



HyspIRI VSWIR Science Measurement and Instrument Concept Baseline

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HyspIRI Measures the Optical Spectrum







HyspIRI VSWIR Decadal Survey Science and Decadal Survey Climate Science



- Key HyspIRI climate objectives from the Decadal Survey and IPCC
 - Ecosystem Measurement for Climate Feedback
 - Black Carbon/Dust Effects on Snow and Ice
 - Carbon Release from Biomass Burning
 - Evapotranspiration and Water Use and Availability
 - Critical Volcanic Eruption Parameters
- Imaging Spectrometer (VSWIR)
 - Pattern and Spatial Distribution of Ecosystems and their Components
 - Ecosystem Function, Physiology and Seasonal Activity
 - Biogeochemical Cycles
 - Changes in Disturbance Activity
 - Ecosystem and Human Health
 - Earth Surface and Shallow Water Substrate Composition
 - Combined Imaging Spectrometer and Multi-Spectral Thermal Science
 - Coastal habitats, and inland aquatic environments
 - Wildfires
 - Volcanoes
 - Ecosystem Function and Diversity
 - Land surface composition and change
 - Human Health and Urbanization







Example: Imaging Spectroscopy Mapping Wetland Dominants 2010 LA









AVIRIS Vegetation Spectra



Vegetation mapped cleanly across scene boundaries

- Phragmites (phau)
- Spartina alterniflora (spal)
- Spartina patens (sppa)
- Vigna luteola (vilu)

D. Roberts, UCSB



Spectroscopy for Species/Functional-type, Biogeochemistry and Physiological Condition





Measuring spectral shift from Lignin : Cellulose ratio







Surface Compositional Derived with Imaging **Spectrometer Measurements**











Tetracorder 3.3 product Sulfates K-Alunite 150c K-Alunite 250c K-Alunite 450c Na82-Alunite 100c Na40-Alunite 400c Jarosite Alunite+Kaolinite and/or Muscovite Kaolinite group clays Kaolinite, wxl Kaolinite, pxl Kaolinite+smectite or muscovite Halloysite Dickite Calcite Calcite +Kaolinite Calcite + montmorillonite Na-Montmorillonite Nontronite (Fe clay) other minerals low-Al muscovite med-Al muscovite high-Al muscovite Chlorite+Musc,Mont Chlorite Buddingtonite Chalcedony: OH Qtz Pyrophyllite +Alunite **♦** N 2 km

Cuprite, Nevada AVIRIS 1995 Data

USGS

Clark & Swayze





Albedo and Black Carbon/Dust Effects on Snow/Ice



What is causing the downwasting and retreat of Himalayan glaciers and

For snow and ice in the Himalaya, increasing temperatures and increasing dust and soot combine in unknown proportions to accelerate melt through their changes in albedo. Imaging Spectrosocpy is the only approach that allows us to attribute changes in albedo into effects from temperature and dust/black carbon and at a fine enough spatial resolution that heterogeneous terrain can be resolved. Multiband sensors such as NPOESS VIIRS have neither capacity.



Required Measurement: Global glacial covered area, full solar spectrum, < 100 m spatial, <20 days revisit









HyspIRI: Coral, Benthic Composition, and Aquatic Vegetation

















Variation in shallow water HyspIRI-type spectral signatures in coral environments.

Variation in HyspIRI-type spectral signatures of floating aquatic vegetation (e.g. Kelp)

Variation in shallow water HyspIRI-type spectral signatures in seagrass beds and benthic habitat materials





1 Km

Emergent vegetation signatures are well suited to the HyspIRI measurement. For example mapping in the Gulf of Mexico coastal region with AVIRIS measurements.



HyspIRI VSWIR Science Measurements







HyspIRI is a global mission, measuring land and shallow aquatic habitats at 60 meters and deep oceans and ice sheets at 1km every 19 days (VSWIR)

HyspIRI's VSWIR imaging spectrometer directly measures the full solar reflected spectrum of the Earth from 380 – 2500nm at 10 nm.





HyspIRI VSWIR SNR and Uniformity Characteristics





Cross Track Sample



Depiction

-Grids are the detectors -Dots are the IFOV centers -Colors are the wavelengths

Requirement

Spectral Cross-Track

Spectral-IFOV-Variation



>95% cross-track uniformity {<0.5 nm min-max over swath}

>95% spectral IFOV uniformity {<5% variation over spectral range}



-150

Longitude

Seasonal and Annual Cloud Probability Maps Validate the HyspIRI Coverage Requirements



150



-150

Longitude

150



HyspIRI VSWIR Coverage



• EO-1 Hyperion acquisitions in 10 years. Technology demonstration sampling mission with deep ESTO contribution.



• HyspIRI VSWIR provides complete terrestrial coverage every 19 days.

• It would take Hyperion 100 years to acquire what HyspIRI measures in 1 year.





VSWIR - Instrument Concept



HyspIRI – VSWIR

OUTLINE

- 1. Introduction
- 2. Key Requirements & Performance
- 3. VSWIR Concept
- 4. Technology Readiness & Heritage



Mass (CBE) 55Kg Power (Ave.) 41Watts



Key VSWIR Requirements



<u>Spectral</u>	Requirement	Status
Range	380 to 2500 nm (solar reflected spectrum)	Demonstrated – AVRIS, MaRS, M3
Sampling	<= 10 nm {uniform over range}	Demonstrated – MaRS, M3
Response	<= 13 nm (FWHM) {uniform over range}	Demonstrated – MaRS, M3
Accuracy	<0.5 nm	Demonstrated – MaRS, M3, CAO-VSWIR

<u>Radiometric</u>	Requirement	Status
Range & Sampling	0 to 1.5 x benchmark radiance, 14 bits	Demonstrated via analysis and 14 bit ADC bread board electronics
Accuracy	>95% absolute radiometric, 98% on-orbit reflectance, 99.5% stability	Demonstrated – AVIRIS, MaRS
Precision (SNR)	See spectral plots at benchmark radiances	Demonstrated via analysis
Linearity	>99% characterized to 0.1 %	Demonstrated via test, MaRS and CAO-VSWIR
Polarization	<2% sensitivity, characterized to 0.5 %	Demonstrated via analysis of design and test data on the grating
Scattered Light	<1:200 characterized to 0.1%	Demonstrated in MaRS and CAO-VSWIR



Key VSWIR Requirements



<u>Spatial</u>	Requirement	Status
Range	>145 km	Demonstrated by design and analysis (150 km)
X-track Sampling	>2400	Demonstrated by design and analysis (2500)
Sampling	<= 60m (Nadir)	Demonstrated by design and analysis
Response	<= 1.2X sampling (FWHM)	Demonstrated by MaRS and M3
<u>Uniformity</u>		
Spectral Cross-Track	>95% cross-track uniformity {<0.5 nm min-max over swath}	Demonstrated by MaRS and M3
Spectral-IFOV- Variation	>95% spectral IFOV uniformity {<5% variation over spectral range}	Demonstrated by MaRS and M3

<u>Other Key</u>	Requirement	Status
Data rate	~ 300 Mbits per second	Met by preliminary architecture and parts selection – all required parts are at or above TRL 6
Compression	3:1 lossless	Met by algorithm test on MaRS data. Algorithms implemented in breadboard electronics and flight FPGA
Pointing Knowledge	60m radius (3σ)	Met by analysis and the use of ground tie points
Mass	<55 kg	Met by current design with 30% margin – working to increase margin



Instrument Approach



• Selected: Offner spectrometer (Hyperion, CRISM, M3, ARTEMIS, COMPASS^{air,} NG^{air})

- Full range from 380 to 2500 nm demonstrated. Efficiency for high SNR optimized with multiple blaze grating demonstrated. Uniformity from design through alignment demonstrated. Snapshot acquisition detector. Dispersion efficiency tunable to optimize use of detector.
- Prism dispersion spectrometer
 - Dispersion is non-uniform. Cross-track and spectral-IFOV uniformity not inherent in optical design.
 Dispersion efficiency not tunable in detail to optimize use of detector full well.
- Wedge/Linear-variable filter spectrometer
 - Full spectral range coverage from 380 to 2500 nm has not been demonstrated maintaining 10 nm spectral sampling and response function. Filter uniformity is a concern over wide spectral and spatial domain. Fast, high throughput, beams interplay with filter spectral bandpass undermines uniformity.
- Fourier Transform Spectrometer
 - Dispersion is non constant with wavelength. Not typically built to operate below 1 micron. Detector dynamic range and photon shot noise concern. Architecture for > 2000 cross track elements and >200 spectral channels not identified. Not well suited for wide or moderate field of view. Requires IMC.
- Liquid Crystal Tunable & Acousto-Optical Tunable
 - Time sequential acquisition undermines uniformity. Low TRL, polarization sensitive. Limited spectral range. Requires IMC.



HyspIRI VSWIR Concept







Concept Technologies are Proven Work Continues to Lower Risk

- 1) Uniform Offner spectrometer (Mouroulis Design)
- 2) Finely adjustable optics and detector mounts that can be locked within fraction of a micron (0.1 microns)
- 3) Electron beam fabricated gratings (large ruling period)
- 4) Electron beam fabricated air slits (non-uniformity < .05 microns)
- 5) Alignment and calibration sources and methodologies to achieve and verify requirements.





















The HyspIRI VSWIR Concept Enables the Full Set of Decadal Survey Science and Science Applications and Climate Science



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Questions?





Backup



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