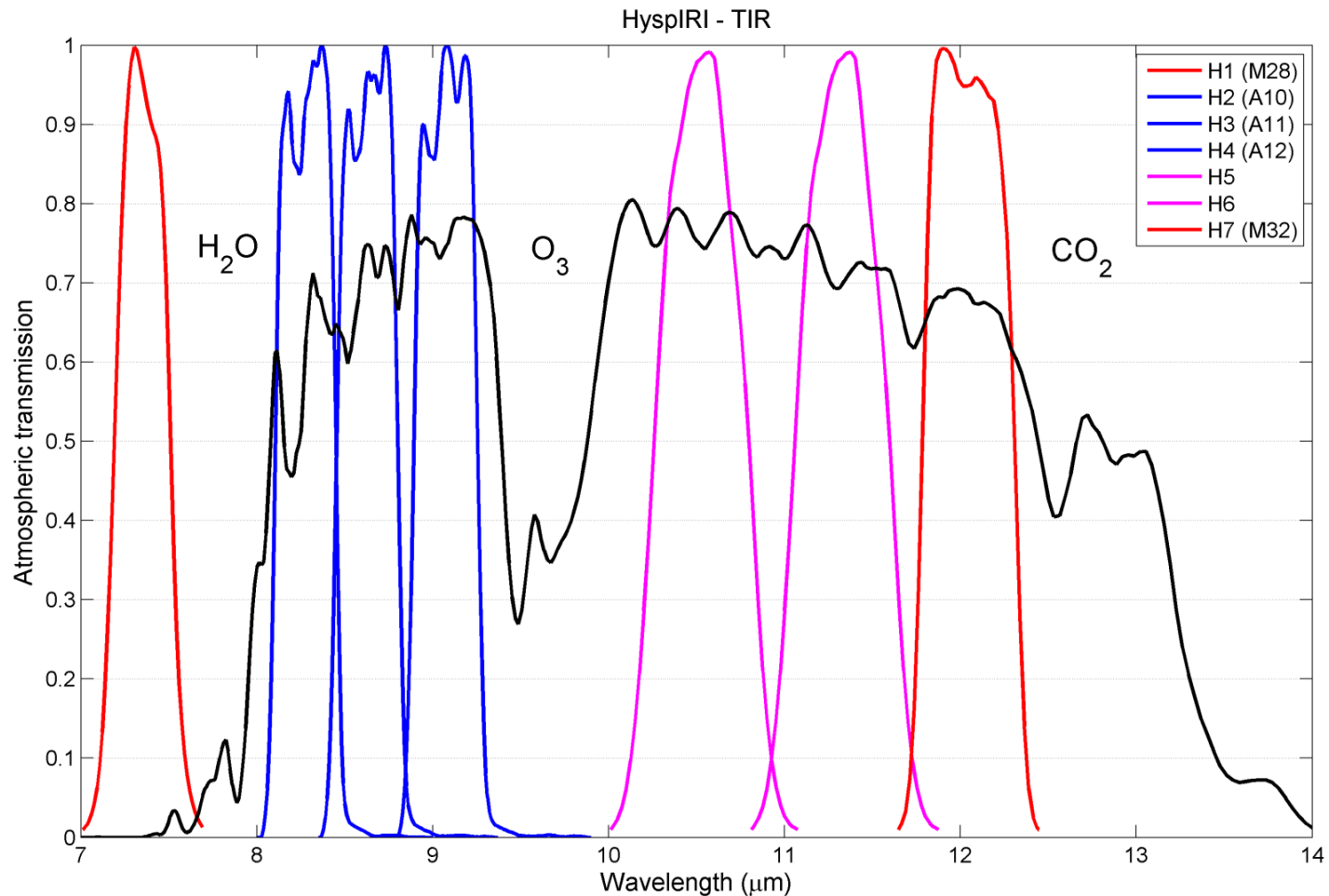
Atmospheric correction: NASA NWP models

NASA GMAO	MERRA-2	GEOS-5 FP
Description	Reanalysis from 1979-present	Analyses and forecasts in near real-time using recent GEOS-5 model
Spatial Resolution	0.5 degree (~50 km) 	0.25 degree (~25 km)
Temporal Resolution	6 hourly 	3 hourly
Vertical Resolution	42 pressure levels	42 pressure levels
Geophysical Data	Temperature, humidity, pressure	Temperature, humidity, pressure
Latency	Monthly 	Daily

Tx Grass(WVS)



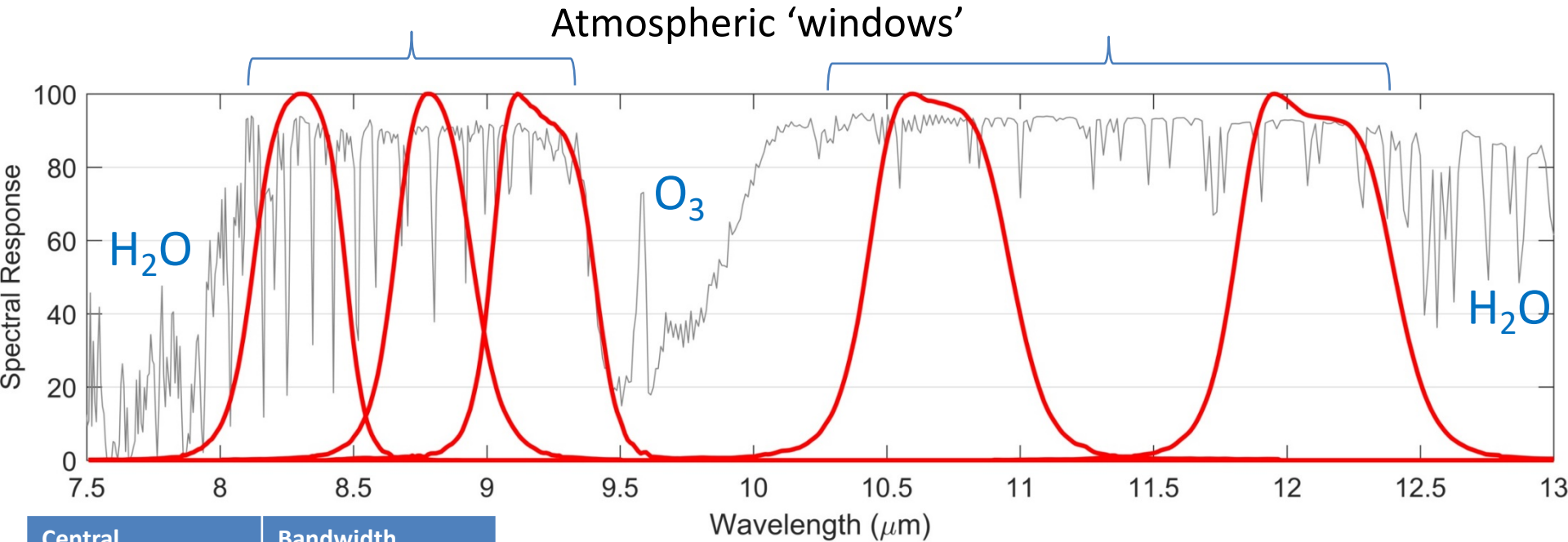
HyspIRI TIR Response Functions (Nominal)



HyspIRI TIR band study report – Ramsey, Realmuto, Hulley

https://hyspiri.jpl.nasa.gov/downloads/reports_whitepapers/

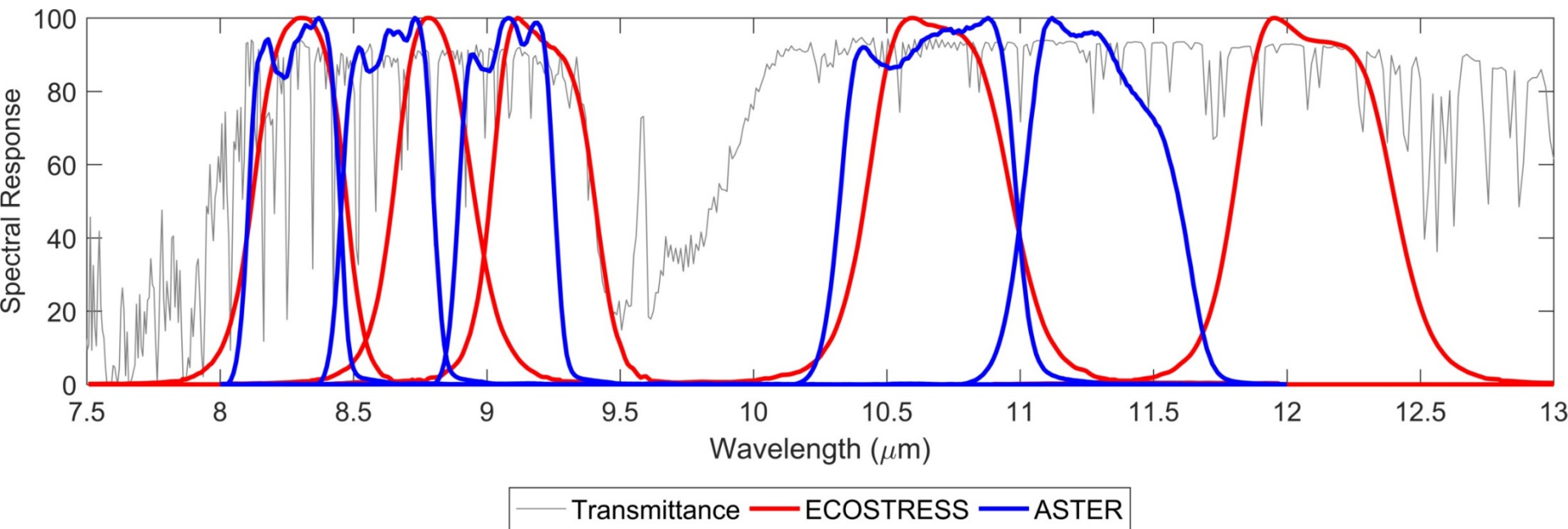
ECOSTRESS



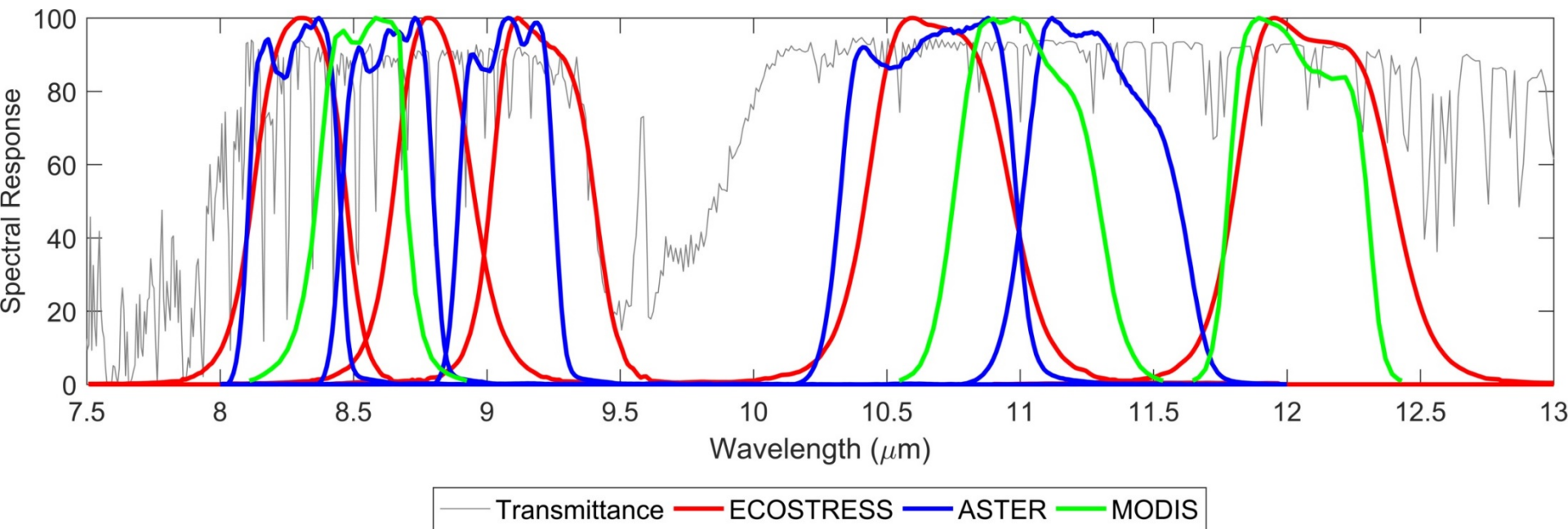
Central Wavelength (μm)	Bandwidth (μm)
8.29	0.355
8.80	0.309
9.20	0.395
10.6	0.553
12.09	0.610

— Transmittance — ECOSTRESS

Spectral Response Functions



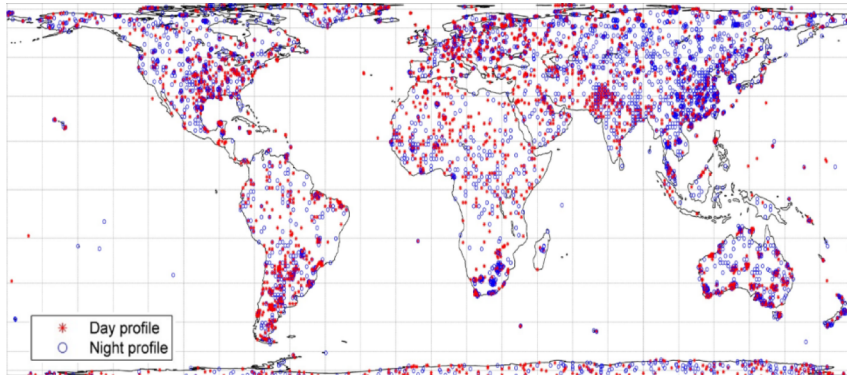
Spectral Response Functions



LST Uncertainty Analysis – number of bands

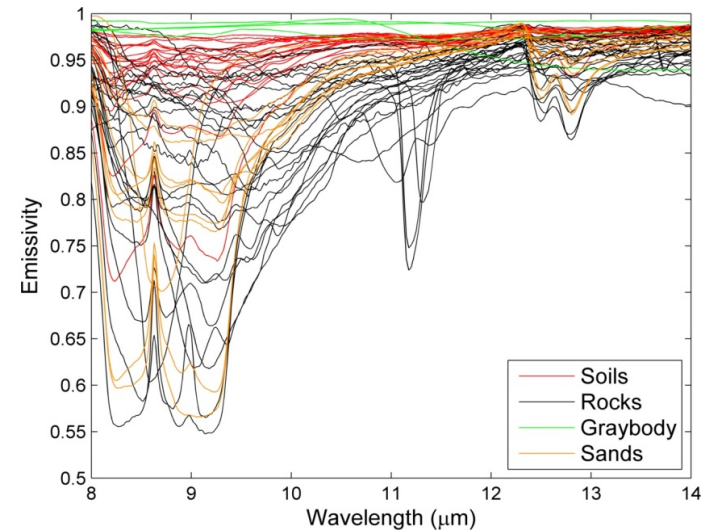
<i>Hulley et al. 2012 (Uncertainty Analysis Study)</i>			LST Uncertainty (K)			
Surface types	Samples	MODTRAN Simulations	2 Bands Split-Window Landsat 8, MODIS, VIIRS	3 bands TES MODIS, VIIRS	5 bands TES ASTER, ECOSTRESS	6/7 bands TES HypsIRI/SBG
Water, vegetation soils, rocks, sands	111	9,158,832	2.66 K	1.49 K	1.13 K	1.05 K

SeeBor Global Radiosonde database
(15,704 profiles)



+

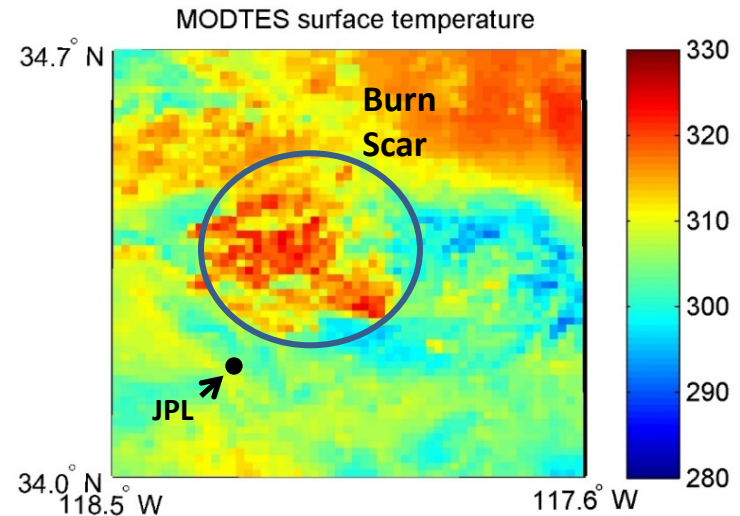
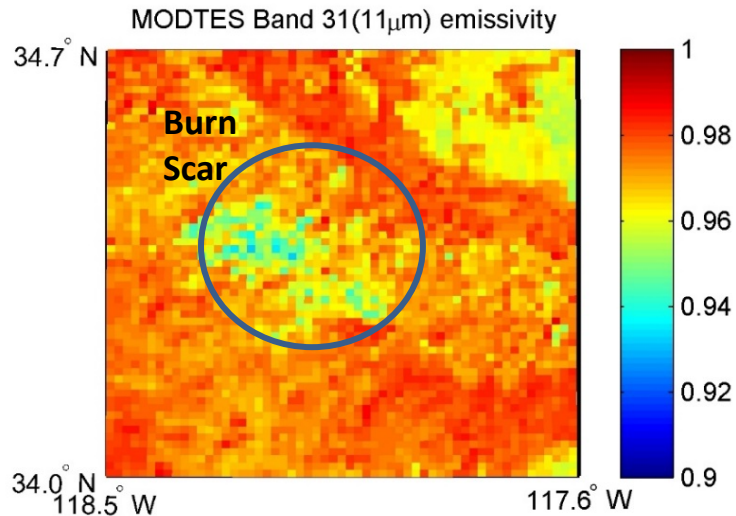
150 Surface types (ASTER speclib)



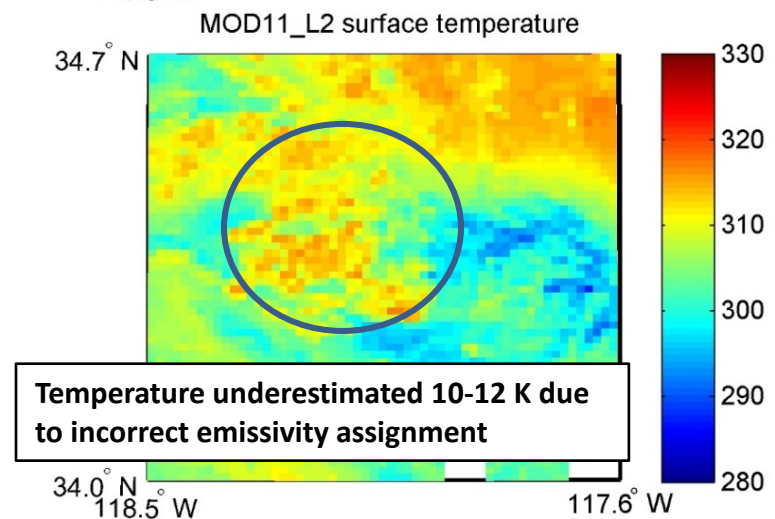
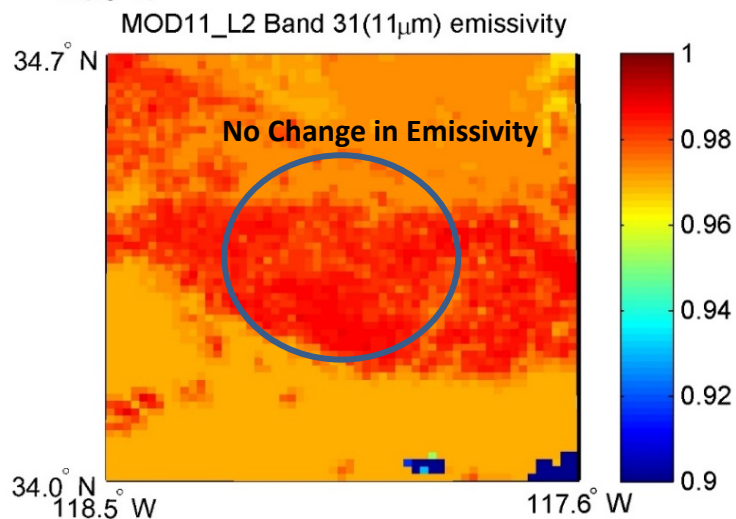
Station Fire: Multispectral capability



MOD21
(TES) →

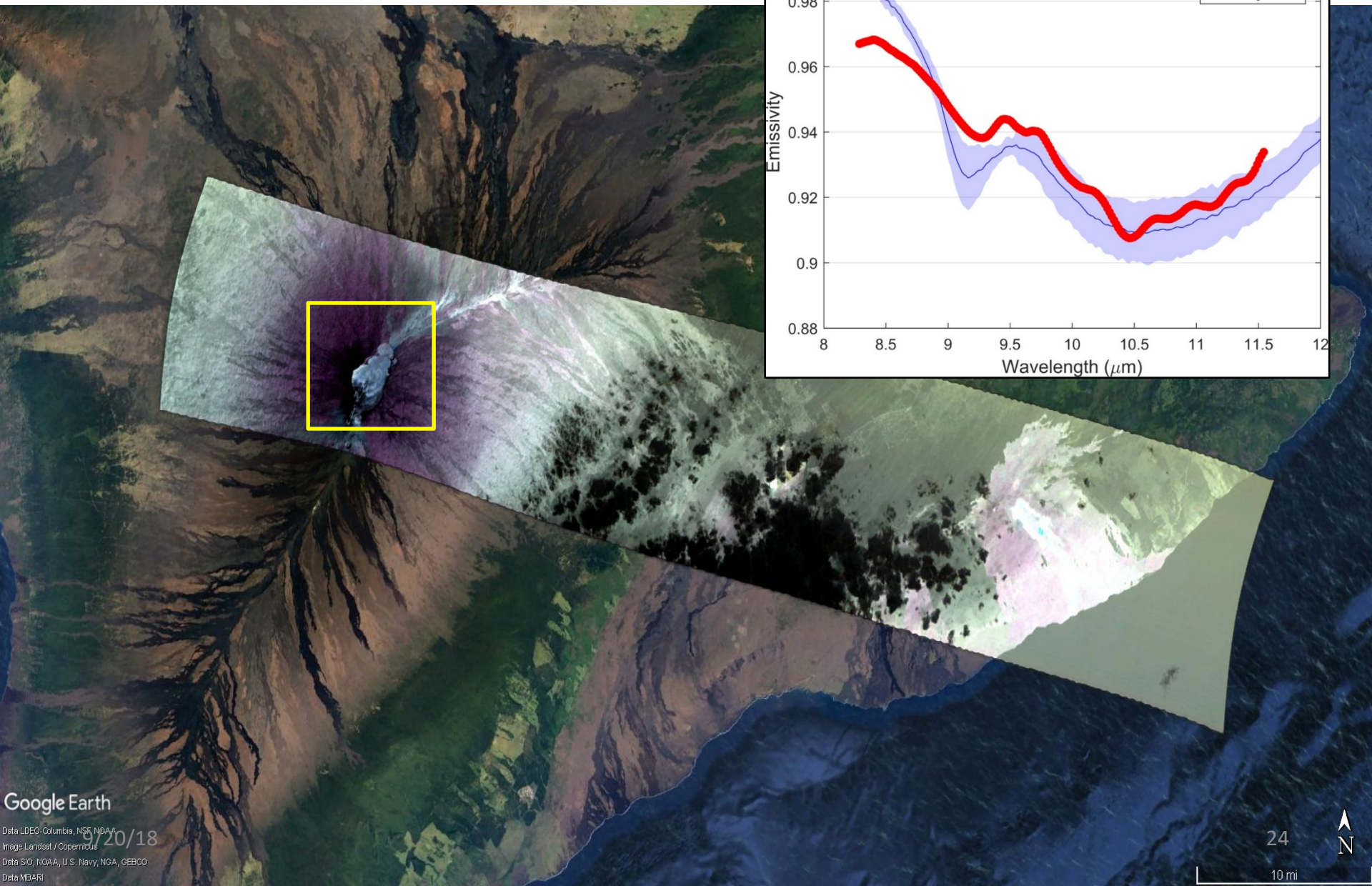
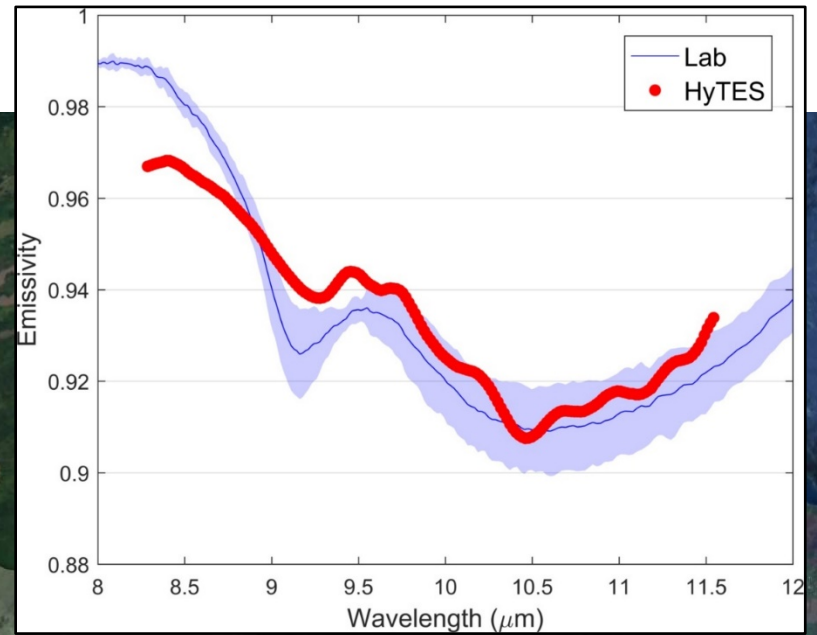


MOD11
(2-band) →

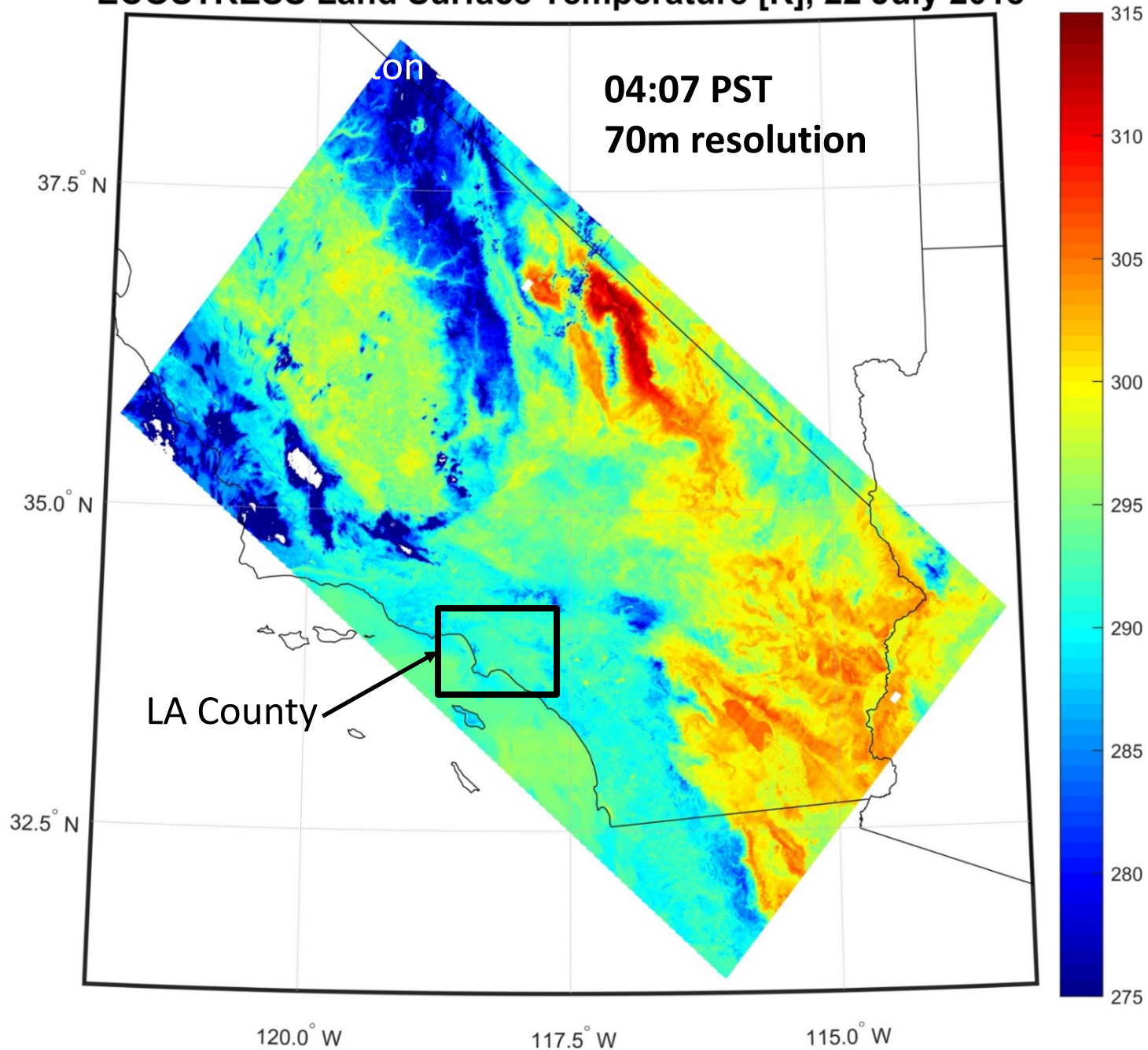


HyTES

20180130t205342_HawaiiHI



ECOSTRESS Land Surface Temperature [K], 22 July 2018



ECOSTRESS LST, 22 July 2018, 04:07 PST

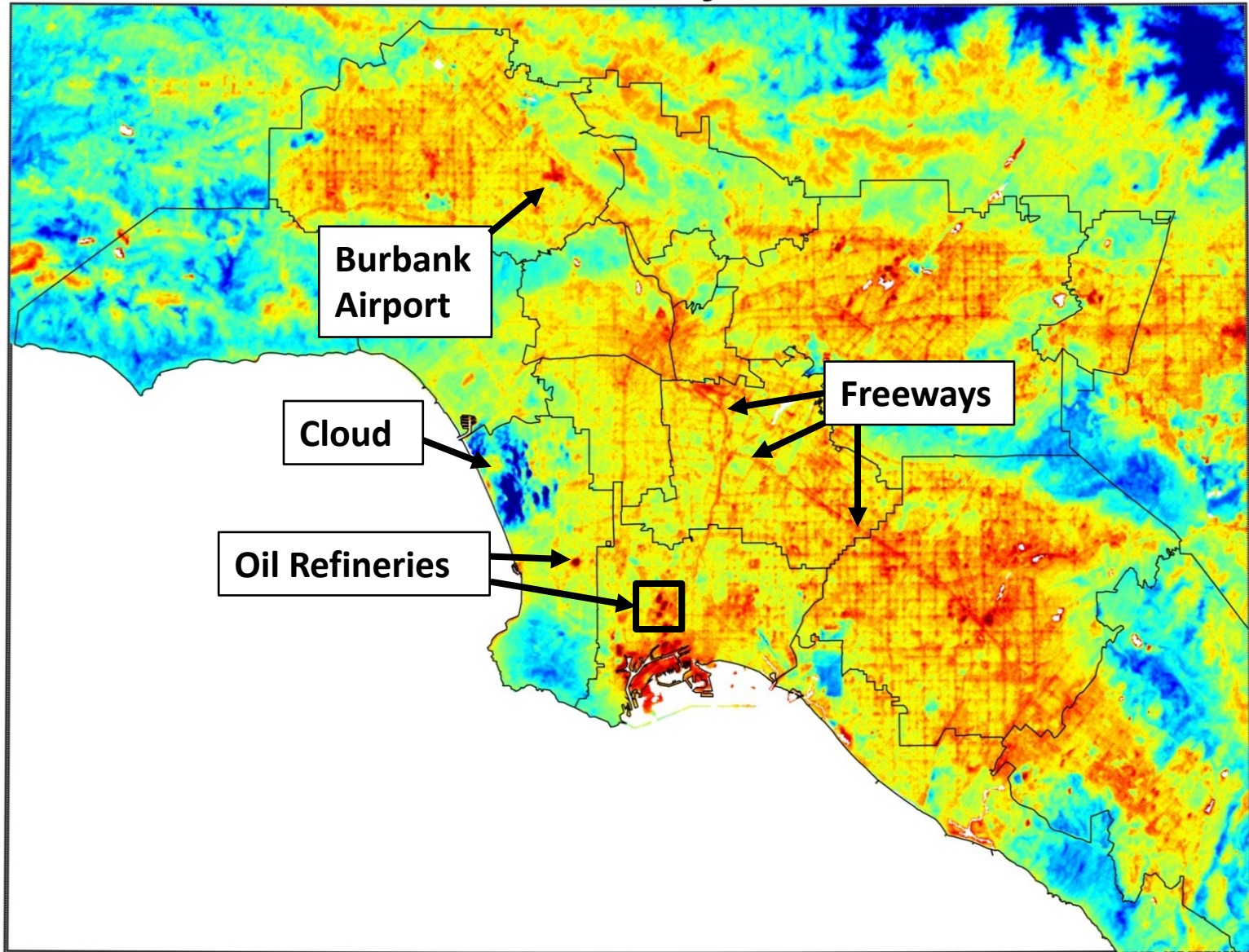
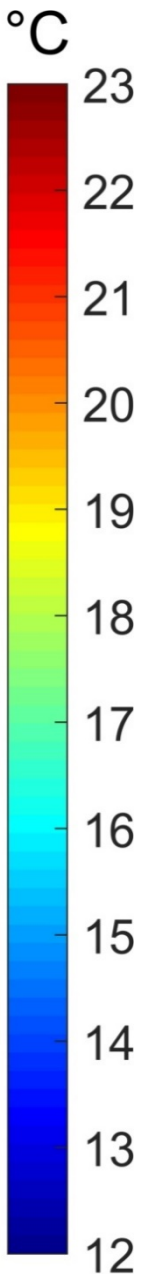
34.4° N

33.5° N

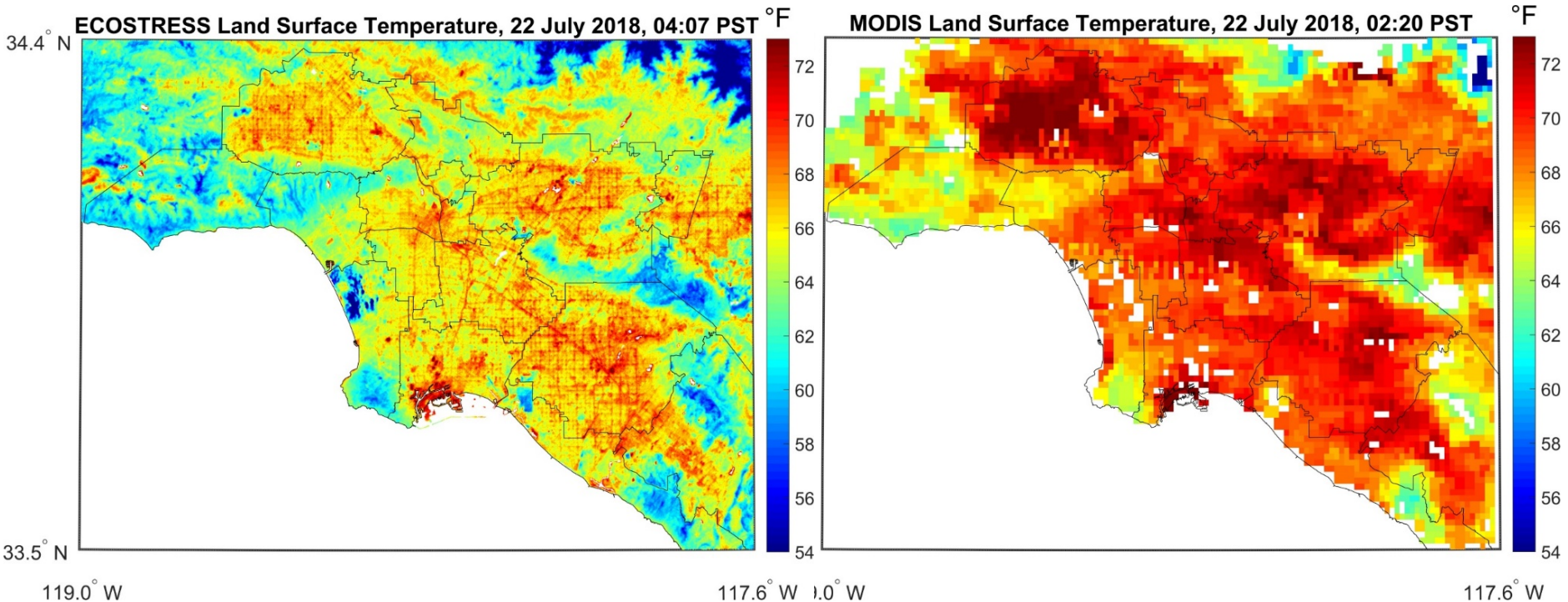
119.0° W

(c) 2018 California Institute of Technology. Government sponsorship acknowledged.

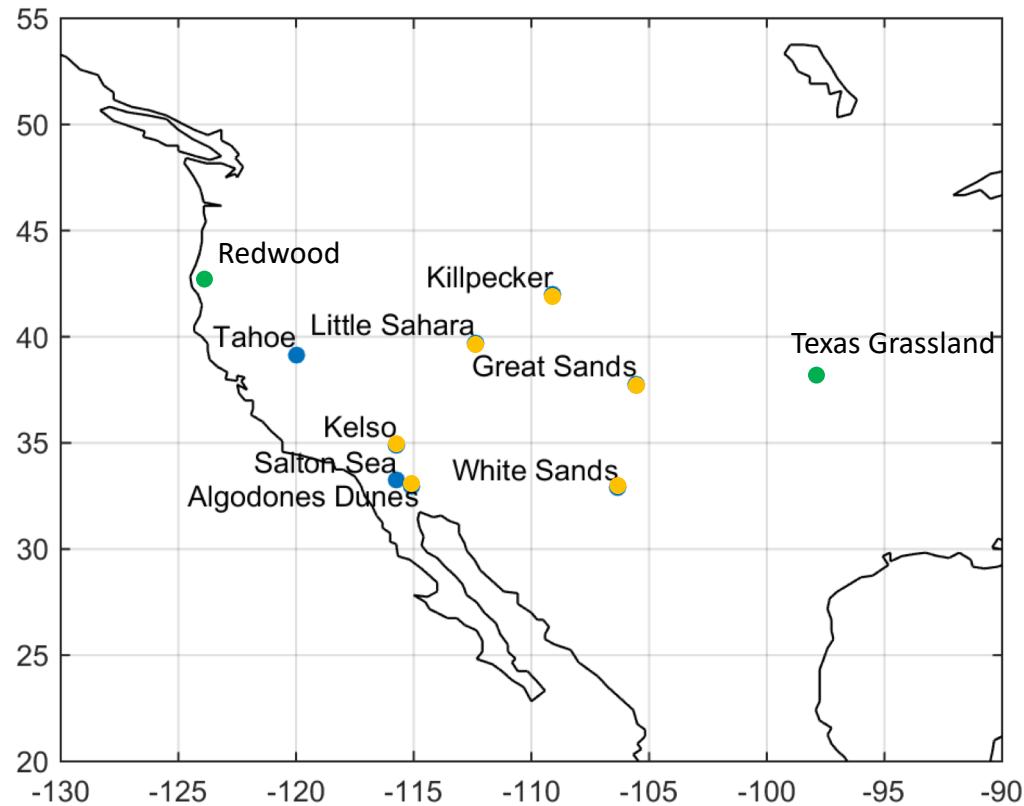
117.6° W



ECOSTRESS (left) compared to MODIS (right) highlight ~200x improved spatial detail of temperature



LST&E Validation Sites (Stage 1)

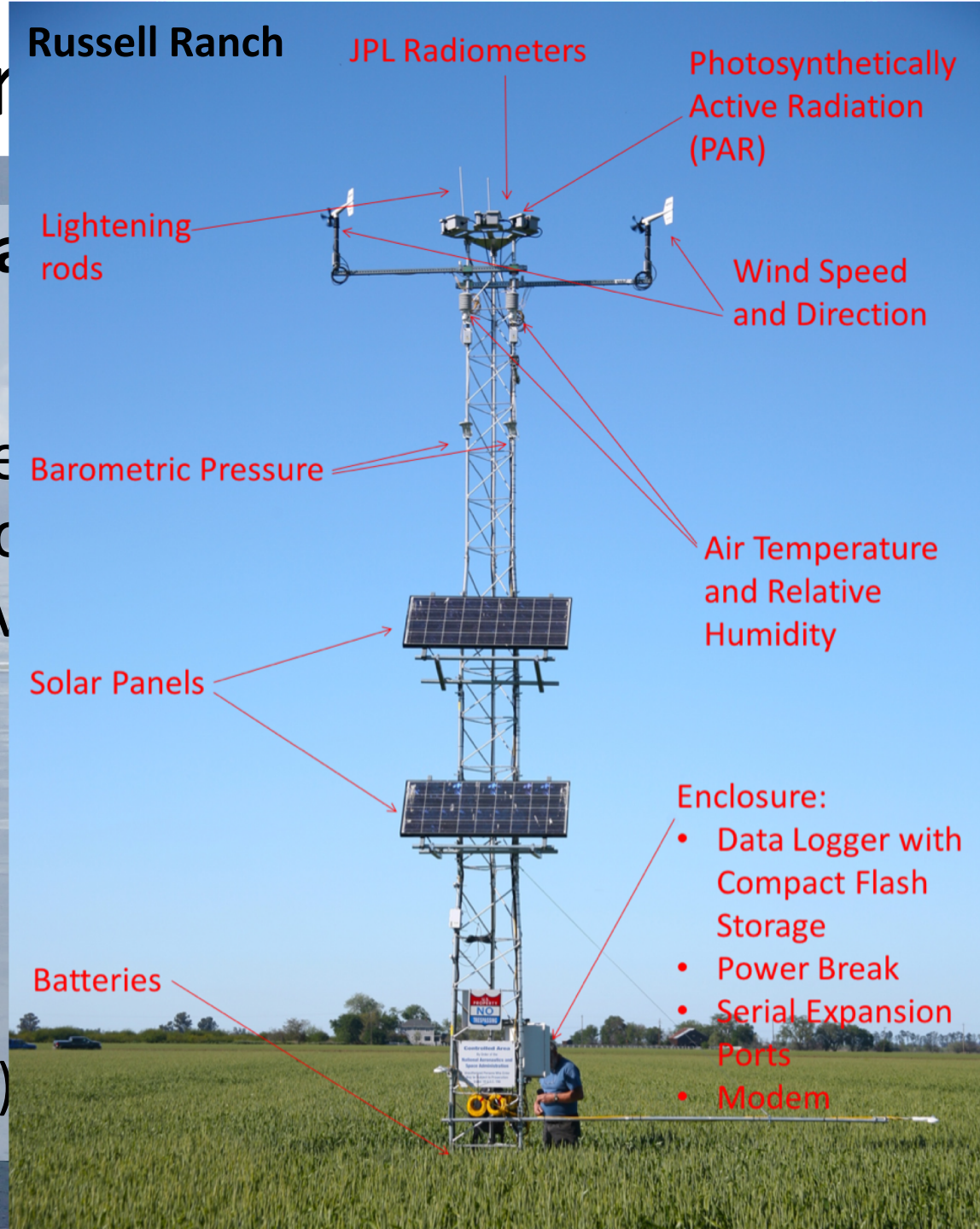


Temperature-based and Radiance-based validation methods

(Wan et al. 2008, Guillevic et al. 2012, Schneider et al. 2013, Hulley et al. 2012, Hook et al. 2007)

Temper

- **Method 1: Temperature**
- Measure the temperature of the overpass and compare to the temperature retrieval
- Core Tval sites:
 - Lake Tahoe
 - Salton Sea
 - Russell Ranch
 - Gobabeb (Namibia)

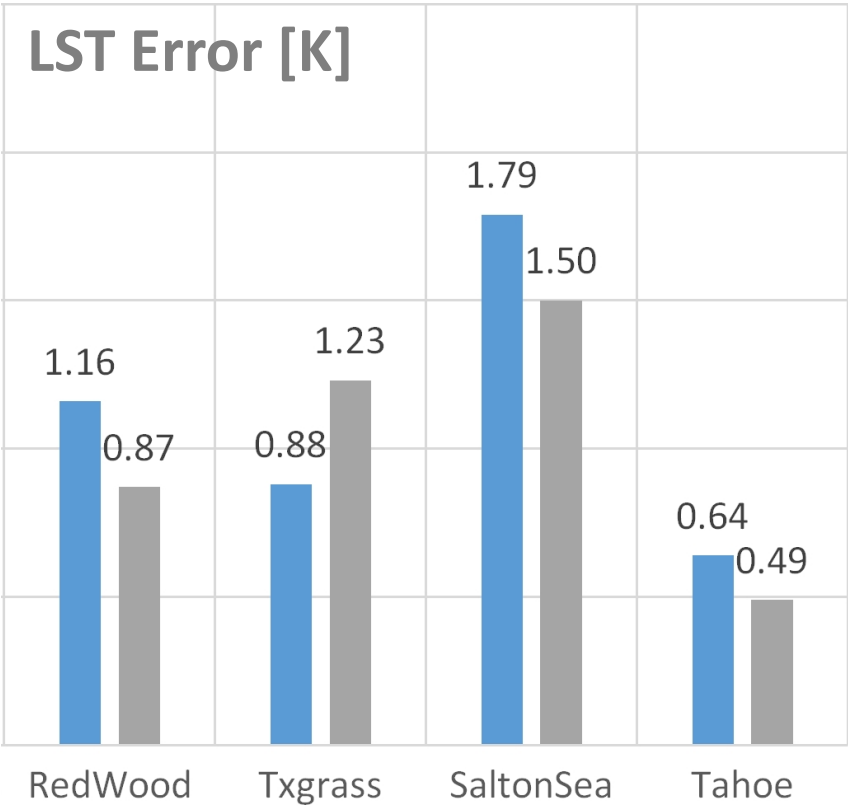
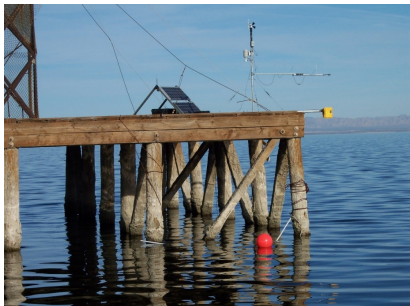


Validation: Graybody Sites

2003-2005
Day and Night

MOD21
(TES)

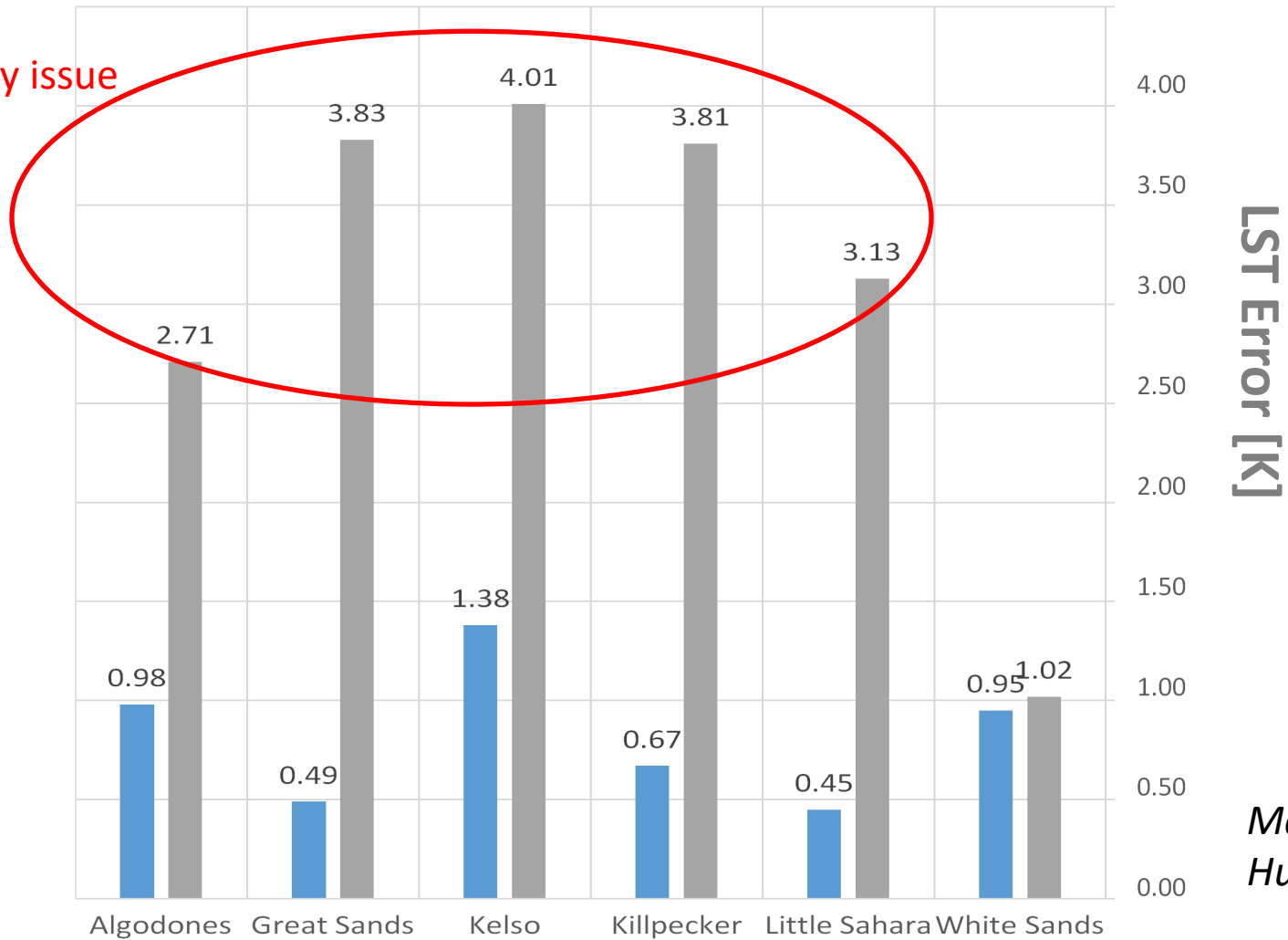
MOD11 (C5)
(Split-Window)



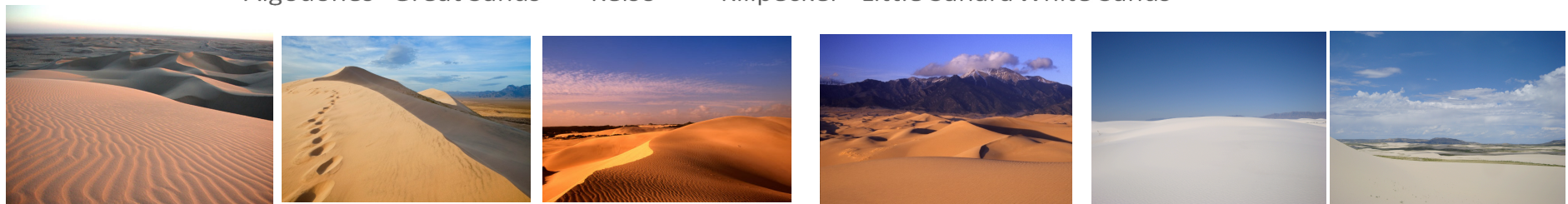
■ MOD21 (TES) ■ MOD11 (C5) (Split-Window)

2003-2005
Day and Night

MOD11
Emissivity issue



*Malakar and
Hulley, 2014*



Samples collected in field

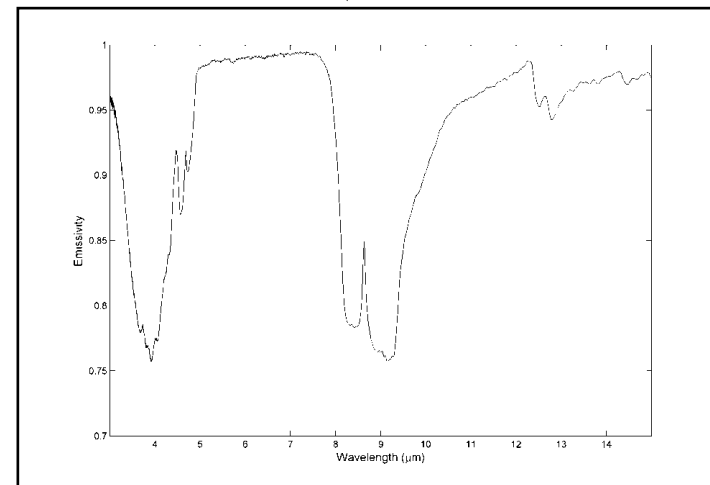
Reflectance measured using Nicolet 520 FTIR spectrometer

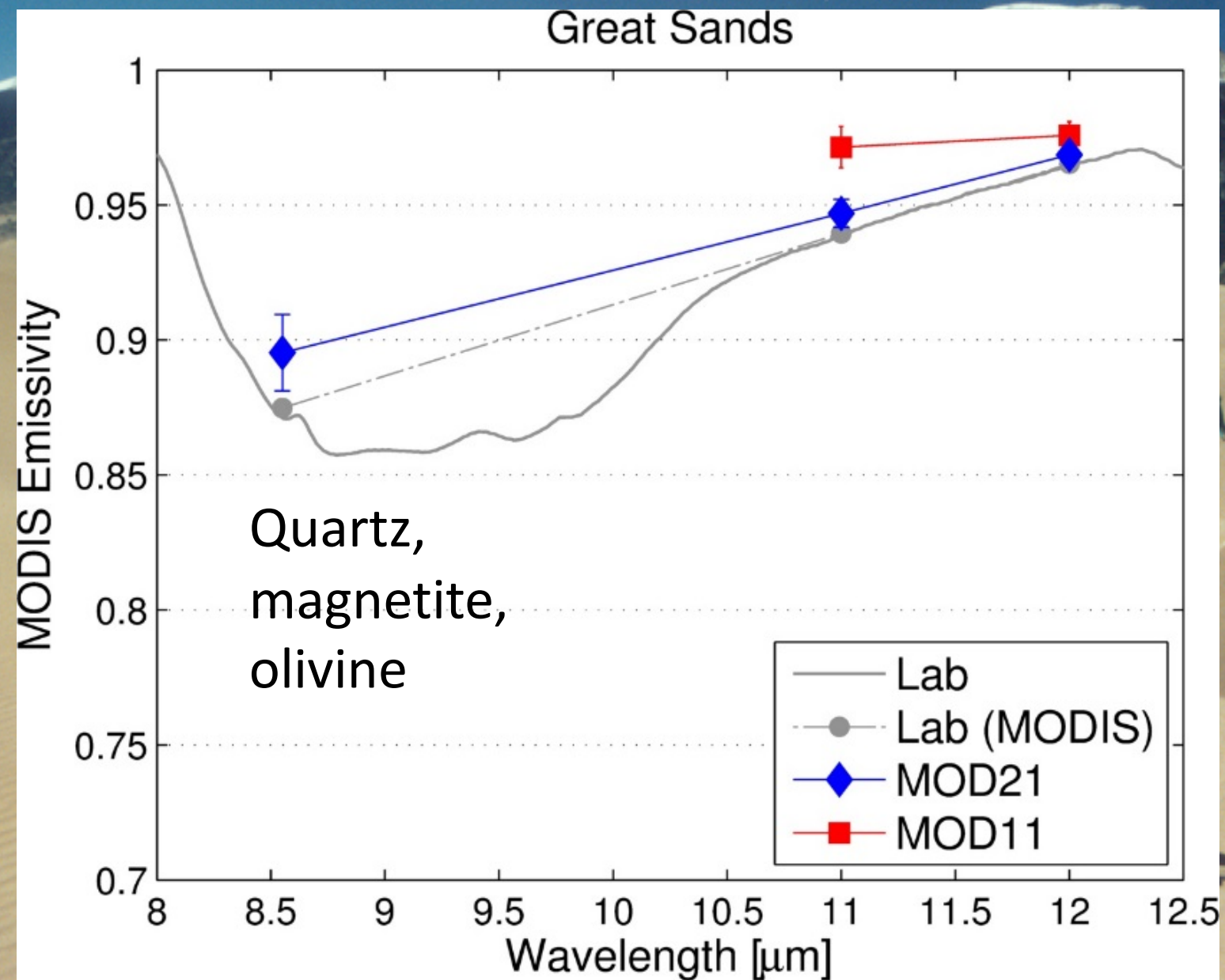


$$\varepsilon = 1 - r$$

Emissivity Validation

spectral range: 2.5 – 15 μm
spectral resolution: 4 cm^{-1}
1000 scans in 10 minutes





New TIR Field Instruments (arriving soon)

1. Nulling field portable thermal infrared imager (FLIR camera)
2. Field FT-IR spectrometer (shown at right).

Technical Equipment, Facilities and Infrastructure Management (**TEFIM**)

SPECIFICATIONS				
Item	Parameter	Value	Units	Comments
1	Spectral Range	2 - 16	micrometers	Standard IR
2	Spectral Resolution (FWHH)	4	wavenumbers	Standard, 1 sec. scan
3	Size (WxDxH)	36x20x23	centimeters	(14"x8"x9")
4	Weight	< 7	kilograms	(< 15 pounds)

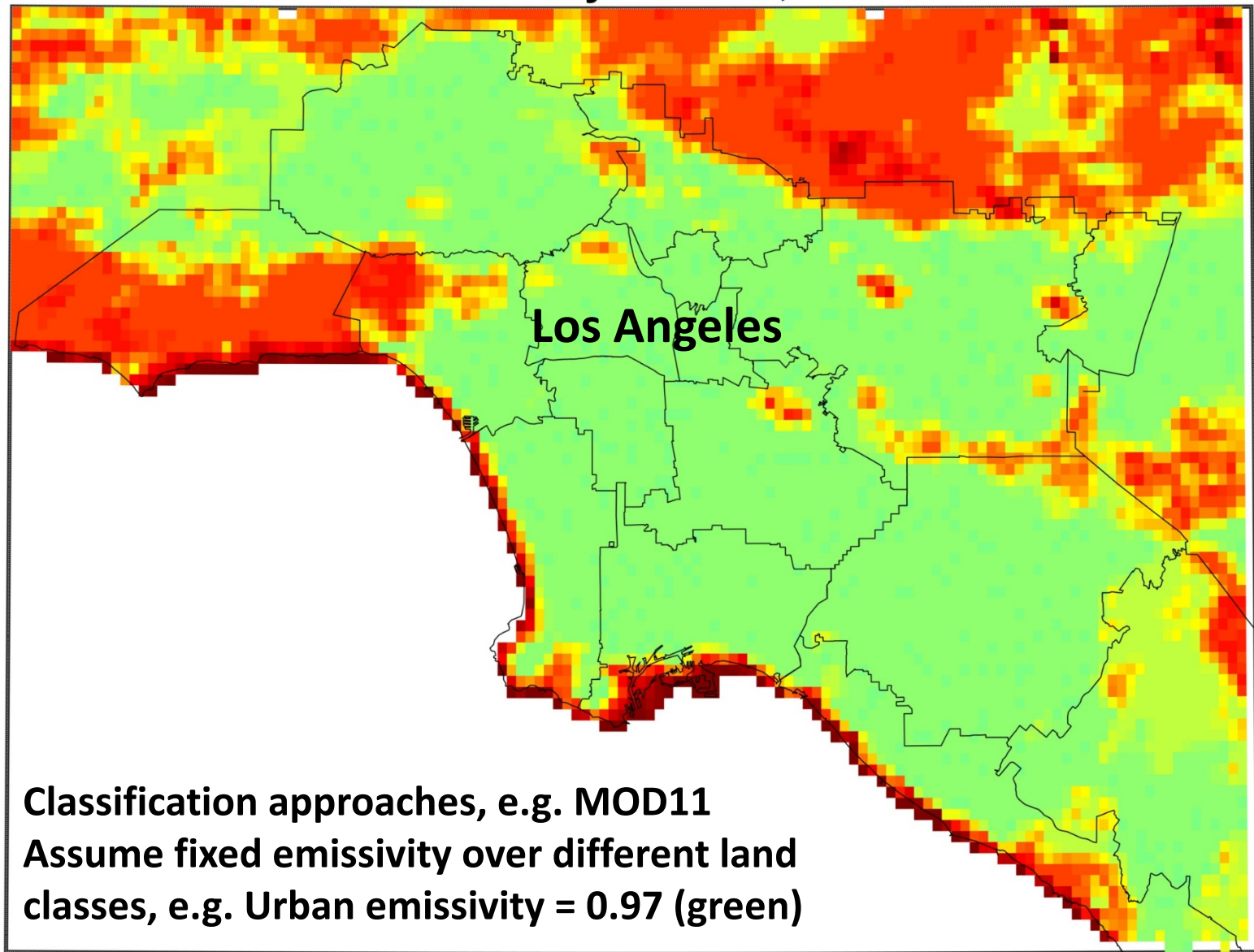


Summary

- HyspIRI L1 instrument and algorithm maturity from ECOSTRESS/PHyTIR 'pathfinder'
- HyspIRI L2 Products:
 - Land Surface Temperature (LST)
 - Spectral Emissivity (6 bands)
 - Cloud Mask
 - Error estimates and Quality Control
- Well defined and strong L2 TES algorithm heritage (ASTER, MODIS, VIIRS, HyTES, MASTER)
- L1/L2 ATBD and other documentation complete
- CalVal sites established

MYD11 Emissivity Band 31, 08/16/2016

34.4° N



0.99
0.985
0.98
0.975
0.97
0.965
0.96
0.955
0.95

Classification approaches, e.g. MOD11
Assume fixed emissivity over different land
classes, e.g. Urban emissivity = 0.97 (green)

33.5° N

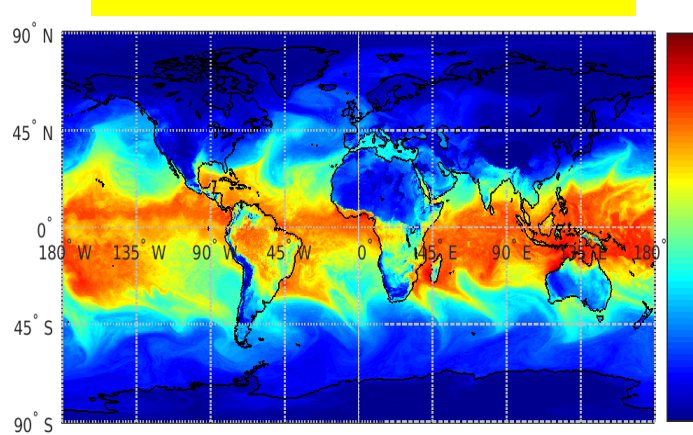
119.0° W

117.6° W

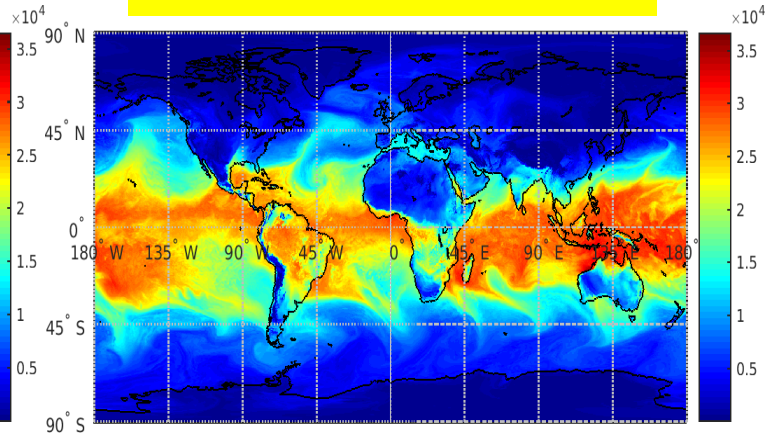
MERRA Global Water Vapor Field (level 1)

(Water vapor at 2m, ppmv)

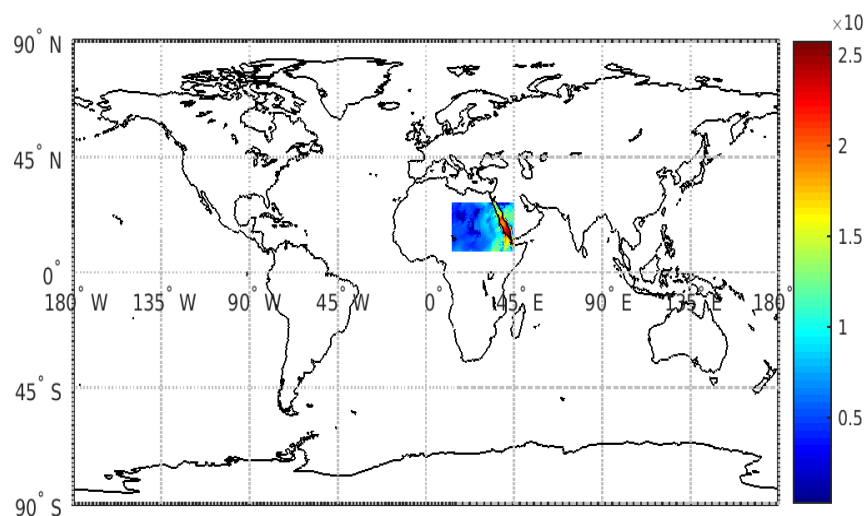
Native Grid 1 – 06hr UTC



Native Grid 2 – 12hr UTC



Native Grid 3 – Clipped and time-interpolated to 'ECOSTRESS' swath at 11hr UTC



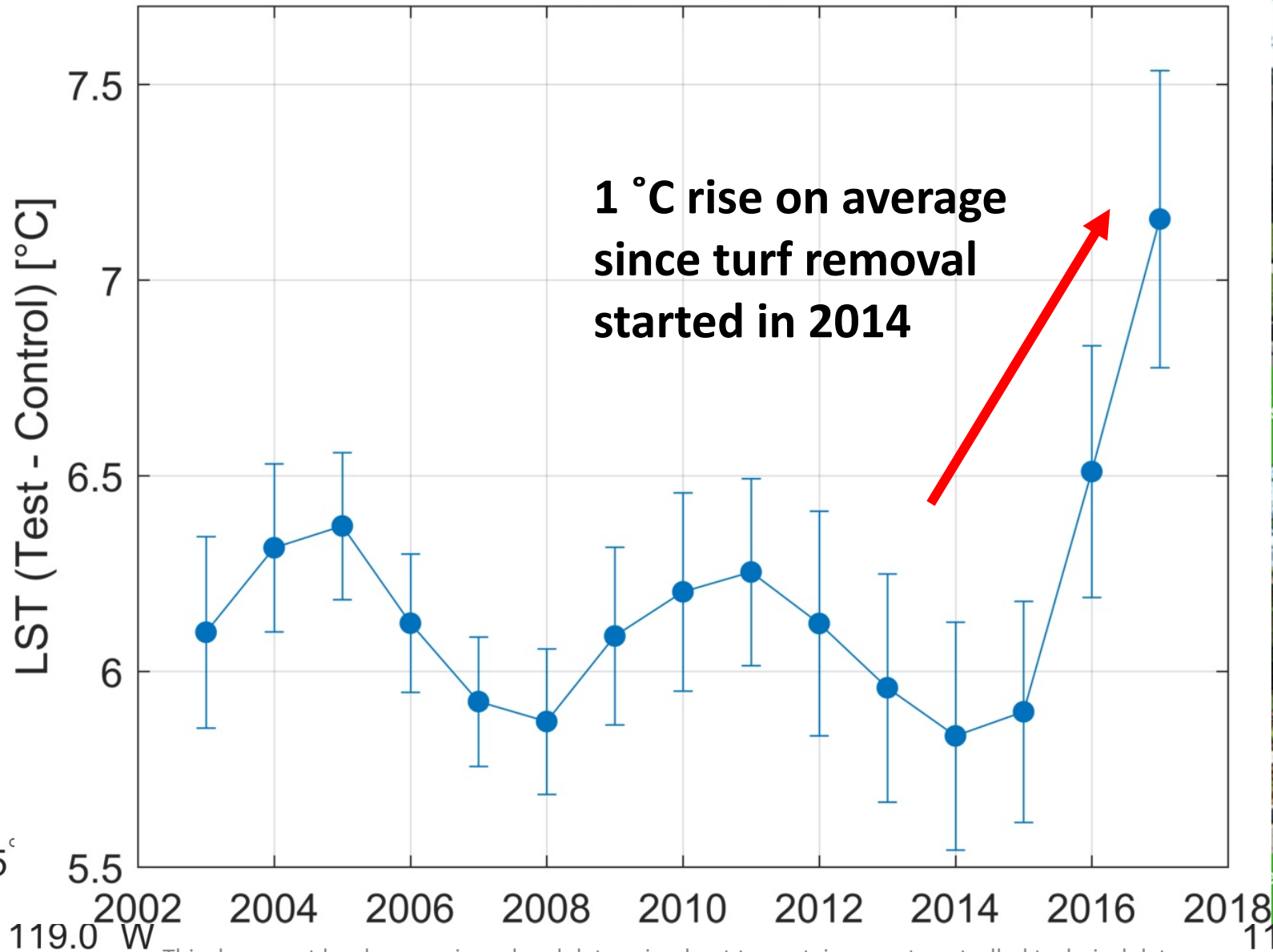
***Profiles and other ancillary data can be saved to hdf, netcdf or other format at this point for use in L3/L4*

MYD21 Emissivity Band 31, 08/16/2016

34.4° N

0.99

Test - Control

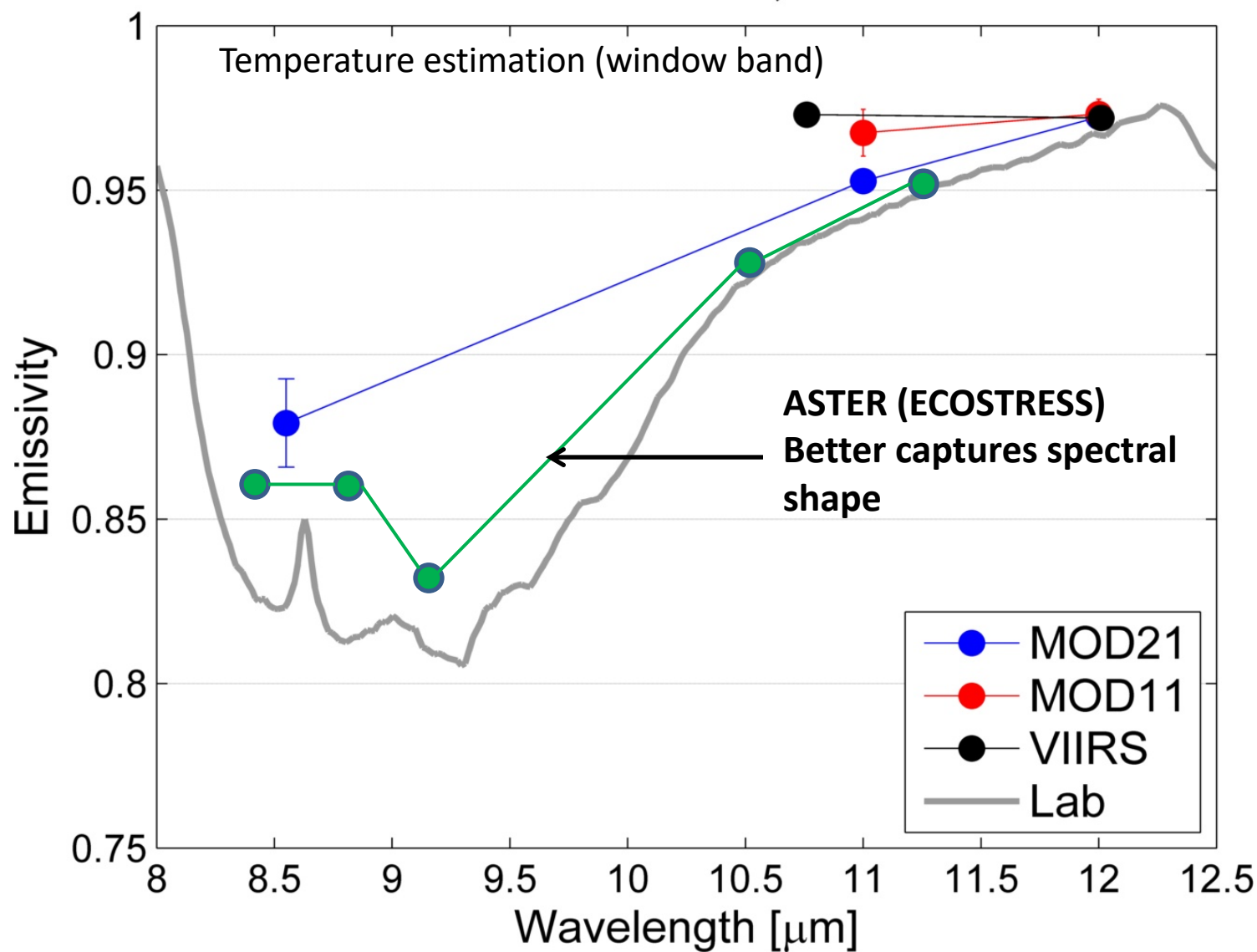


yard where the
sed as reference



ect is completed.

Kelso Dunes, CA





Theoretical Basis: Surface Temperature

Radiometric ('Brightness') Temperature

$$T_{\lambda}(\theta) = B_{\lambda}^{-1}(L_{\lambda}(\theta))$$

Land Surface ('Skin') Temperature (LST)

$$T_s = B_{\lambda}^{-1}\left(\frac{L_{\lambda}(\theta) - \rho_{\lambda}L^{\downarrow}}{\varepsilon_{\lambda}}\right)$$

where :

B_{λ} = blackbody spectral radiance

λ = wavelength

T_s = Surface Temperature

$T_{\lambda}(\theta)$ = Surface Brightness Temperature

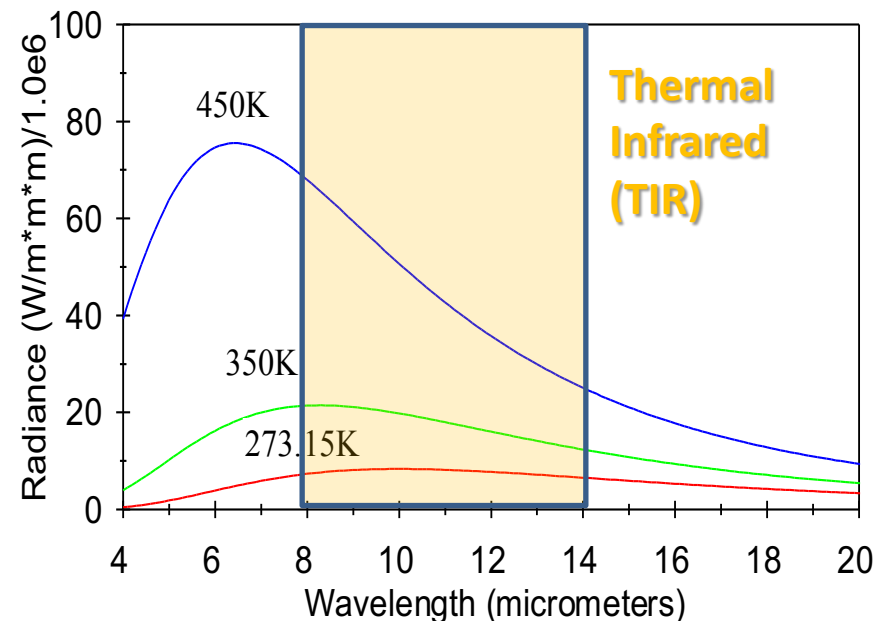
ρ_{λ} = Surface Reflection

L^{\downarrow} = Downwelling Sky Irradiance

ε_{λ} = Emissivity

Planck Function

$$B_{\lambda} = \frac{C_1}{\lambda^5 \left[\exp\left(\frac{C_2}{\lambda T_s}\right) - 1 \right]}$$



in the Planck function shifts to shorter and shorter wavelengths

Theoretical Basis: Land Surface Temperature (LST)

1. Radiometric ('Brightness') Temperature

$$T_{\lambda}(\theta) = B_{\lambda}^{-1}(L_{\lambda}(\theta))$$

2. Land Surface ('Skin') Temperature (LST)

$$T_s = B_{\lambda}^{-1}\left(\frac{L_{\lambda}(\theta) - \rho_{\lambda}L^{\downarrow}}{\varepsilon_{\lambda}}\right)$$

where :

B_{λ} = blackbody spectral radiance

λ = wavelength

T_s = Surface Temperature

$T_{\lambda}(\theta)$ = Surface Brightness Temperature

ρ_{λ} = Surface Reflection

L^{\downarrow} = Downwelling Sky Irradiance

ε_{λ} = Emissivity

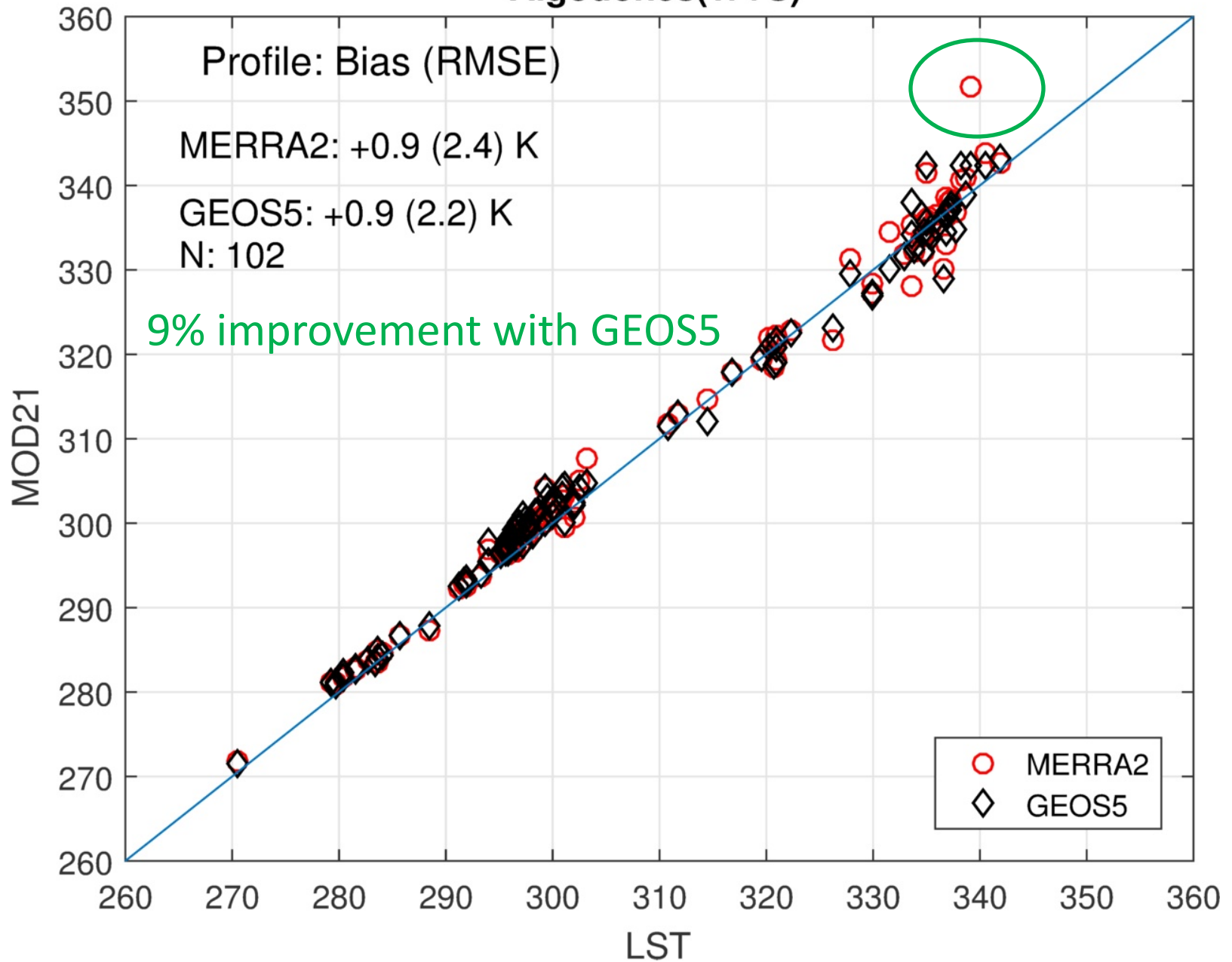
Temperature Validation

- **Method 2: Radiance-based (Rval)**

Measure the emissivity and use an independent atmospheric profile (e.g. radiosonde). Forward calculate the ground temperature needed to match the at-sensor radiance and compare to retrieved temperature (Radiative Closure Simulation).

Works day/night and globally when accurate in situ emissivity information is available

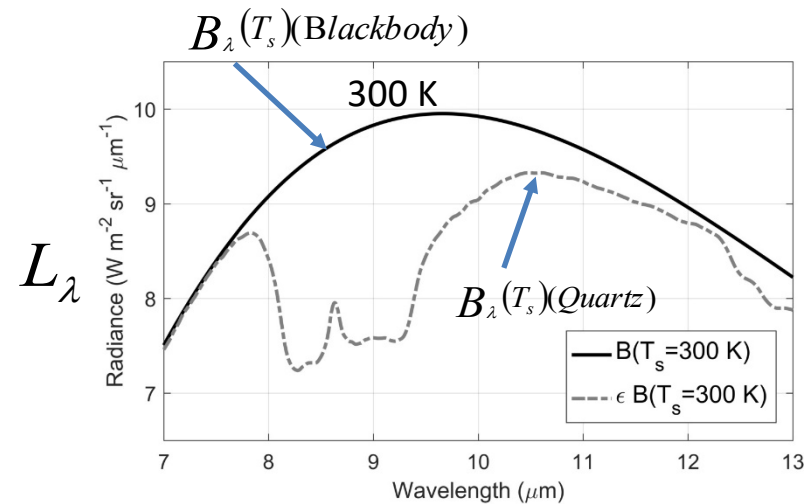
Algodones(WVS)



Theoretical Basis: Surface Emissivity

Blackbody Radiance

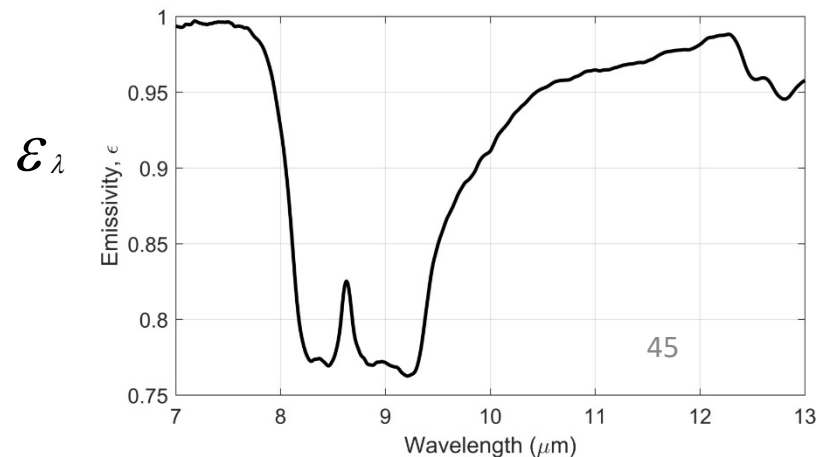
$$L_{\lambda} = \underset{\substack{\nearrow \\ \text{Planck}}}{B_{\lambda}}(T_s)$$



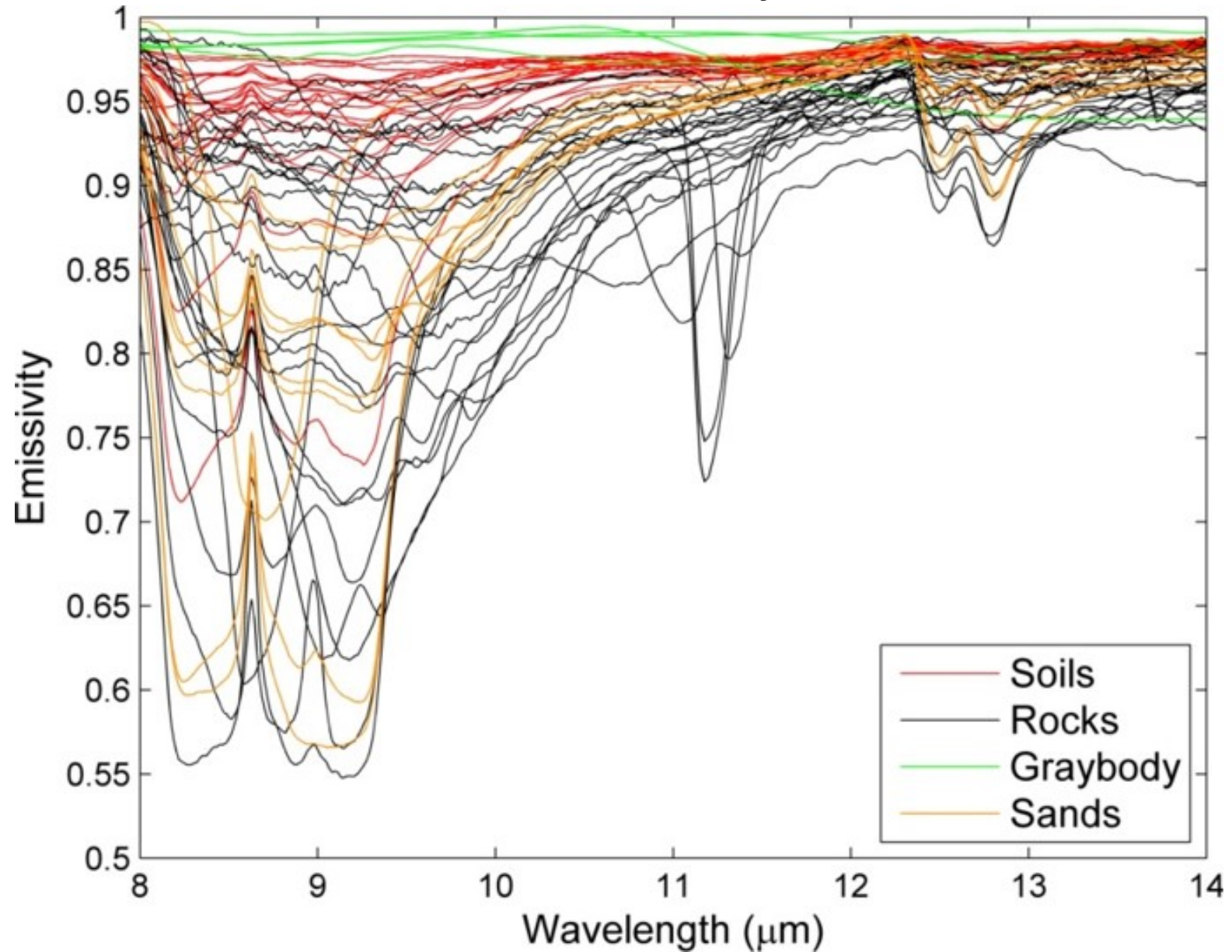
Spectral Emissivity:

ratio of the spectral radiance of a material to that of a blackbody at the same temperature:

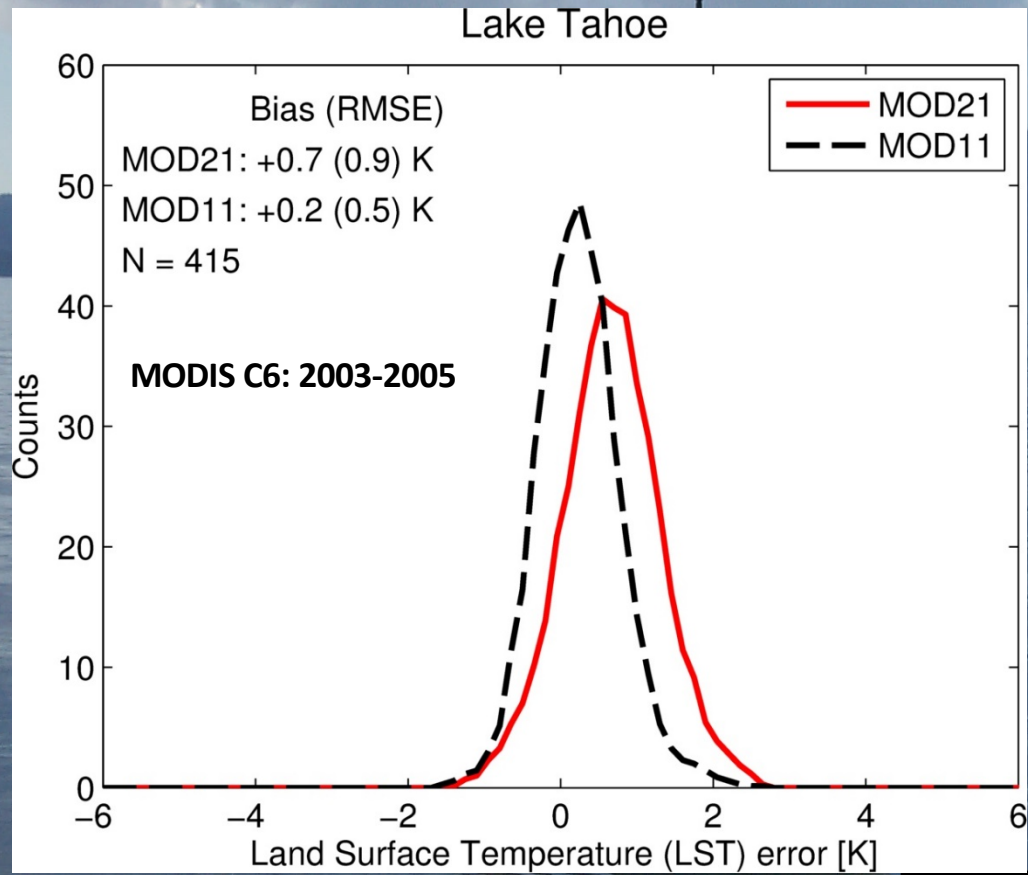
$$\epsilon_{\lambda} = \frac{L_{\lambda}(\text{Material})}{L_{\lambda}(\text{Blackbody})}$$



Emissivity

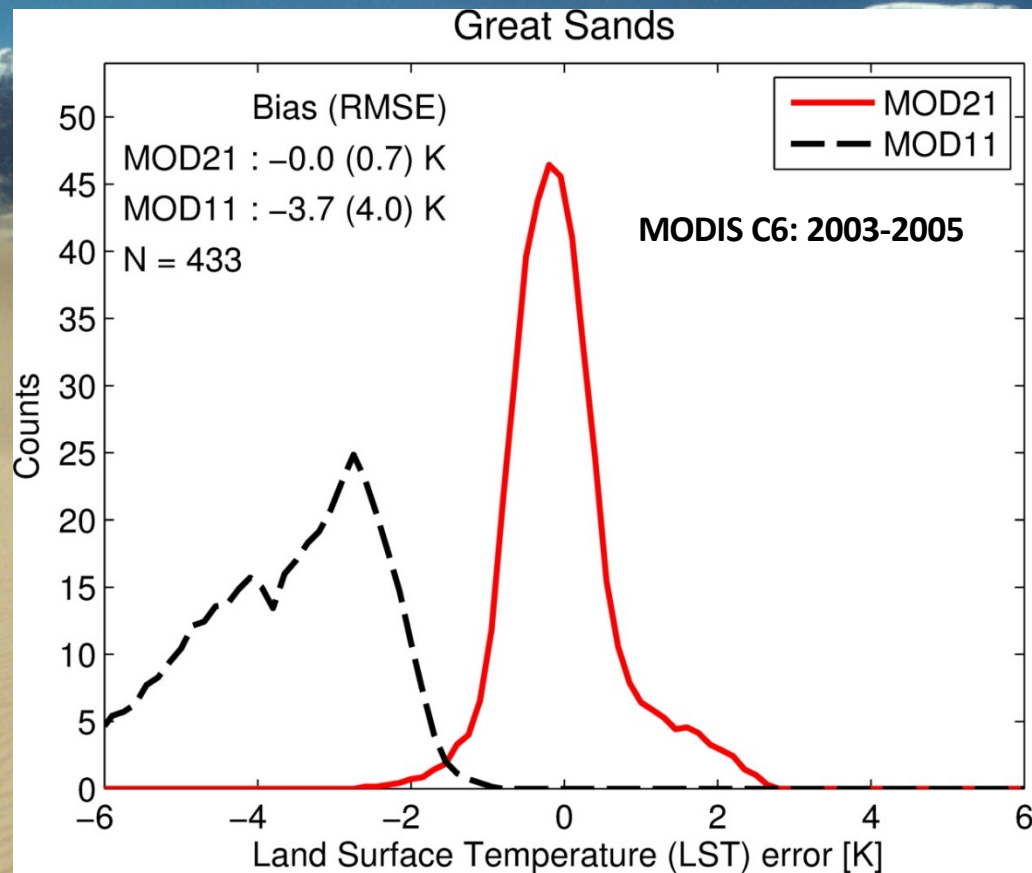


Lake Tahoe, CA (2003-2005)



6/17/2003 1:20pm

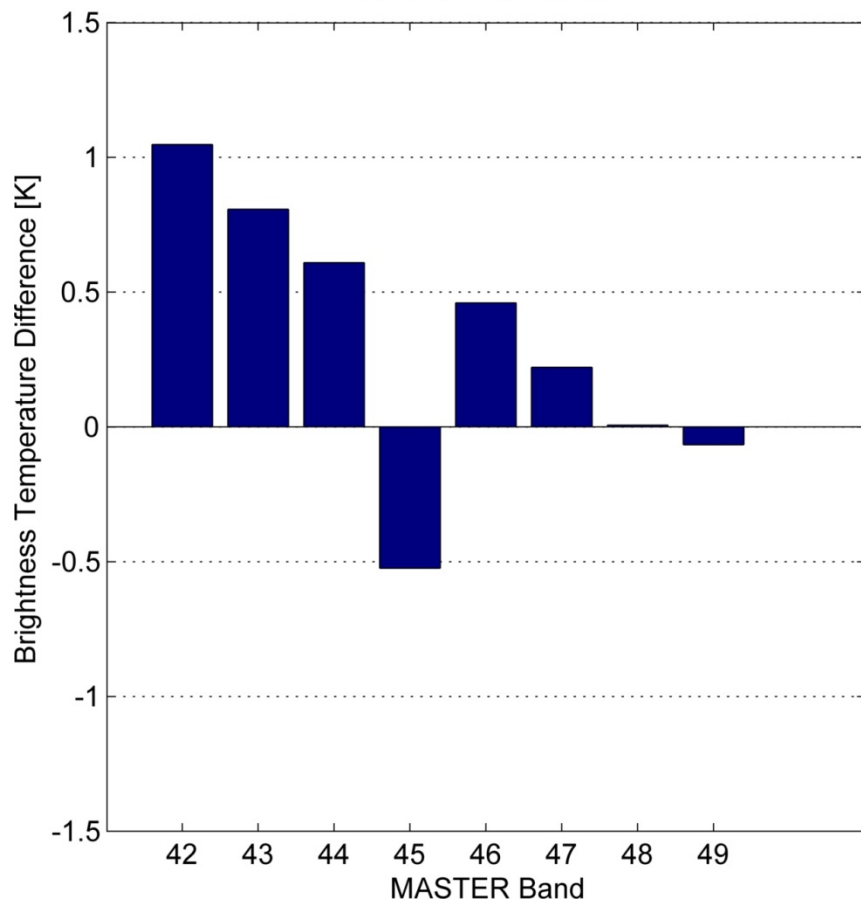
Great Sands, CO (2003-2005)



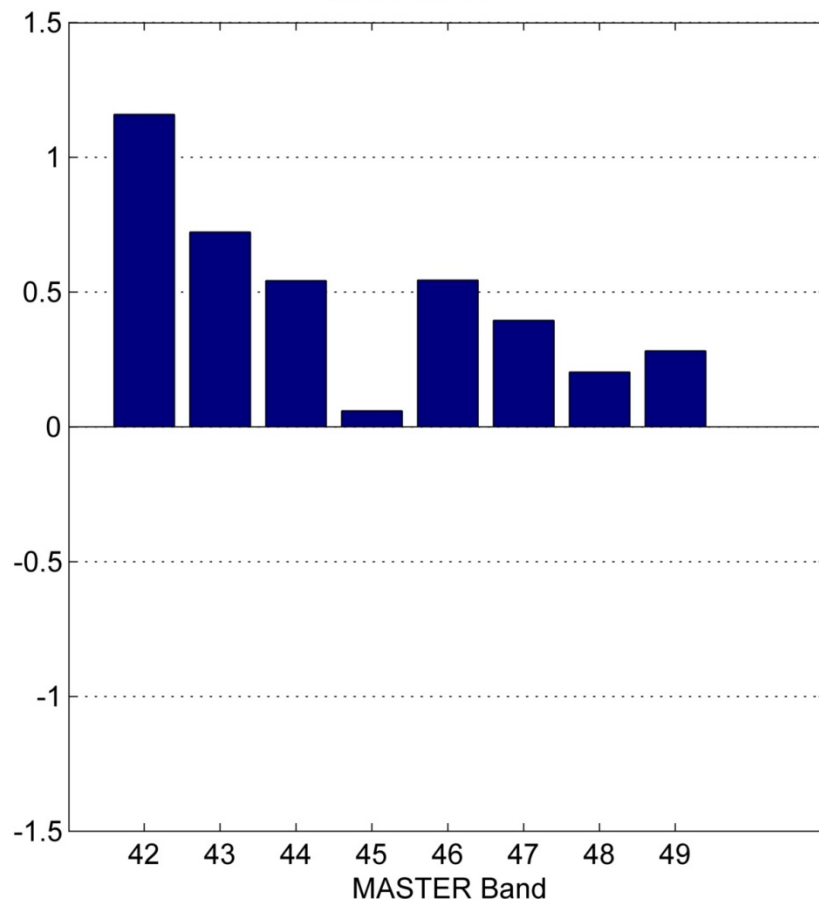


Salton Sea L1 Calibration

Salton Sea 2013



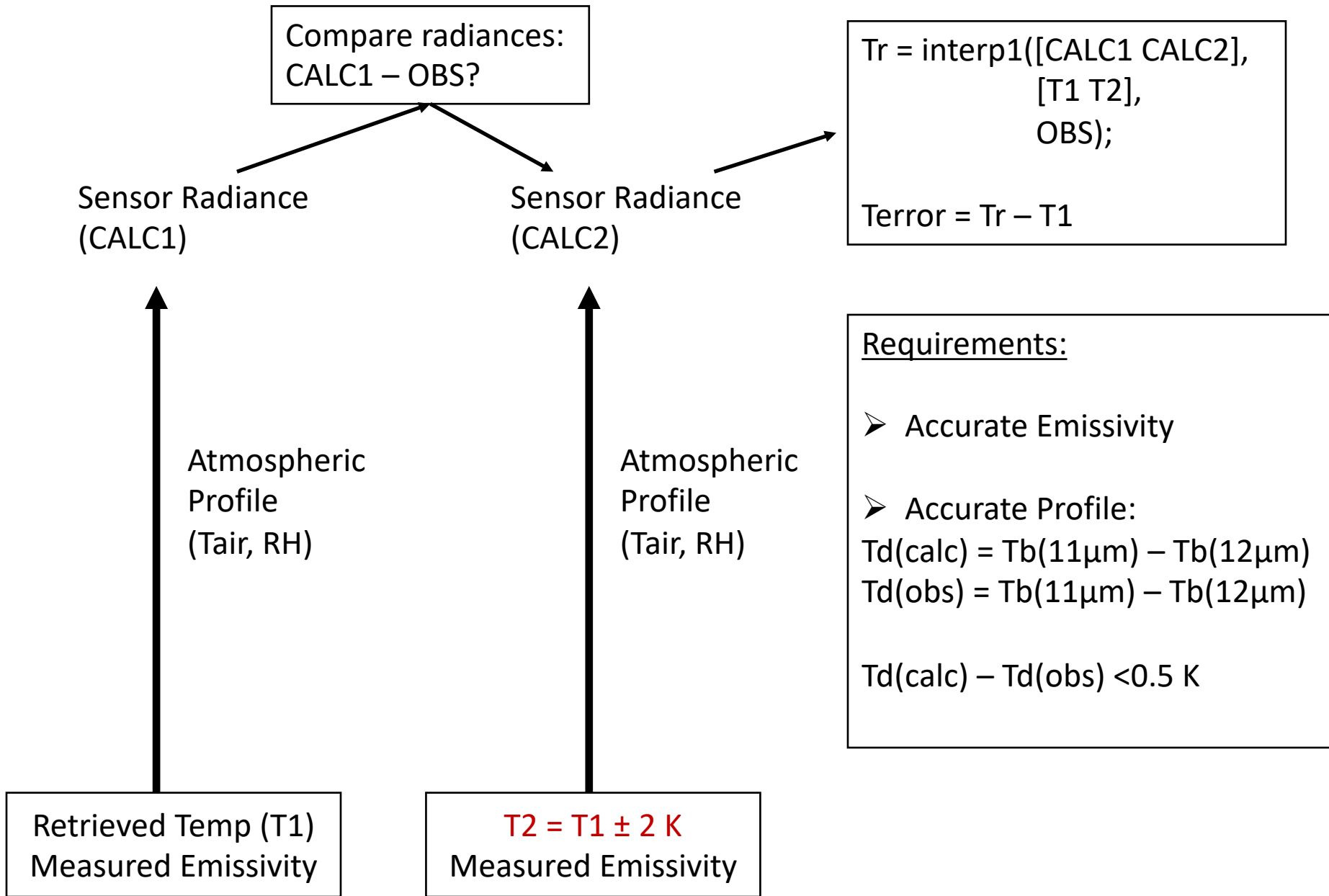
Salton Sea 2014



Baseline	Threshold (Mission Success)
VSWIR:380 to 2500 nm at ≤ 10 nm sampling at the specified signal-to-noise ratio and accuracy with <u>>95%</u> spectral/spatial uniformity at ≤ 30 m nadir spatial sampling with <16 day revisit to provide	VSWIR:380 to 2500 nm at ≤ 10 nm sampling at <u>>80%</u> of the specified signal-to-noise ratio and accuracy with <u>>90%</u> spectral/spatial uniformity at ≤ 30 m nadir spatial sampling with <16 day revisit to provide
VSWIR: <u>>60%</u> seasonal and <u>>80%</u> annual coverage of the terrestrial and shallow water regions of the Earth	VSWIR: <u>> 50%</u> seasonal and <u>>70%</u> annual coverage of the terrestrial and shallow water regions of the Earth
<u>three years</u> with a subset of measurements available <u>near-real-time</u> for designated science and applications.	<u>two years.</u>
TIR: 8 spectral bands from the 3-5 micron and 8-12 micron regions of the spectrum at the specified noise-equivalent-delta-temperature and accuracy at ≤ 60 m nadir spatial sampling with ≤ 5 day revisit	TIR: 8 spectral bands from the 3-5 micron and 8-12 micron regions of the spectrum at <u>>80%</u> the specified noise-equivalent-delta-temperature and accuracy at ≤ 60 m nadir spatial sampling with ≤ 5 day revisit
<u>TIR: >60% Monthly, >70% seasonal and >85% annual coverage of the terrestrial and shallow water regions of the Earth</u>	<u>TIR: > 40% Monthly, > 60% seasonal and >70% annual coverage of the terrestrial and shallow water regions of the Earth</u>

*30 m, 16 Day VSWIR

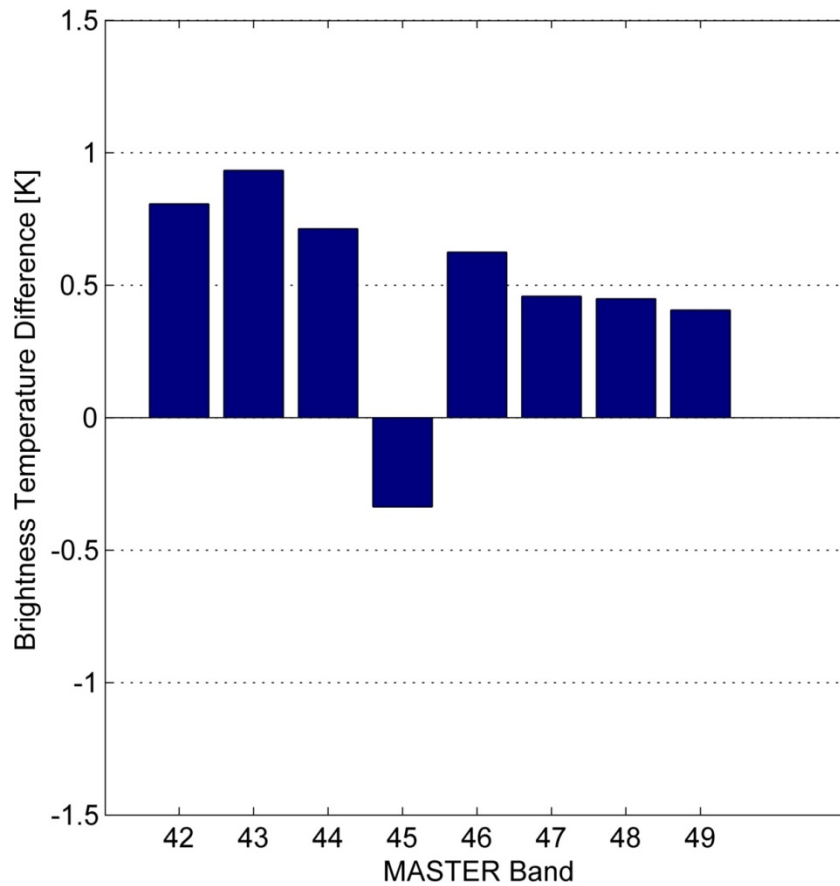
Radiance-Based Temperature Validation



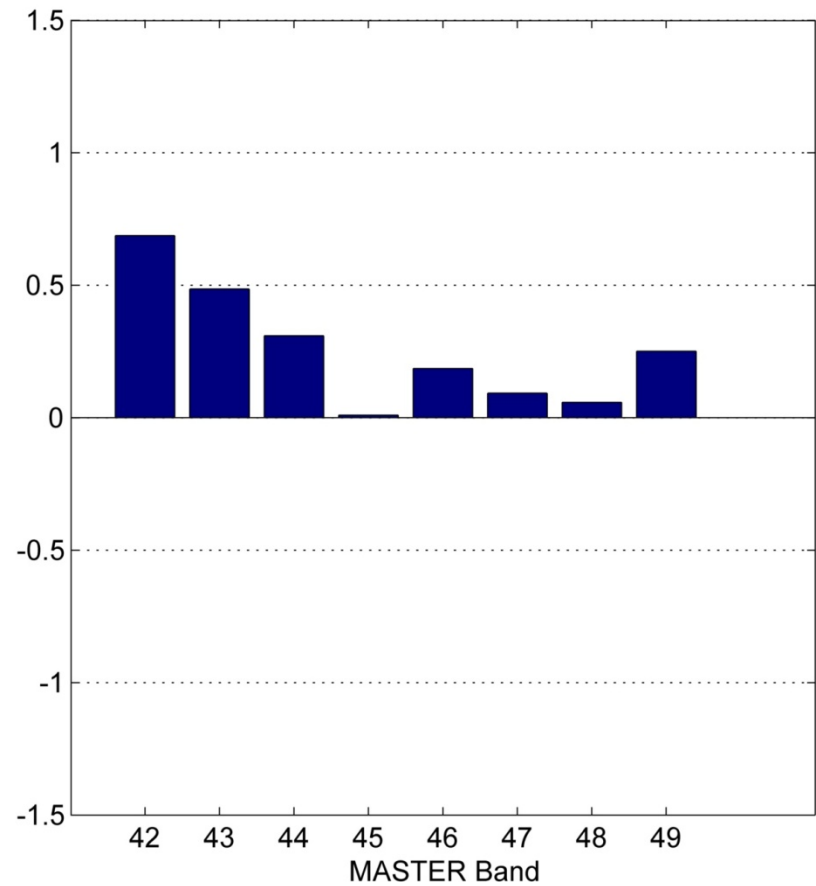
MASTER TIR Calibration



Lake Tahoe 2013



Lake Tahoe 2014



L1A Radiometric Calibration PGE

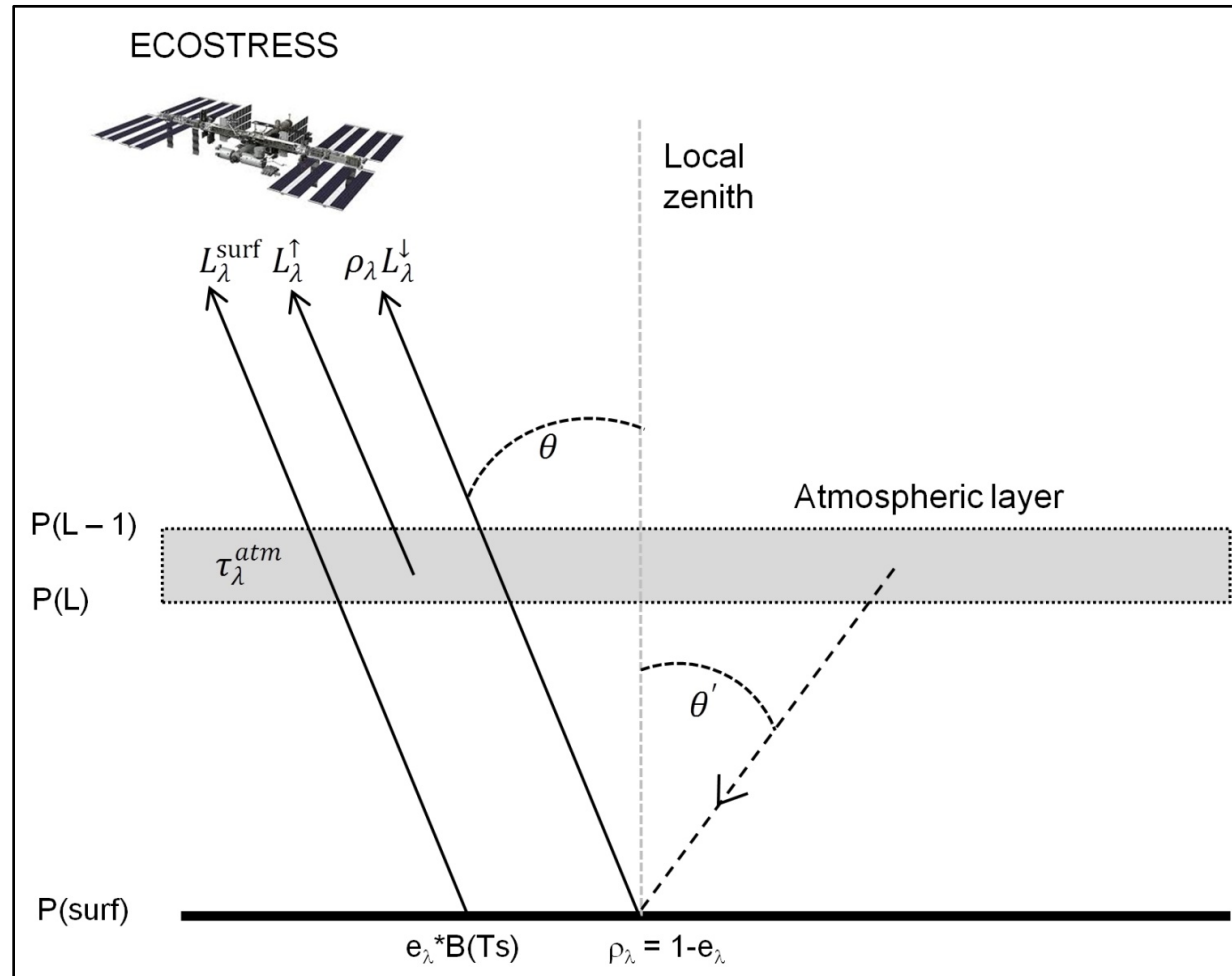
L1A Radiometric Calibration Steps*

- **Purpose: Convert Image TIR DNs to Radiance:**
 - Procedure for each image:
 - Read temperatures from Sensor's Cold (~295K) and Hot (~325K) Blackbodies.
 - Create synthetic FPA temperature images of Cold and Hot Blackbodies and convert them to Radiance (Watt/m²/sr/um) using the center wavelength of each TIR band and the Planck function.
 - Collect push-whisk FPA Digital Number (DN) scans of the Cold and Hot Blackbodies And Ground for all wavelengths.
 - Using the FPA Radiance values and corresponding FPA DNs, use a two-point affine transformation (creating gain/offset coefficients) to convert each Ground pixel's DN to Radiance.
- Accuracy is expected to be better than the 1.0 Kelvin requirement.
- Ground Radiance and Temperature images can be generated for Validation and Verification purposes as necessary (optional parameters).

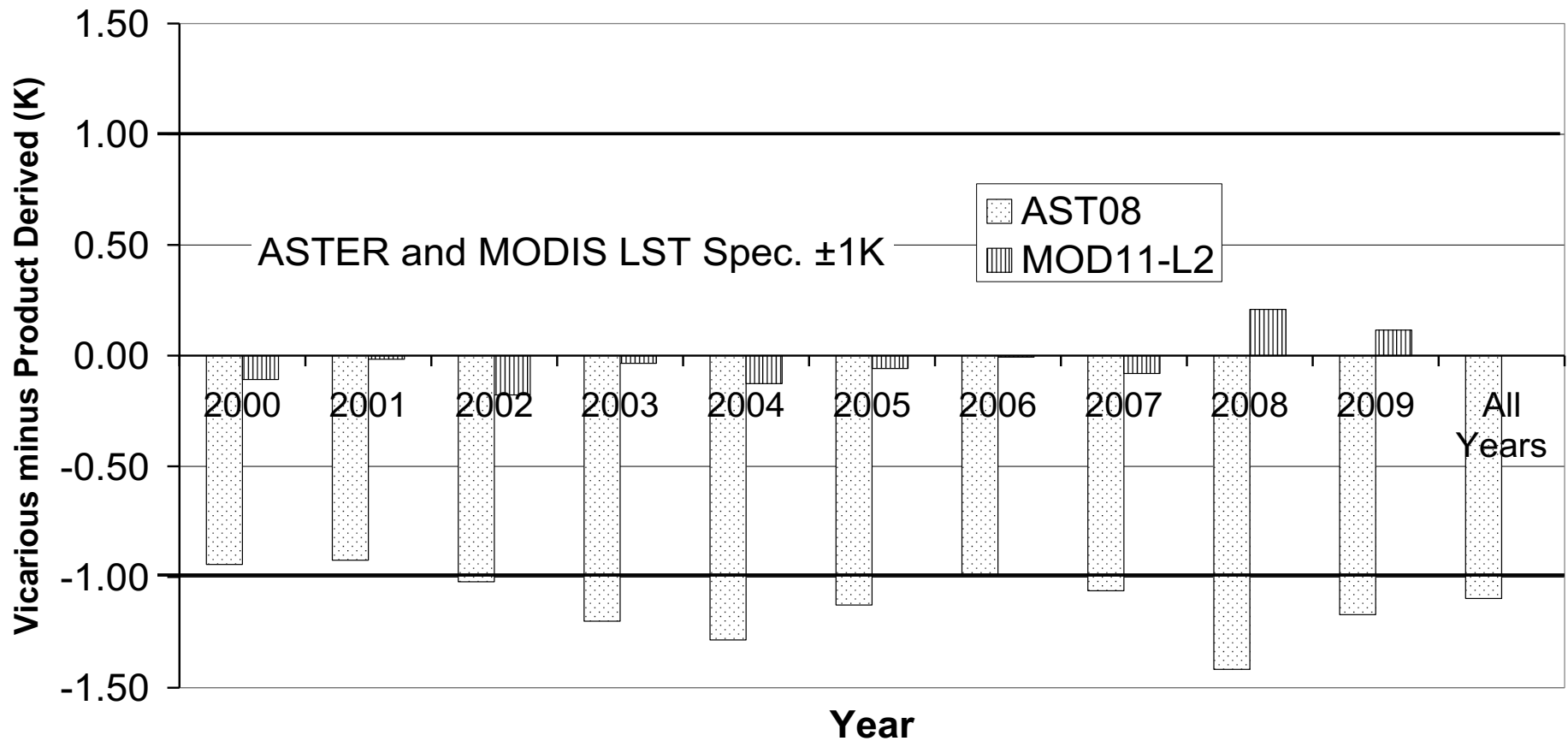
*Documented in: "Level-1 Focal Plane Array and Radiometric Calibration Algorithm Theoretical Basis Document (ATBD)," JPL D-94803.

Atmospheric correction + Temperature Emissivity Separation (TES)

$$L_{\lambda}(\theta) = \tau_{\lambda}(\theta) \left[L_{\lambda}^{surf} + (1 - \varepsilon_{\lambda}) \bar{L}_{\lambda}^{\downarrow} \right] + L_{\lambda}^{\uparrow} \quad L_{\lambda}^{surf} = \varepsilon_{\lambda} B(T_s)$$



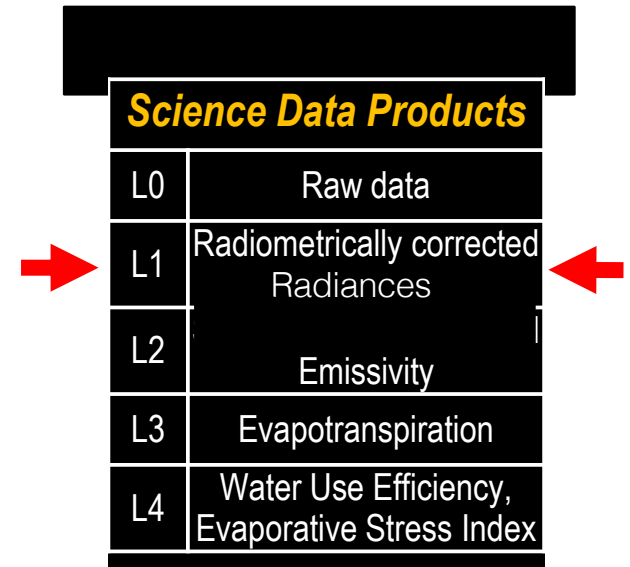
Tahoe LST validation



The MODIS product is accurate to ($\pm 0.2\text{K}$), while the ASTER product has a bias of 1-2 K due to residual atmospheric correction effects

L1 TIR Overview

- Level-1 (L1) is part of the Science Data System (SDS), where the SDS:
 - Creates L0, L1, L2, L3, and L4 products, and
 - Delivers products to the Land Process DAAC (Sioux Falls, SD)
- **Level-1 Inputs include:**
 - L0 Data
 - Raw Image Data Packets
 - Spacecraft Orbital Metadata
 - Ancillary Data
 - Landsat Ortho-Rectified Image Base (geolocation)
 - Digital Terrain Models (pass-through)
 - Elevation
 - Land/Water Mask
- **Level-1 Outputs include:**
 - Calibrated Radiance images with
 - Geolocation (position) and
 - Associated metadata



Outline

1. L1 Algorithm Maturity
2. L1 Cal/Val
3. L2 Algorithm Maturity
4. L2 Cal/Val



TIR Instrument Characteristics

Spectral	
Bands (8) μm	3.98 μm , 7.35 μm , 8.28 μm , 8.63 μm , 9.07 μm , 10.53 μm , 11.33 μm , 12.05 μm
Bandwidth	0.084 μm , 0.32 μm , 0.34 μm , 0.35 μm , 0.36 μm , 0.54 μm , 0.54 μm , 0.52 μm
Accuracy	<0.01 μm
Radiometric	
Range	Bands 2–8 = 200 K – 500 K; Band 1 = 1200 K
Resolution	< 0.05 K, linear quantization to 14 bits
Accuracy	< 0.5 K 3-sigma at 250 K
Precision (NE Δ T)	< 0.2 K
Linearity	>99% characterized to 0.1 %
Spatial	
IFOV	60 m at nadir
MTF	>0.65 at FNy
Scan Type	Push-Whisk
Swath Width	600 km ($\pm 25.5^\circ$ at 623-km altitude)
Cross Track Samples	9,300
Swath Length	15.4 km (± 0.7 degrees at 623 km altitude)
Down Track Samples	256
Band to Band Co-Registration	0.2 pixels (12 m)
Pointing Knowledge	10 arcsec (0.5 pixels) (approximate value, currently under evaluation)
Temporal	
Orbit Crossing	11 a.m. Sun synchronous descending
Global Land Repeat	5 days at Equator
On Orbit Calibration	
Lunar views	1 per month {radiometric}
Blackbody views	1 per scan {radiometric}
Deep Space views	1 per scan {radiometric}
Surface Cal Experiments	2 (day/night) every 5 days {radiometric}
Spectral Surface Cal Experiments	1 per year
Data Collection	
Time Coverage	Day and Night
Land Coverage	Land surface above sea level
Water Coverage	Coastal zone minus 50 m and shallower
Open Ocean	Averaged to 1-km spatial sampling
Compression	2:1 lossless