

Simulating the response of the HypsIRI 4 μm channel to active lavas, using EO-1 Hyperion

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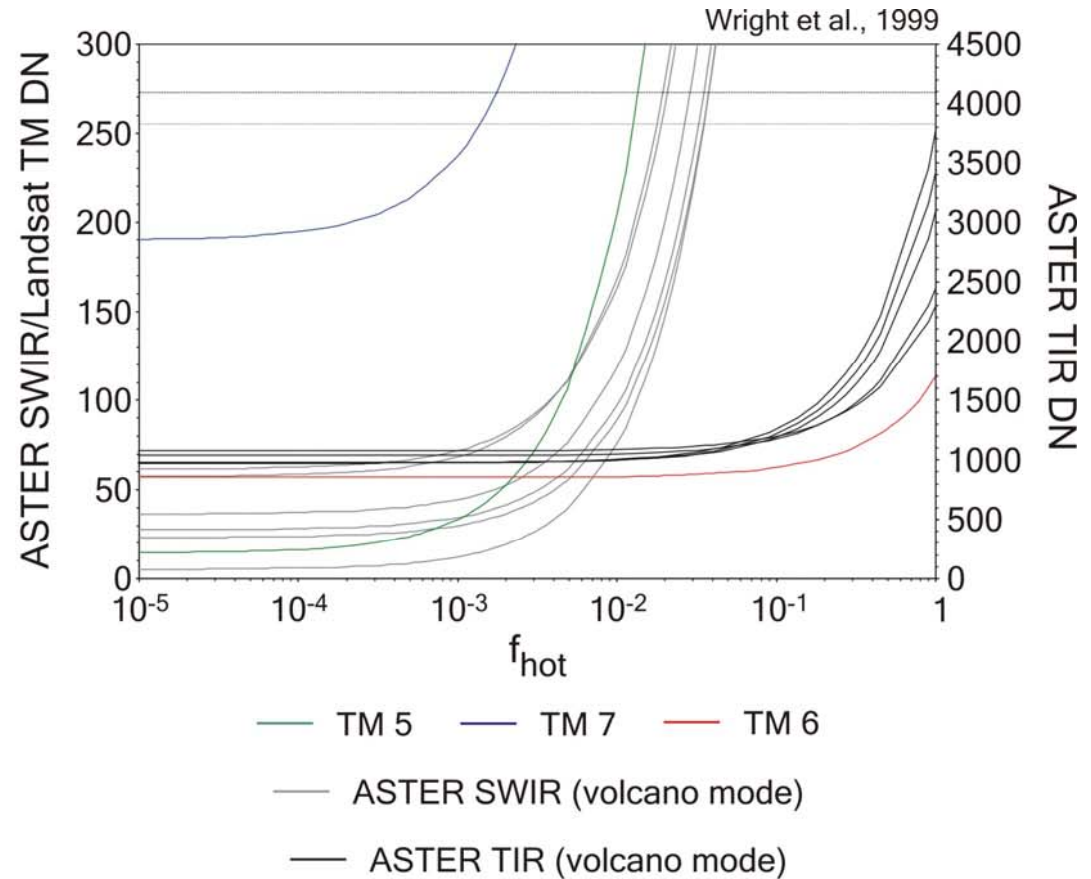
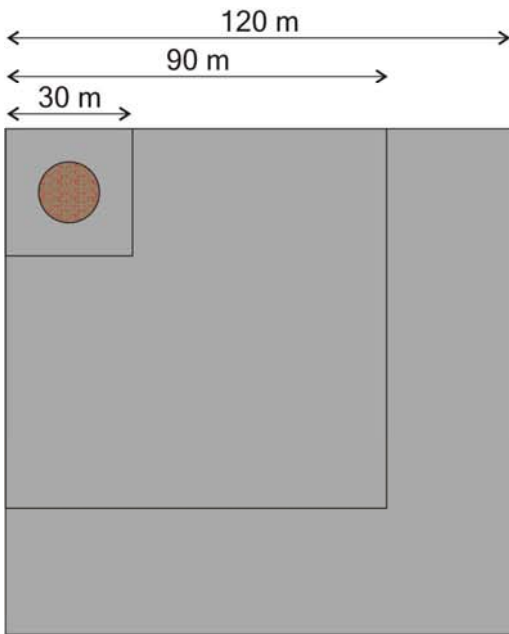
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Topics

- Data driven predictions of surface leaving radiance from active lavas based on non-linear unmixing of Hyperion data
 - Simulating HypsIRI at-satellite radiance from these predictions
- The propensity of HypsIRI's 4 μm channel to saturate over terrestrial lavas: recommendation for setting L_{max}
 - Some science results

What is the surface leaving radiance from Earth's active lavas?

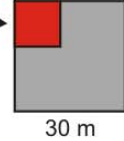
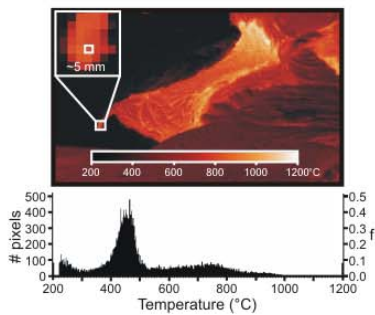
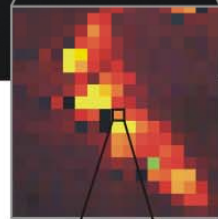
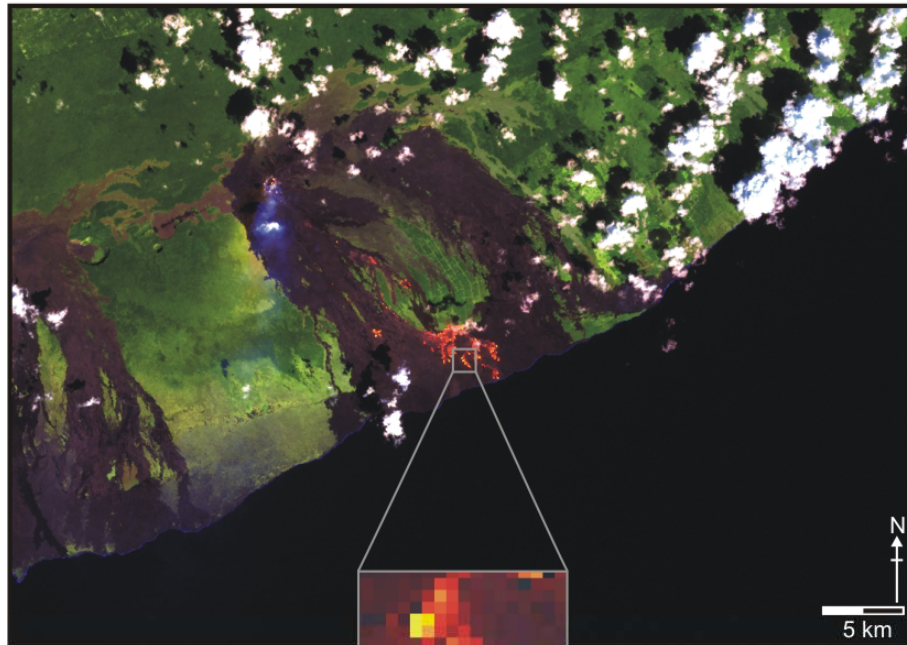
a) simple modeling



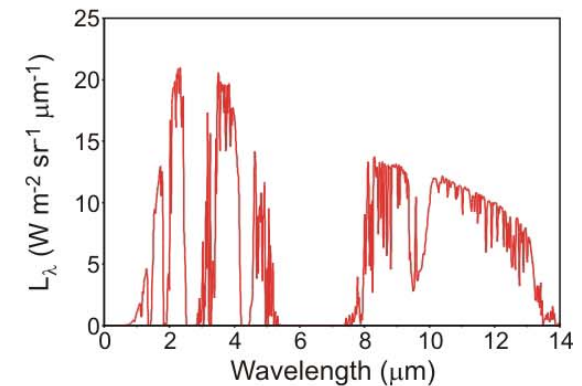
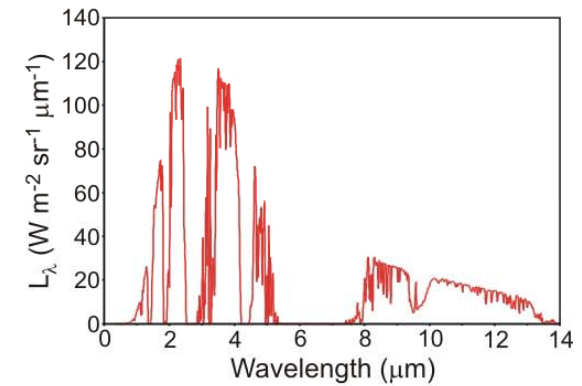
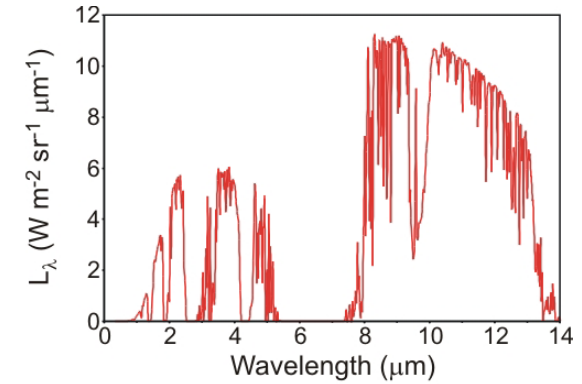
- Largely unconstrained by data
- How large? How hot? In what proportions?

What is the surface leaving radiance from Earth's active lavas?

b) combining modeling with real data

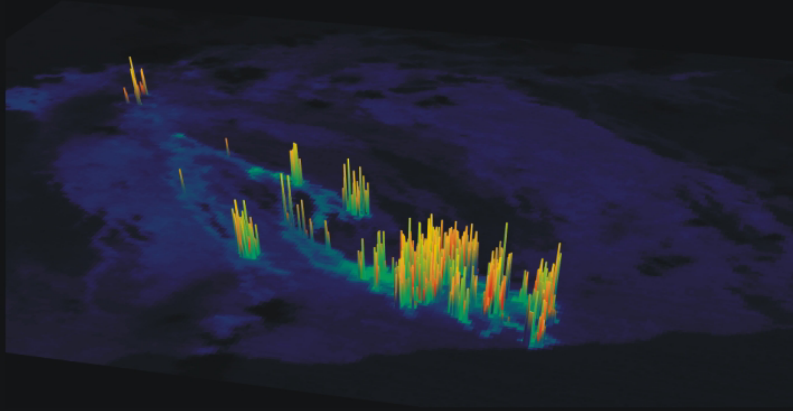


ETM+ 6 (10.4–12.5 μm)

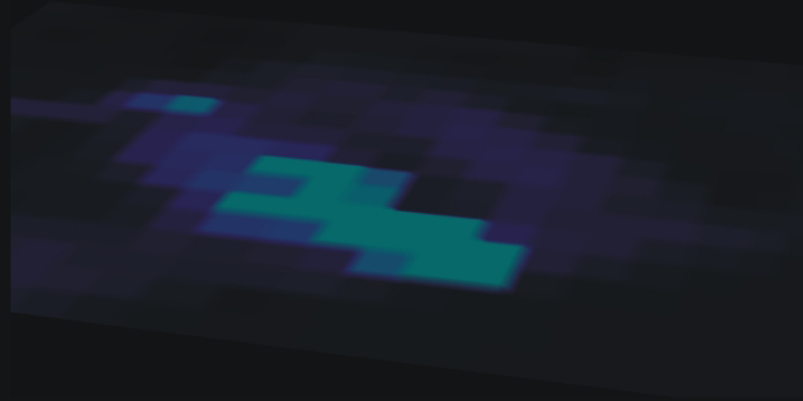


14 February 2000

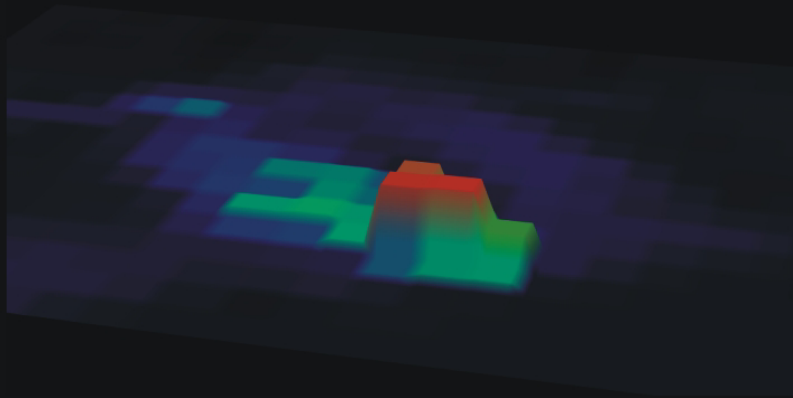
Simulated 4 μm spectral radiance



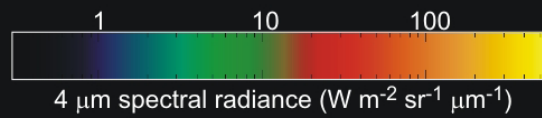
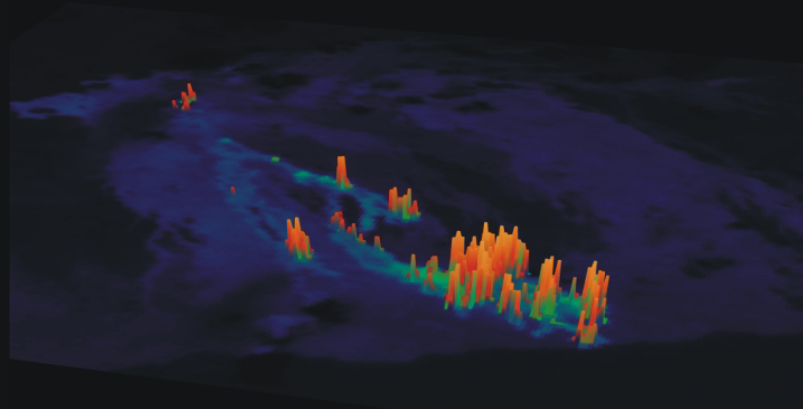
MODIS band 22



MODIS band 21

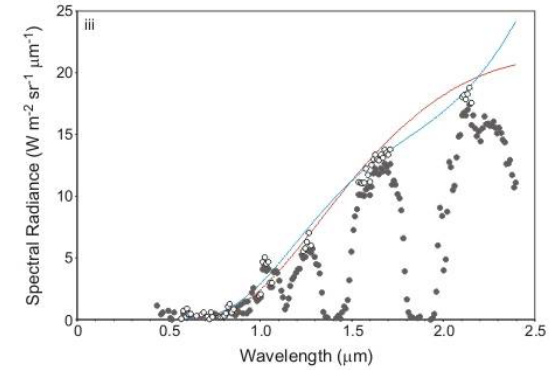
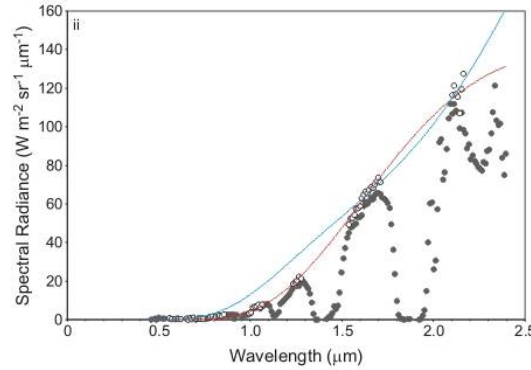
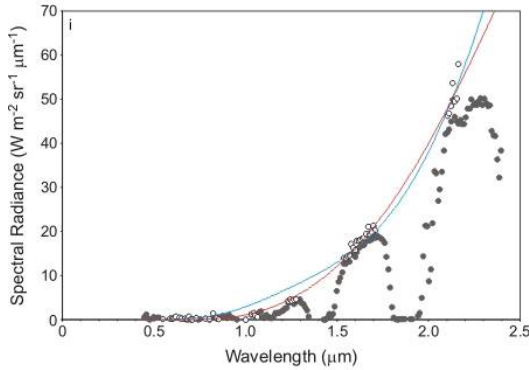


HyspIRI 4 μm channel



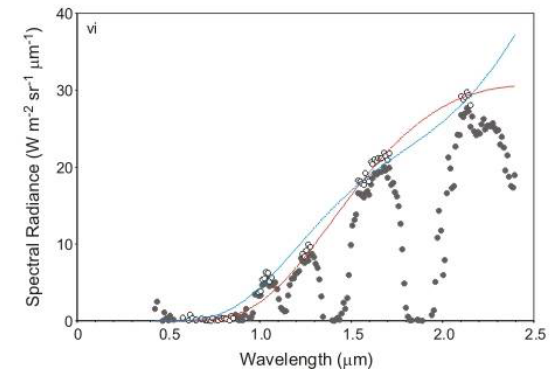
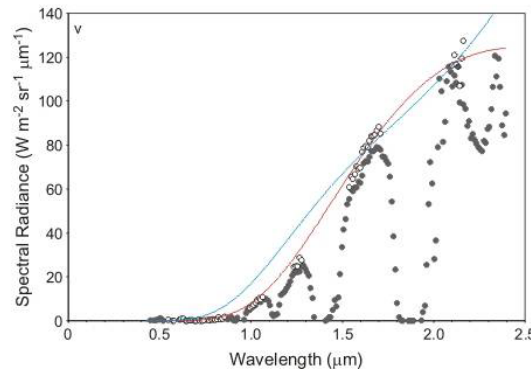
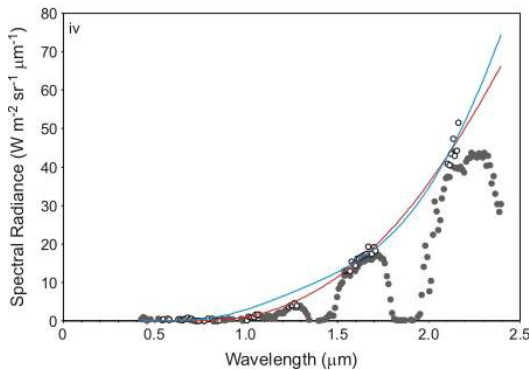
What is the surface leaving radiance from Earth's active lavas?

c) simulating surface leaving radiance using real data

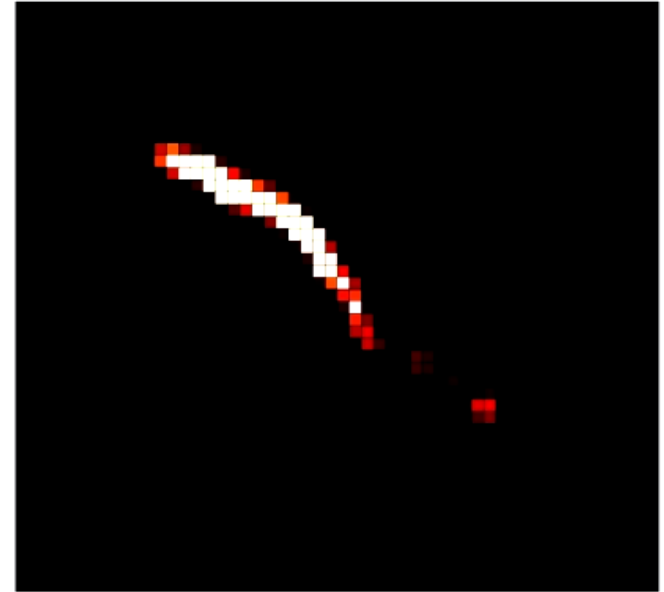
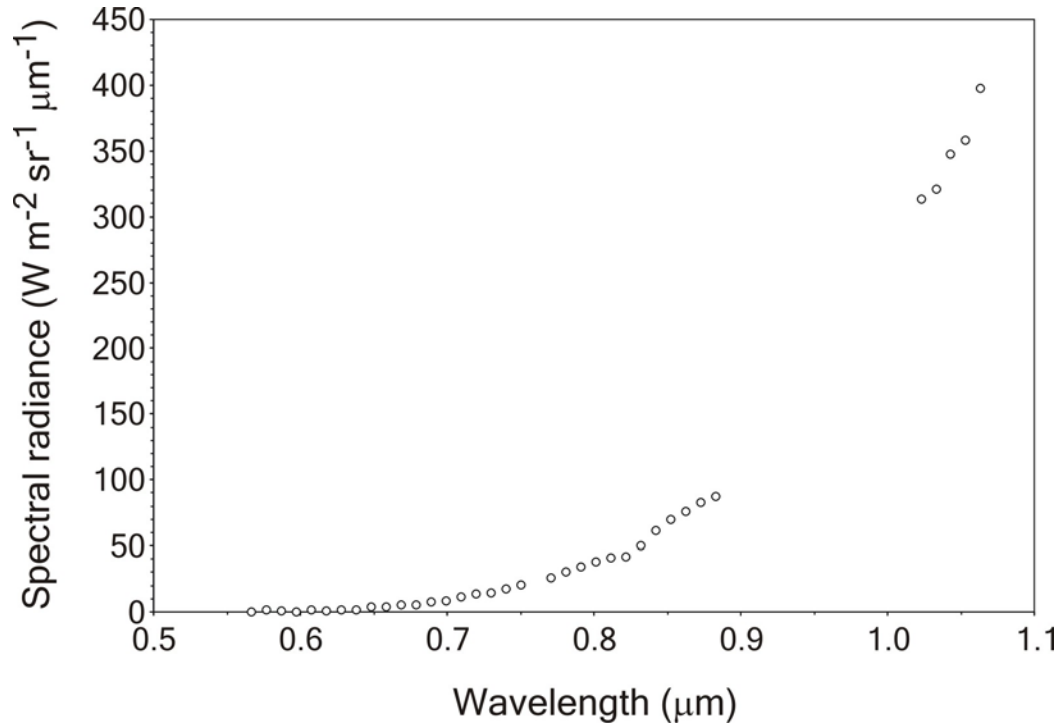


Wright et al., 2010

Pixel	i	ii	iii	iv	v	vi
Dual-band						
T_1	1150°C	1150°C	1150°C	1150°C	1150°C	1150°C
f_1	0.0006	0.0027	0.0006	0.0006	0.0035	0.0009
T_2	328°C	351°C	233°C	320°C	328°C	254°C
f_2	0.9994	0.9973	0.9994	0.9994	0.9965	0.9991
rms	1.3	3.6	0.5	1.2	4.8	0.8
Unconstrained						
T_1	874°C	1067°C	1138°C	1001°C	920°C	963°C
f_1	0.003	0.0002	0.0005	0.0008	0.008	0.002
T_2	421°C	790°C	697°C	500°C	899°C	587°C
f_2	0.2	0.02	0.003	0.05	0.005	0.001
T_3	128°C	489°C	*	377°C	*	230°C
f_3	0.6	0.002	*	0.2	*	0.06
rms	0.7	1.2	0.4	0.7	1.8	0.7

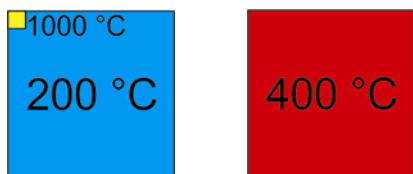
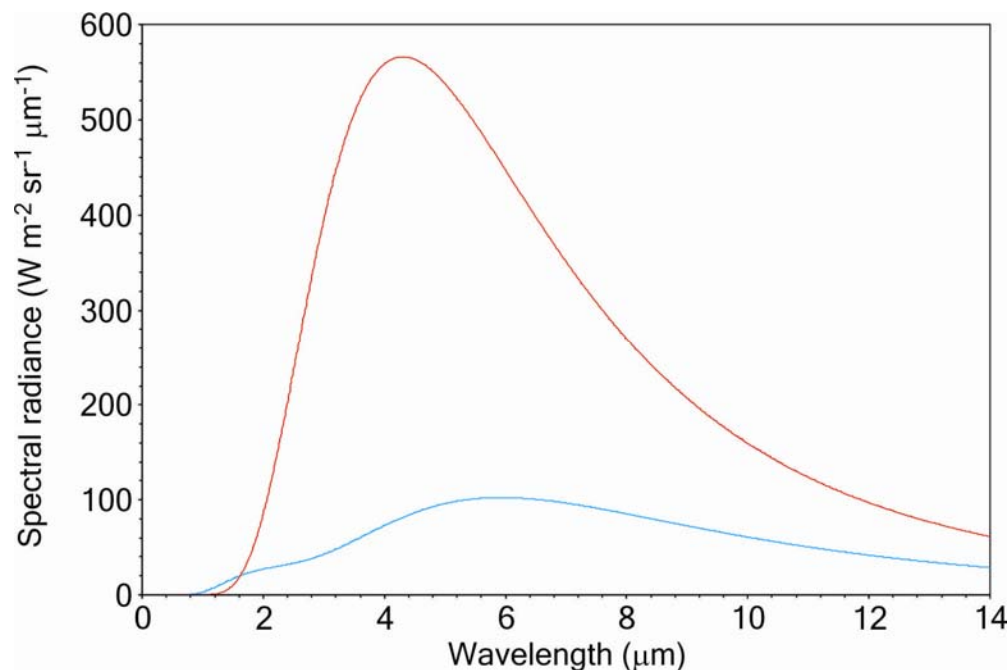
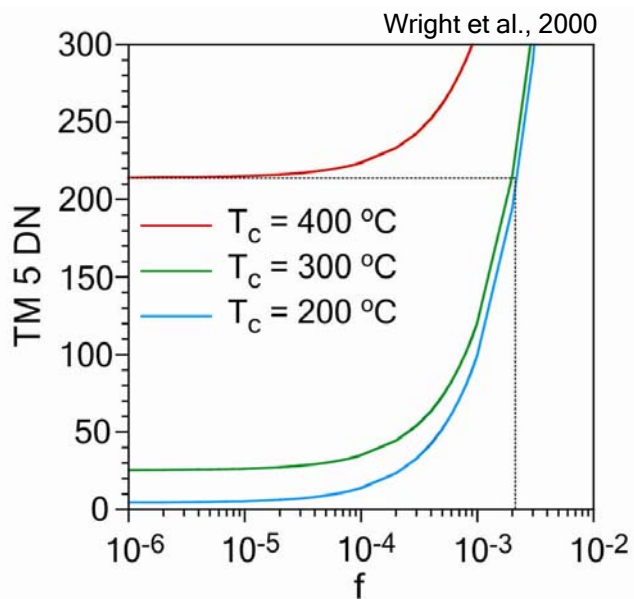


The ability to retrieve sub-pixel temperature characteristics using an imaging spectrometer



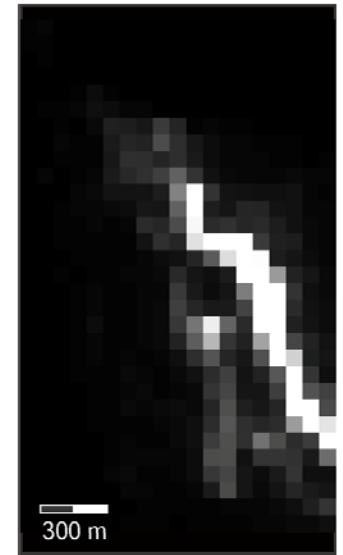
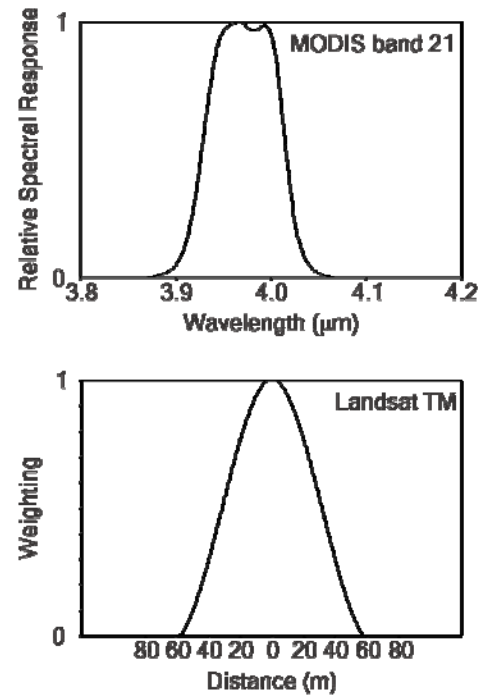
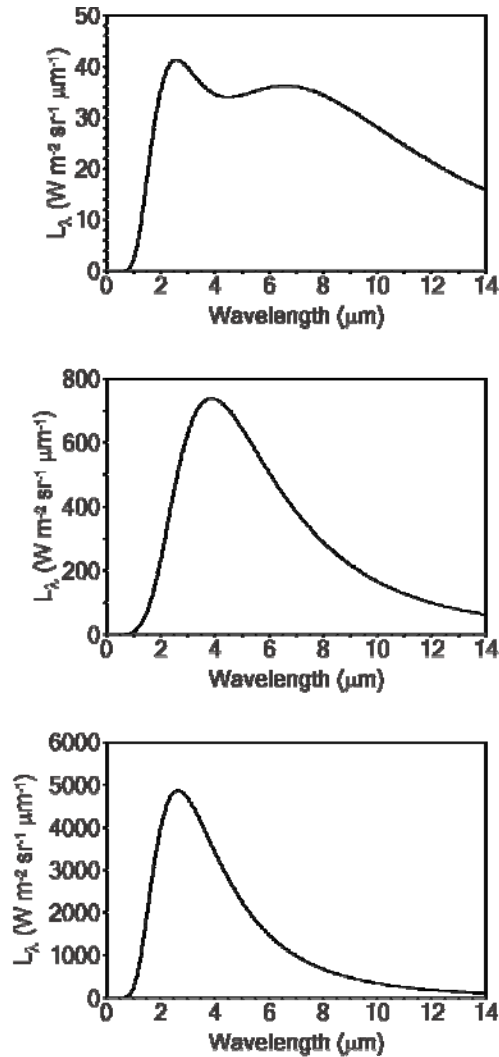
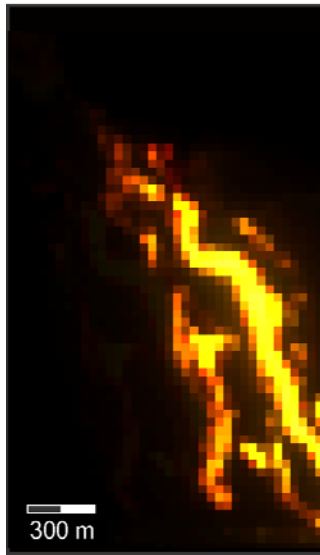
- Sub-pixel temperature characteristics can be determined for even the most radiant active lava surfaces

The desirability to quantify sub-pixel temperature mixing



- Whole pixel/"Dozier" temperatures can easily misrepresent the nature of emitted spectral radiance from thermally complex active lava surfaces

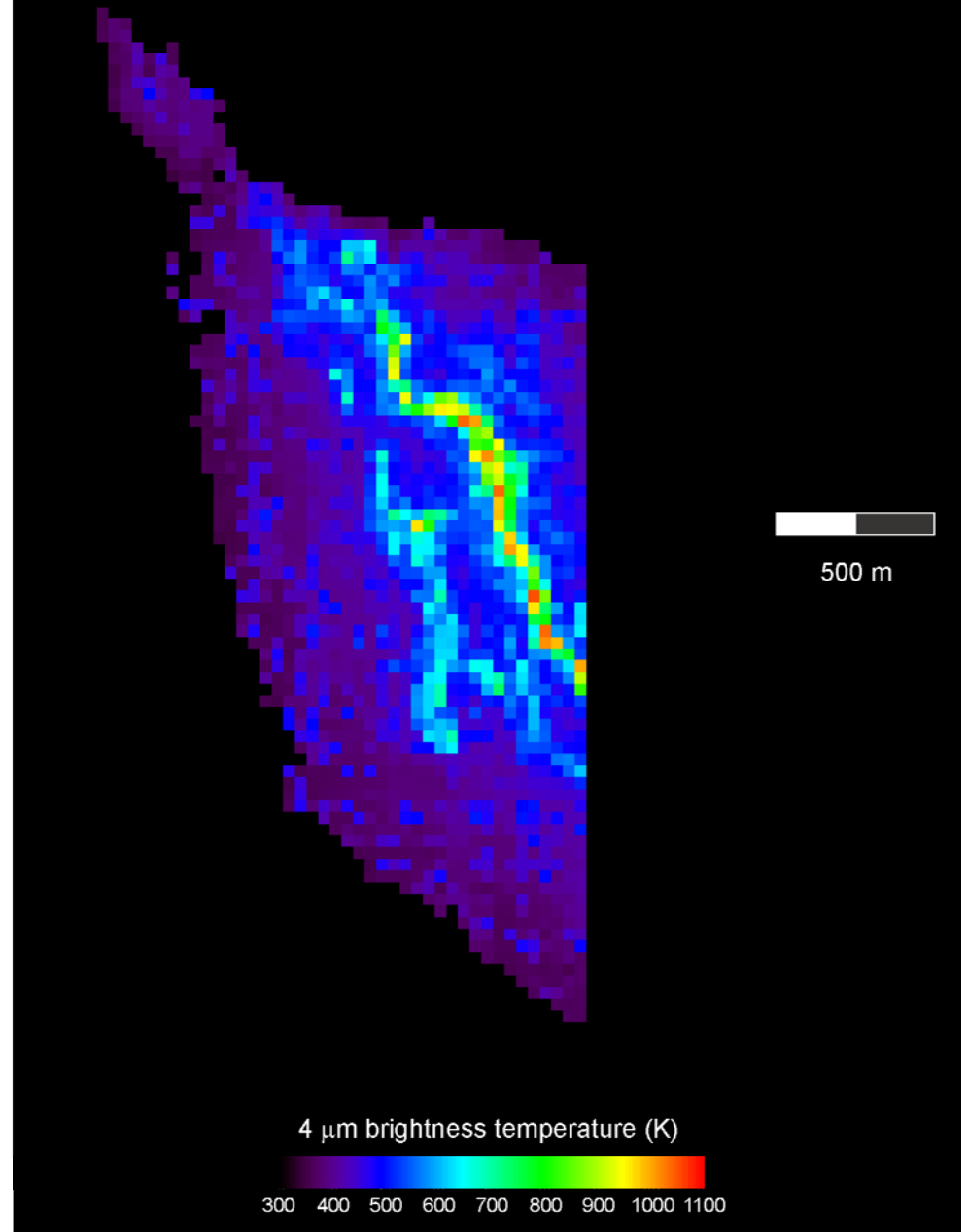
Simulating HypsIRI data from Hyperion data



- Analyzed 60 Hyperion images encompassing the full spectrum of terrestrial active lava bodies

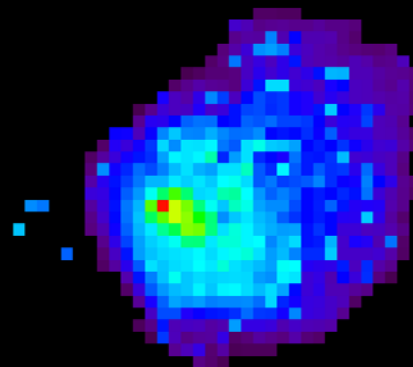


Nyamuragira (Democratic Republic of Congo)





Nyiragongo (Democratic Republic of Congo)



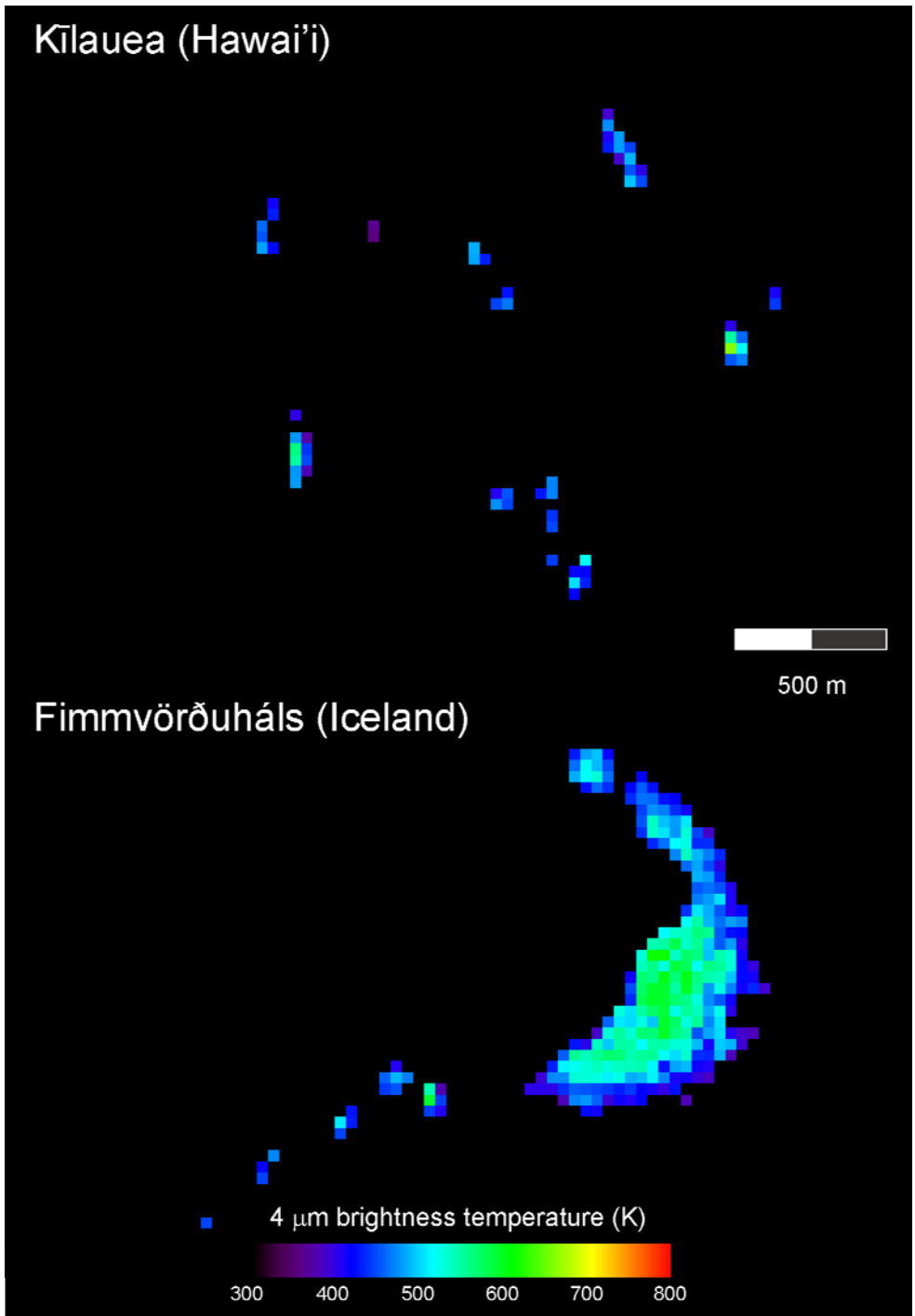
500 m

Erta Ale (Ethiopia)



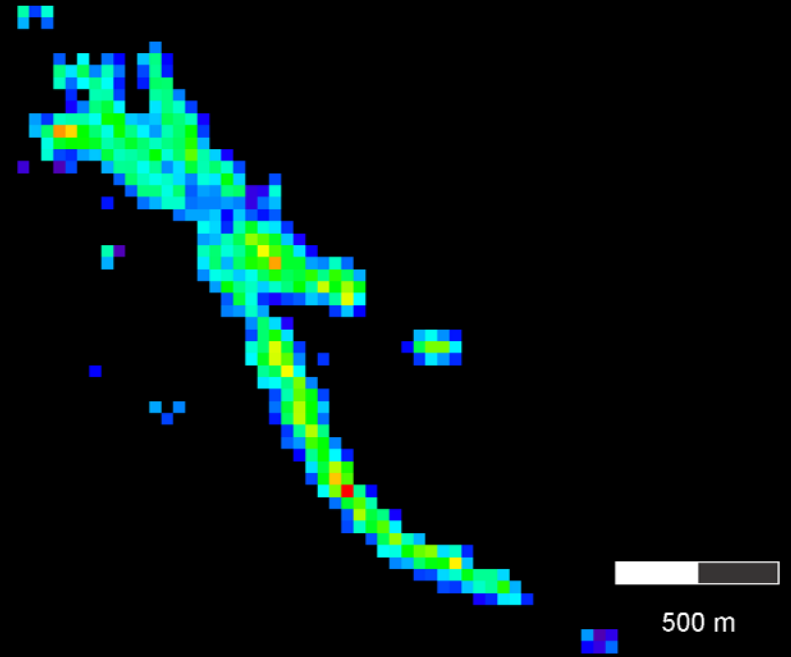
4 μm brightness temperature (K)



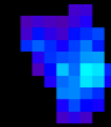




Etna (Italy)



Soufriere Hills (Montserrat)

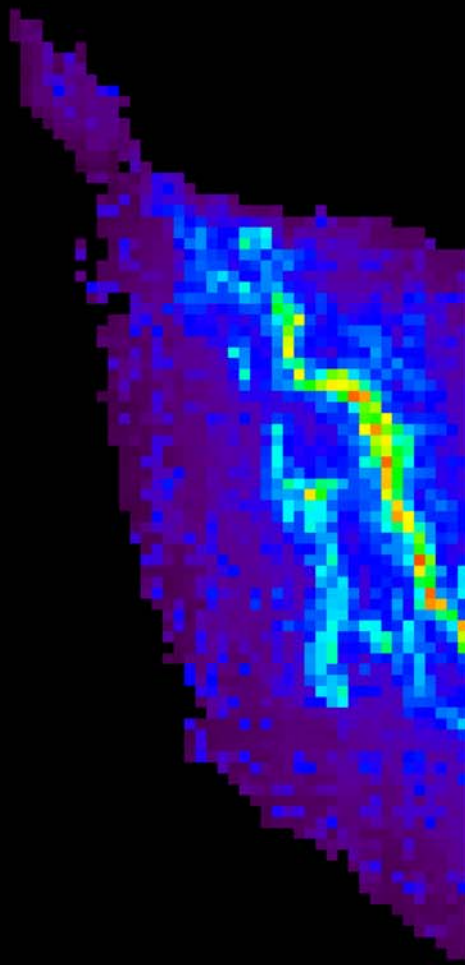


4 μm brightness temperature (K)

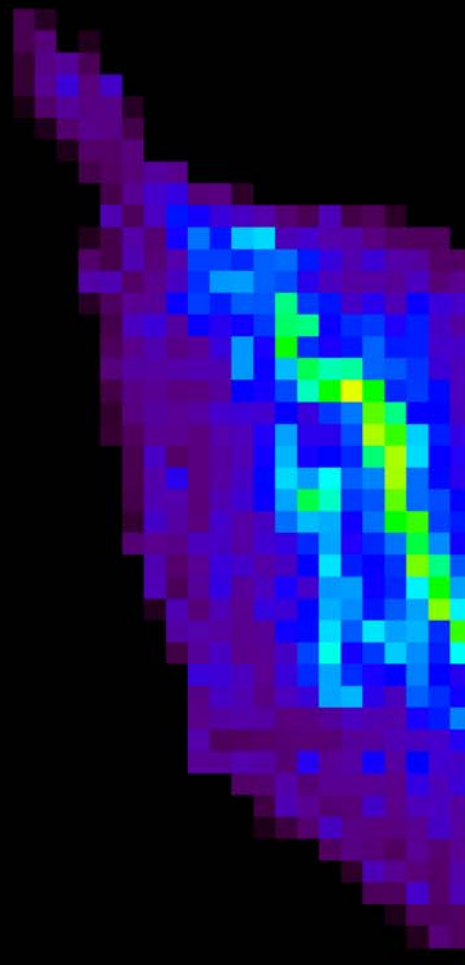


Nyamuragira (Democratic Republic of Congo)

4 μm brightness temperature (K)



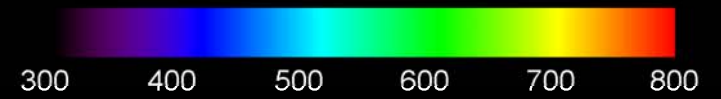
30 m



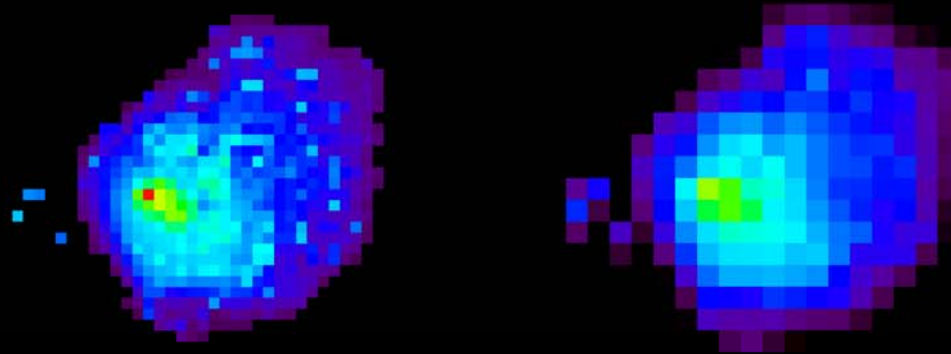
60 m



4 μm brightness temperature (K)



Nyiragongo (Democratic Republic of Congo)



500 m

Erta Ale (Ethiopia)

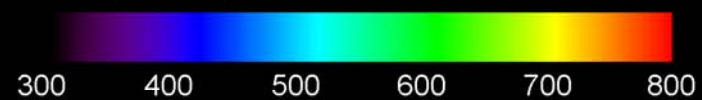


30 m

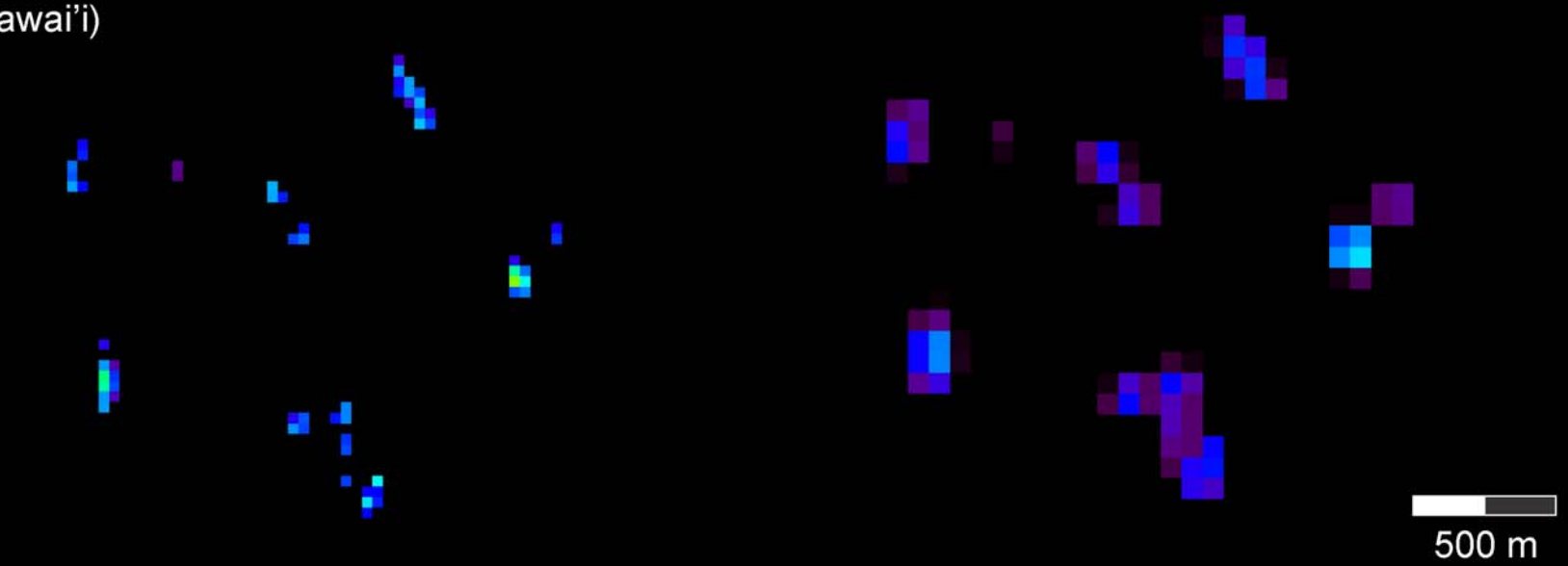


60 m

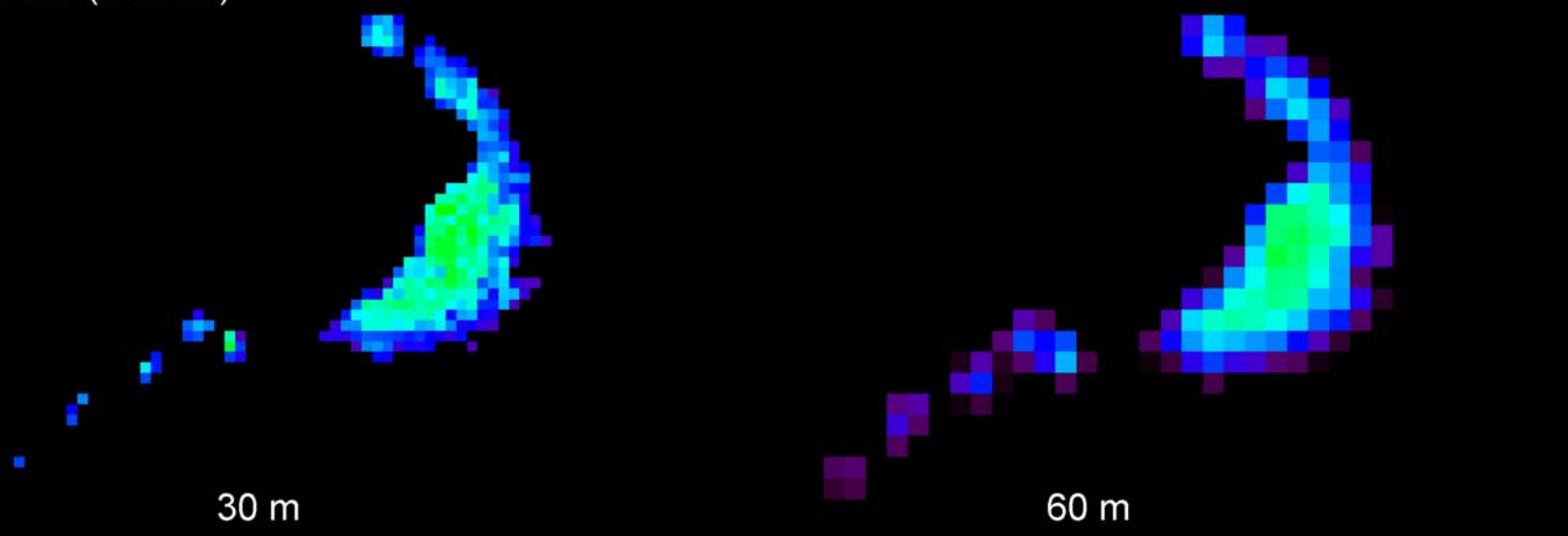
4 μm brightness temperature (K)



Kīlauea (Hawai'i)



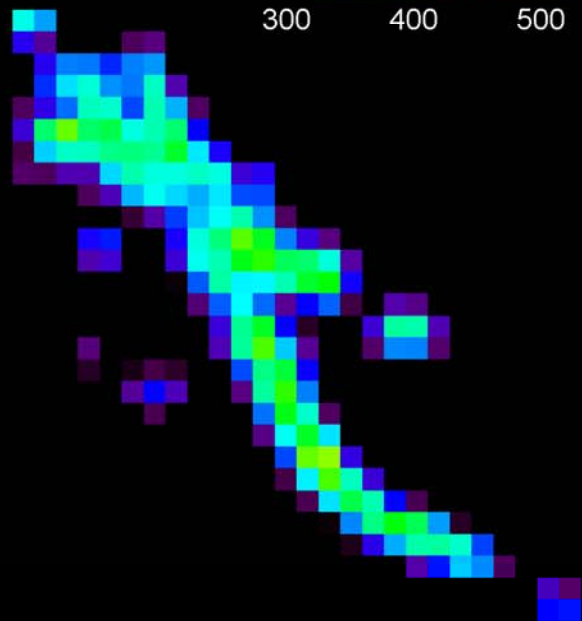
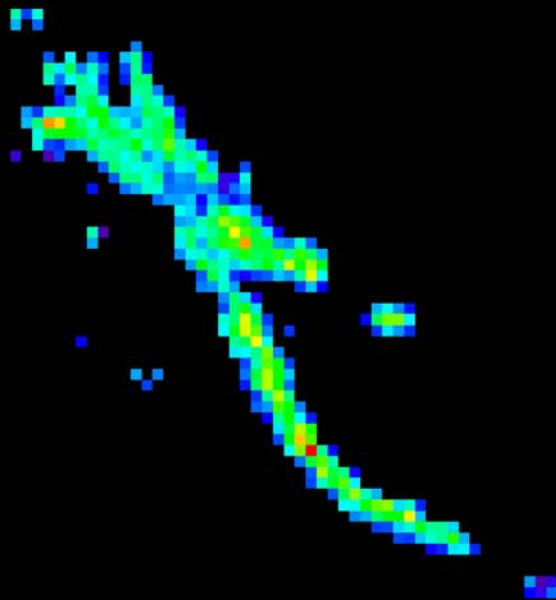
Fimmvörðuháls (Iceland)



4 μm brightness temperature (K)



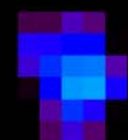
Etna (Italy)



Soufriere Hills (Montserrat)

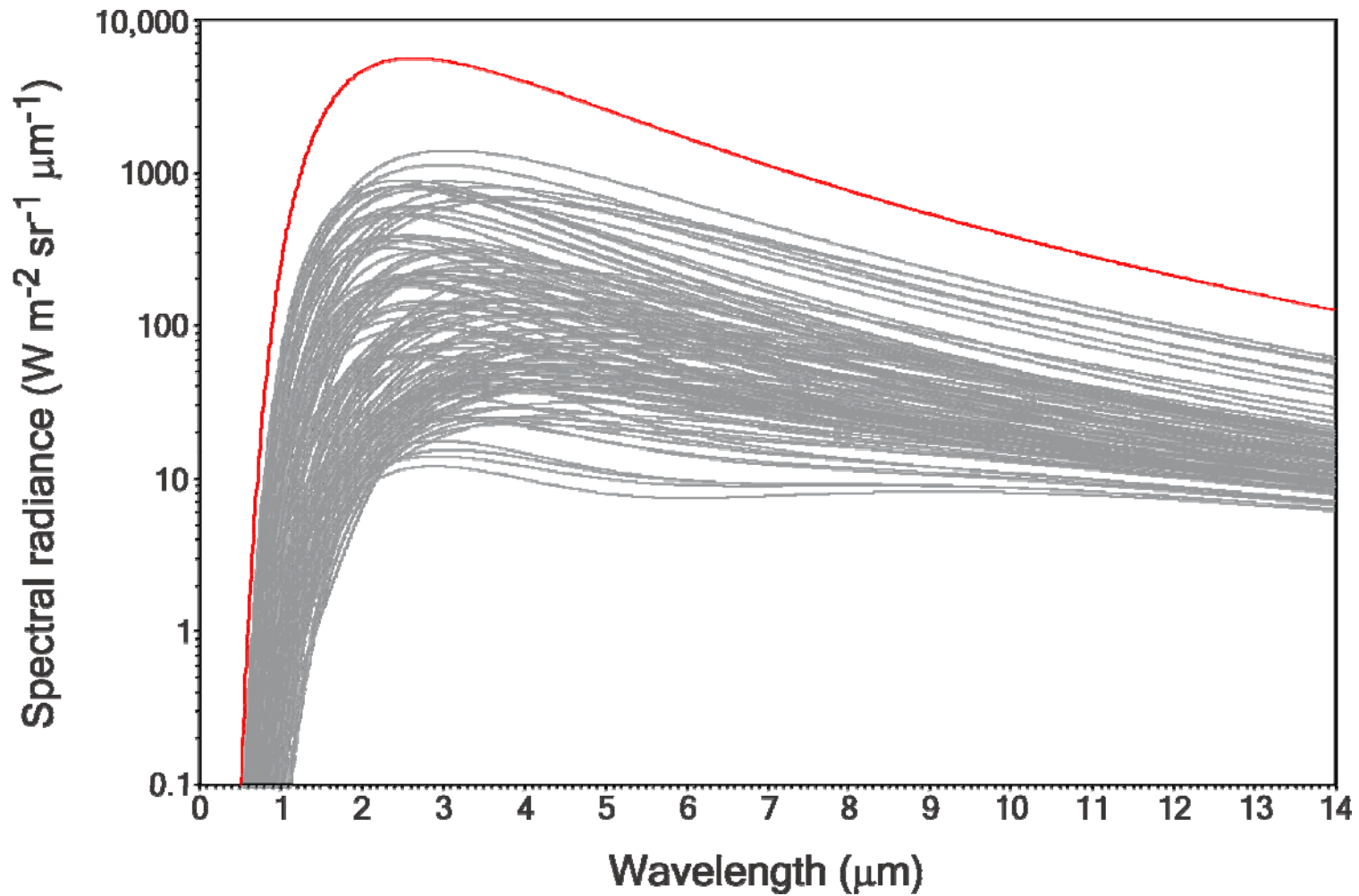


30 m



60 m

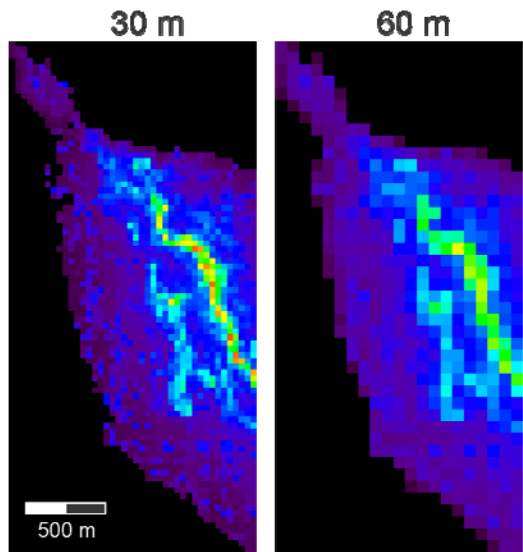
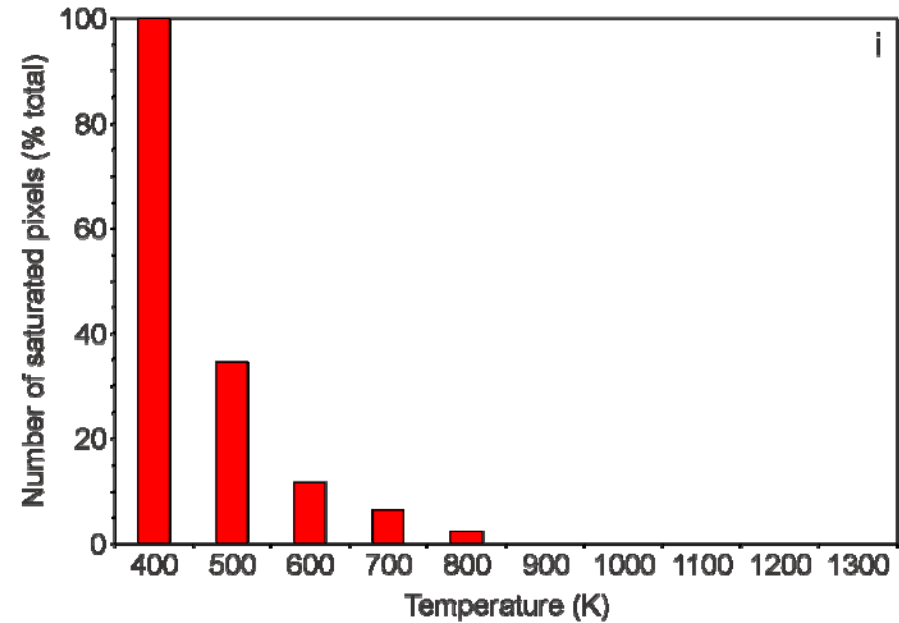
T_{sat} for the HypsIRI 4 μm channel



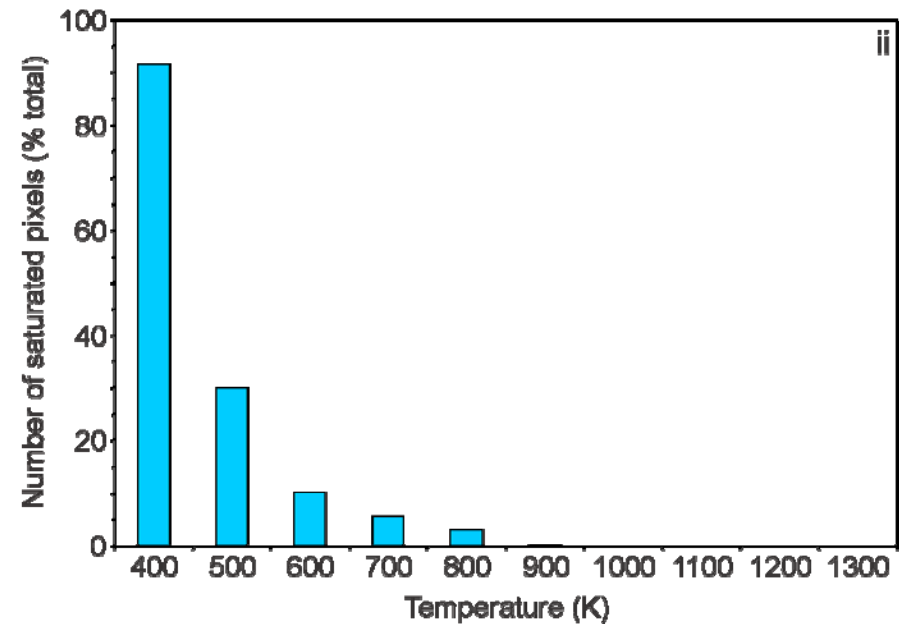
- What should the HypsIRI 4 μm channel saturation be set to?

Incidence of saturation as a function of assumed T_{sat}

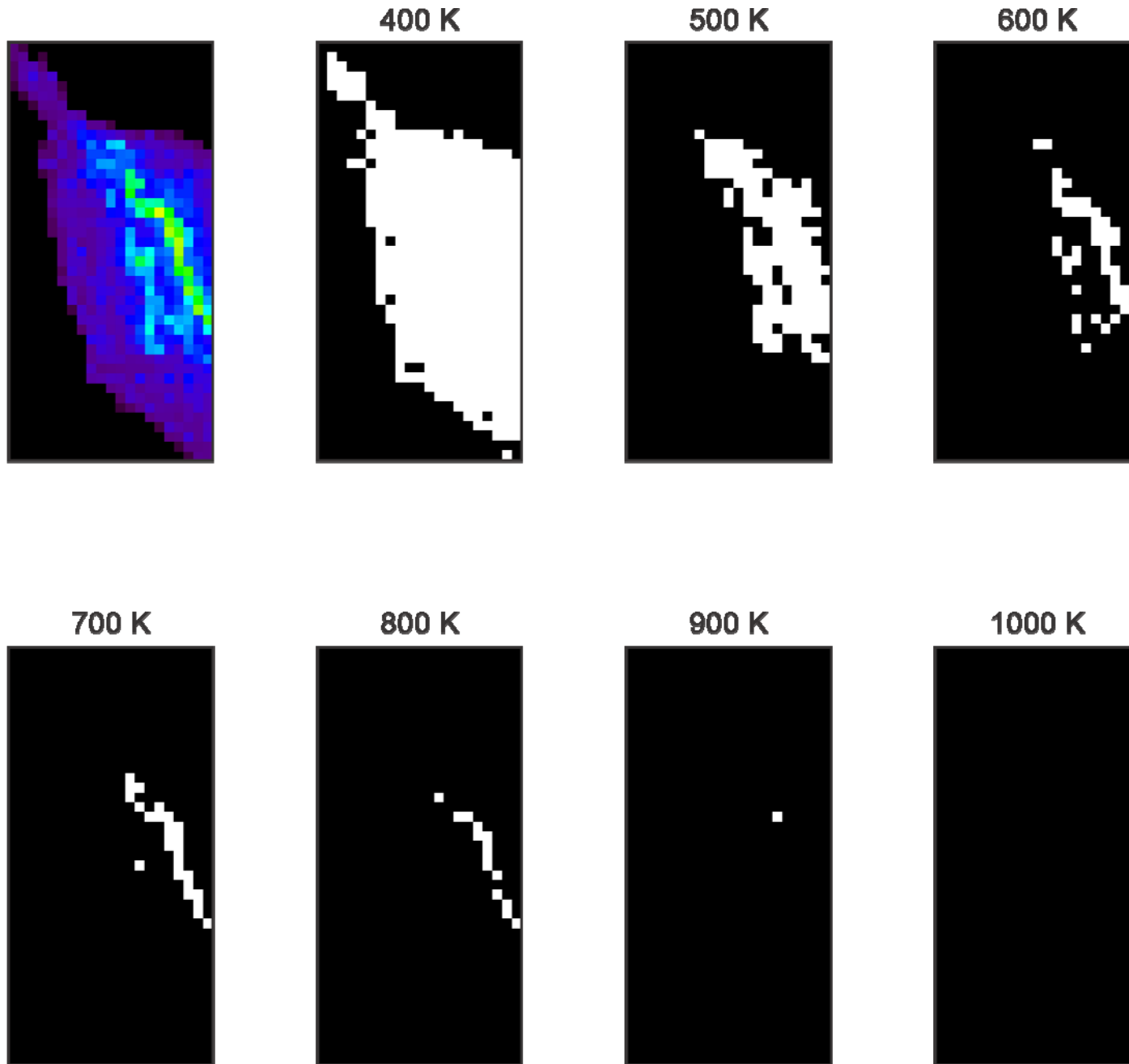
T_{sat} (K)	# _{sat} (i)	# _{sat} (ii)
400	471	432
500	163	142
600	55	48
700	30	26
800	11	14
900	0	1
1000	0	0



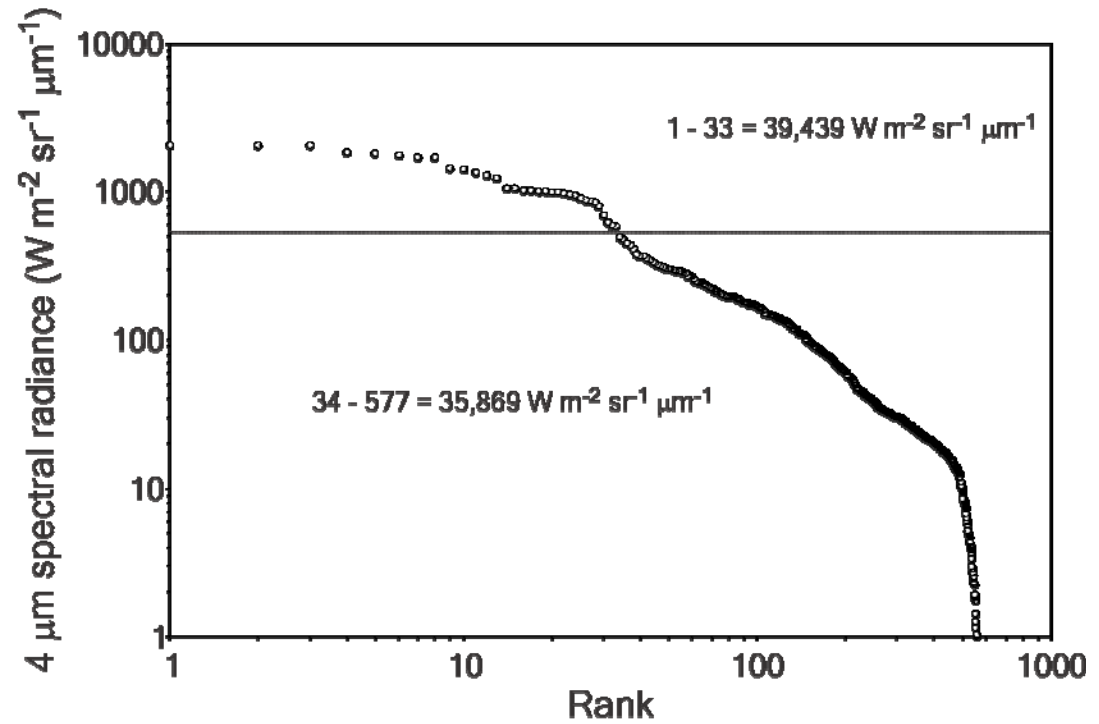
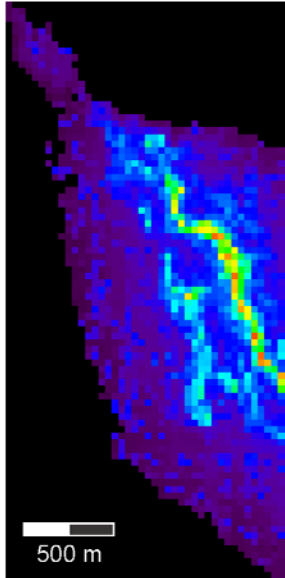
4 μm brightness temperature (K)



The important pixels are the ones which drive T_{sat} higher

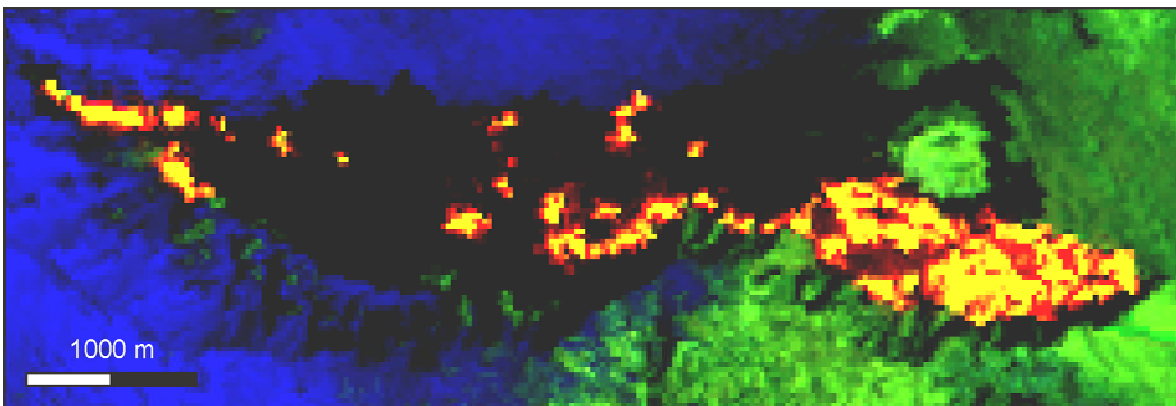
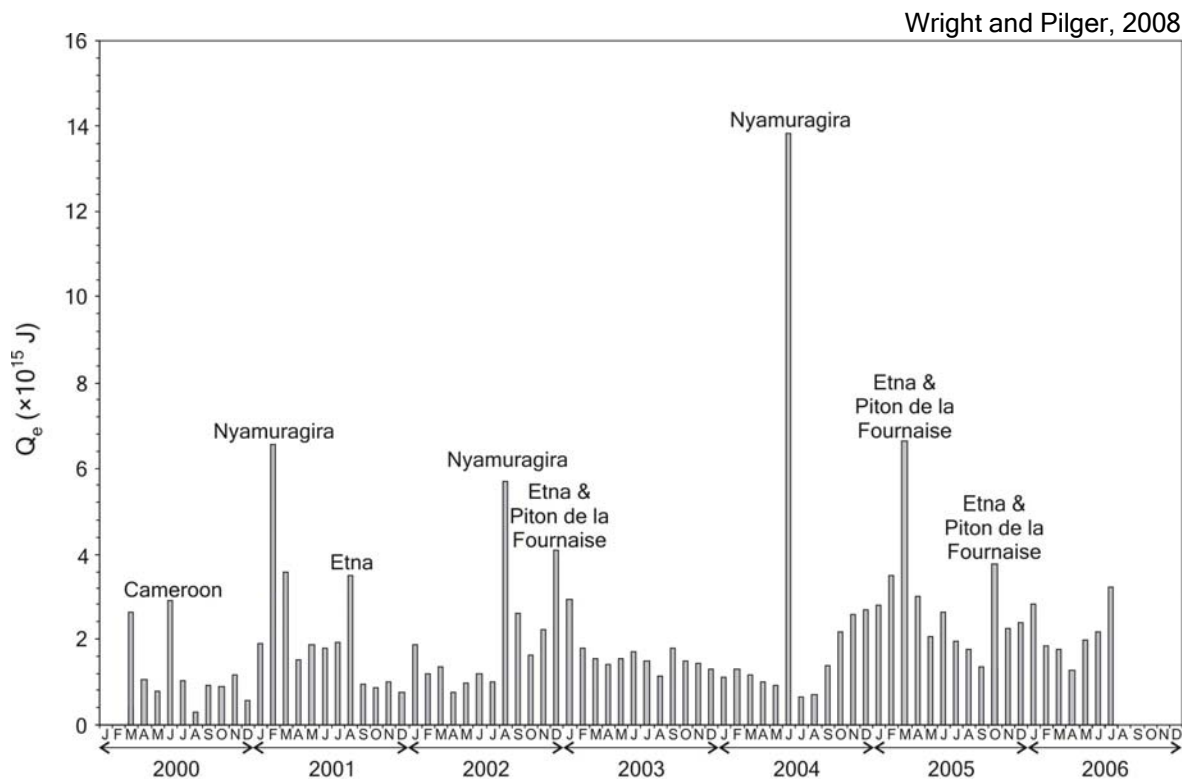


Recommendation: HypsIRI 4 μm T_{sat} not less than 1100 K



Aggregation of four most radiant pixels in simulated 30 m data set yields a maximum temperature of 1041 K at 60 m

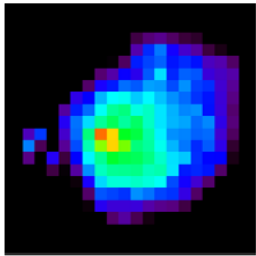
Large lava-flow-forming eruptions are not uncommon



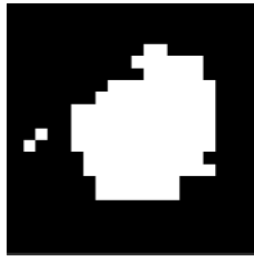
Mount Etna, 1991-1993 eruption

Lower predicted incidence of saturation over lava lakes and domes

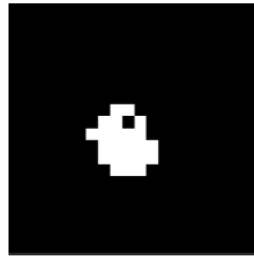
Nyiragongo



400 K



500 K



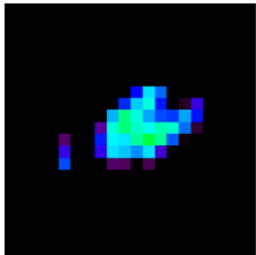
600 K



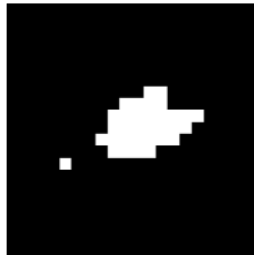
700 K



Shiveluch



400 K



500 K

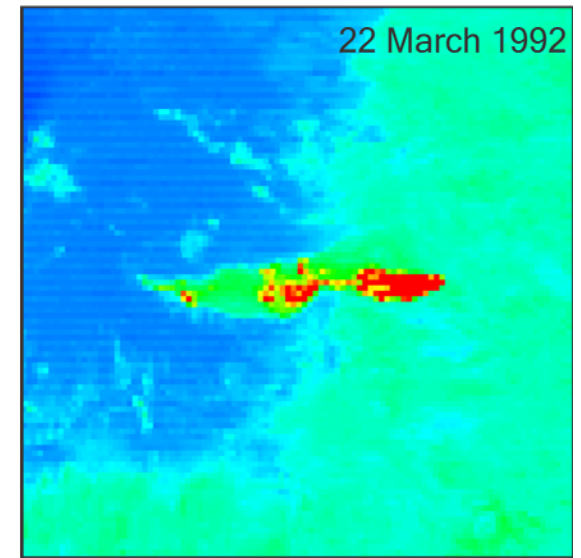
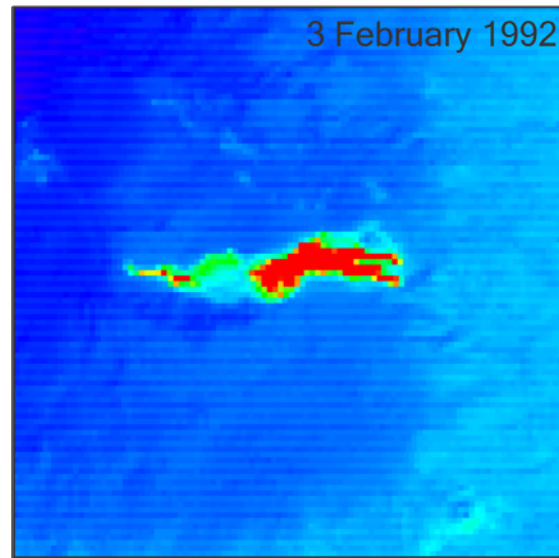
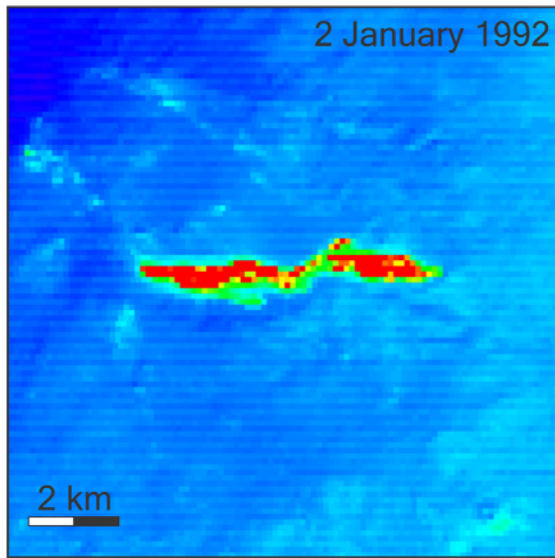


600 K



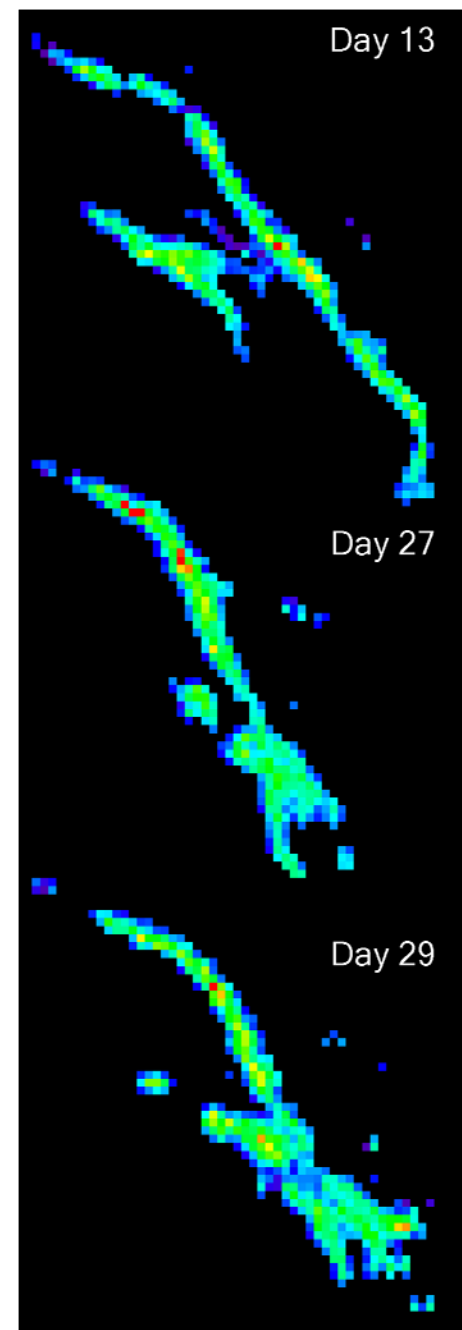
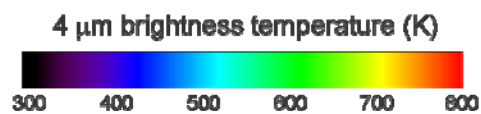
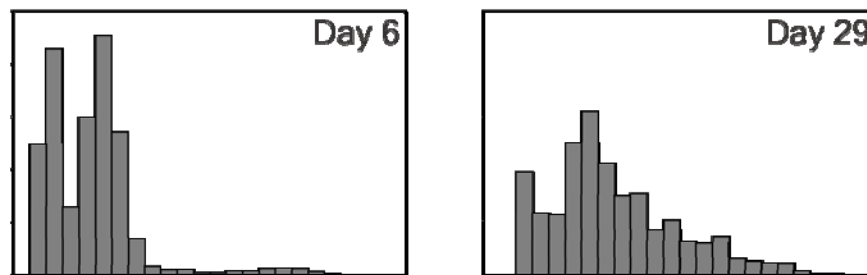
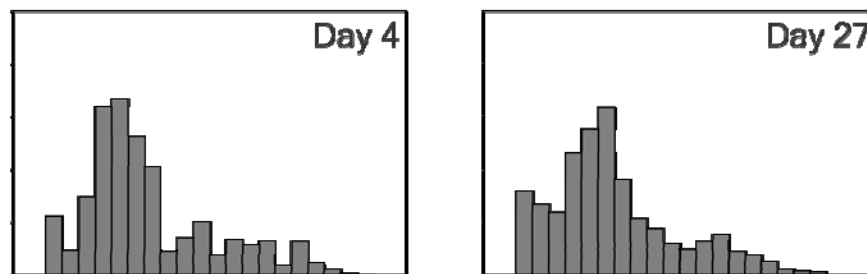
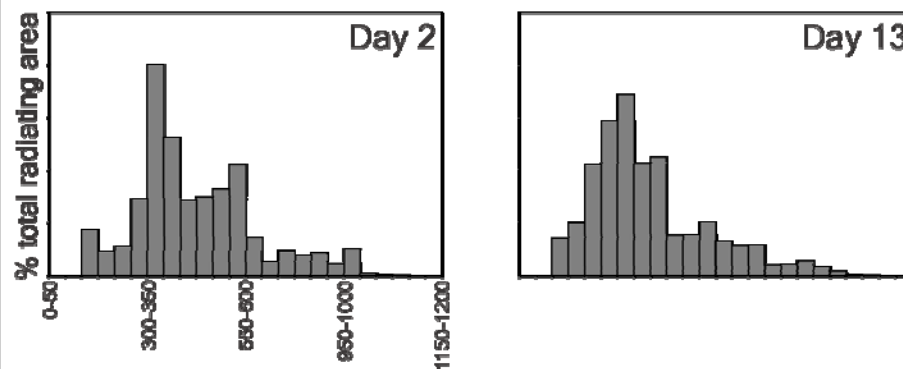
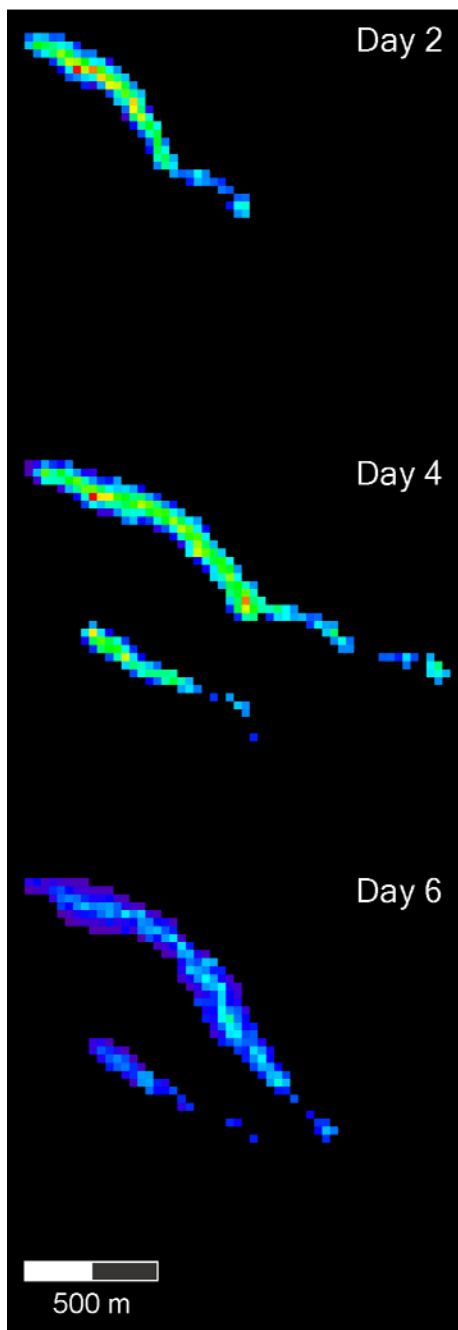
HyspIRI 11-14 μm data should prove useful for volcanologists

Mount Etna, Sicily

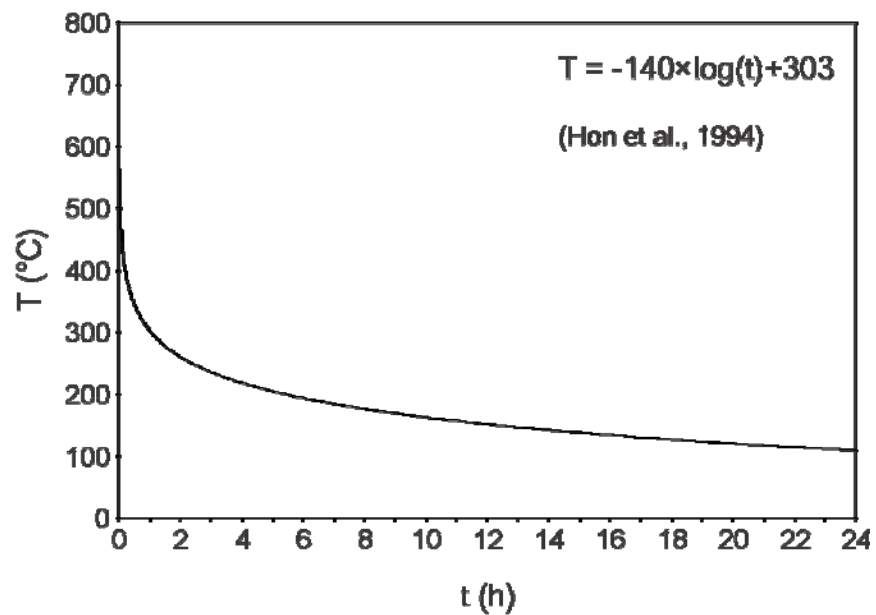
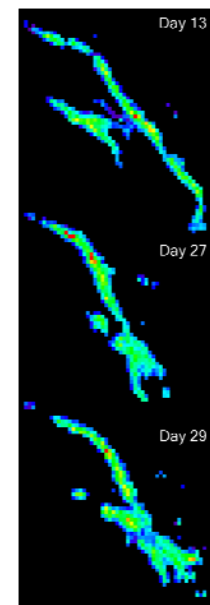
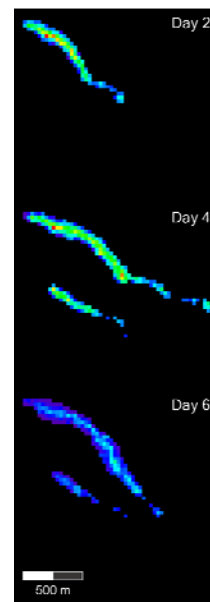
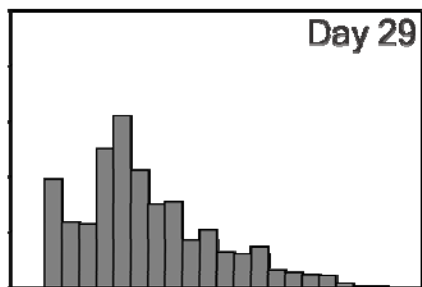
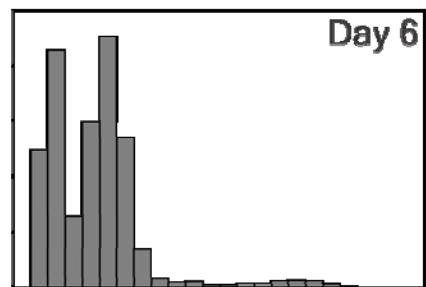
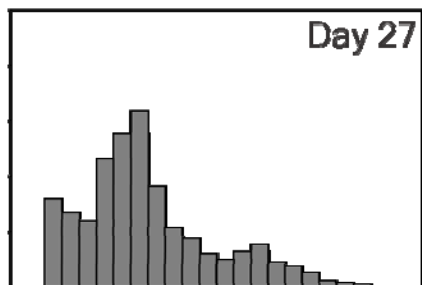
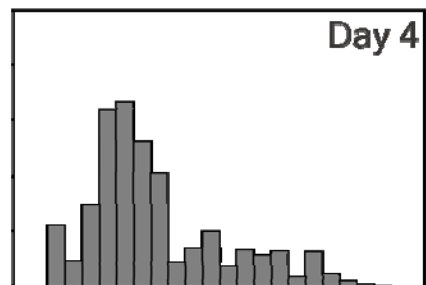
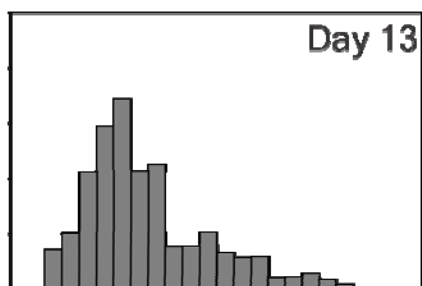
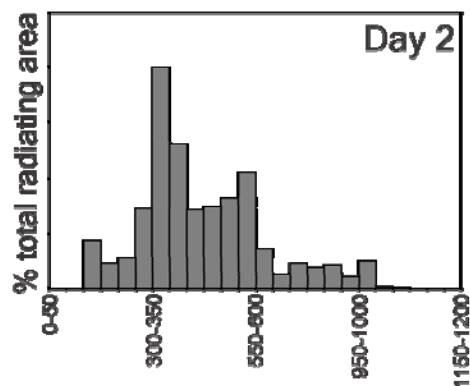


Red pixels are saturated in Landsat TM band 6, indicating a whole pixel temperature exceeding $\sim 80^{\circ}\text{C}$.

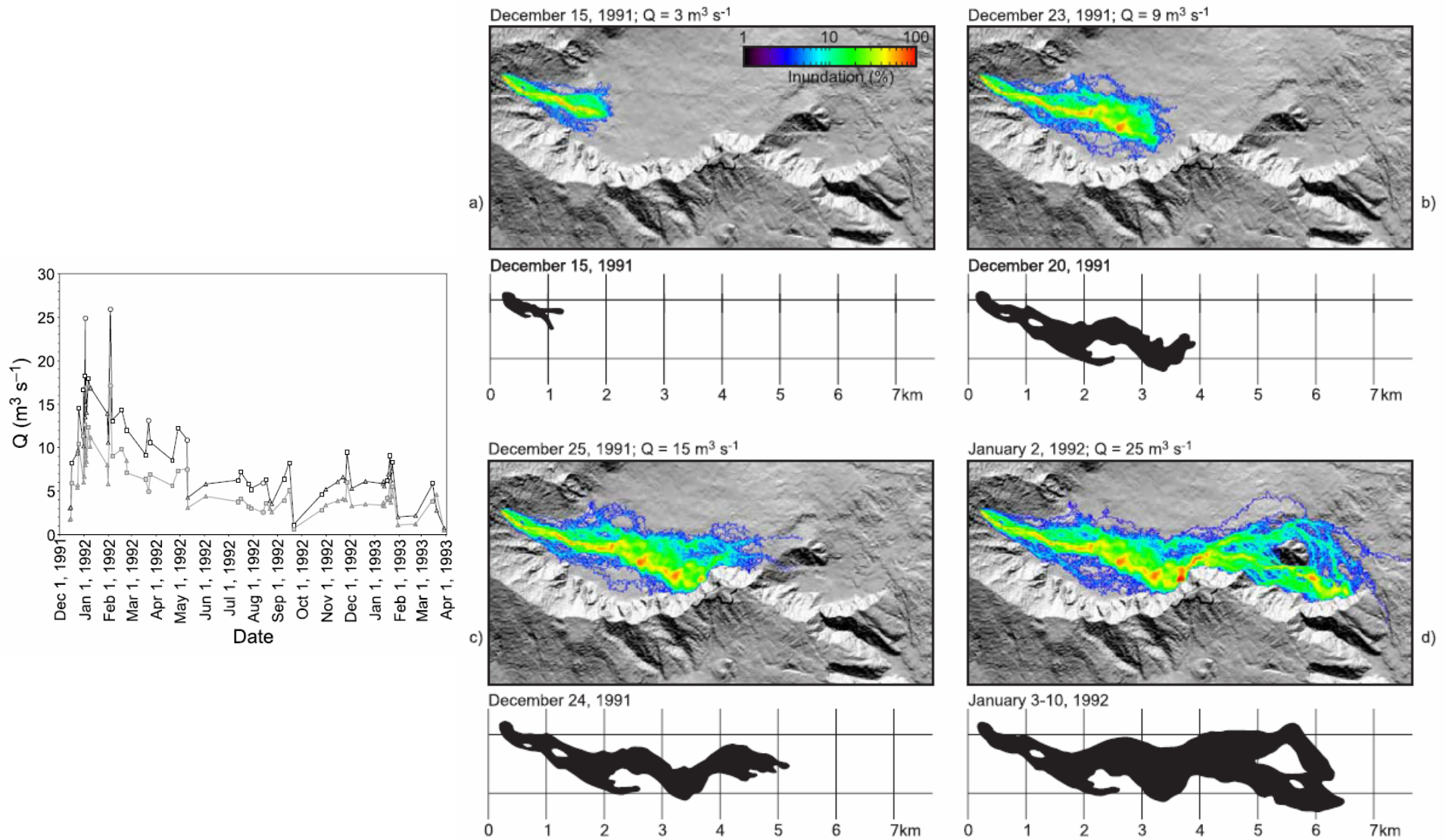
Temporal evolution of lava flow surface temperatures



Lava temperature, age, and process that thermally renew active lava surfaces



Driving numerical lava flow hazard predictions using satellite data



The temperature of Earth's volcanoes

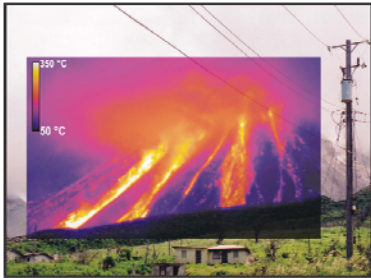
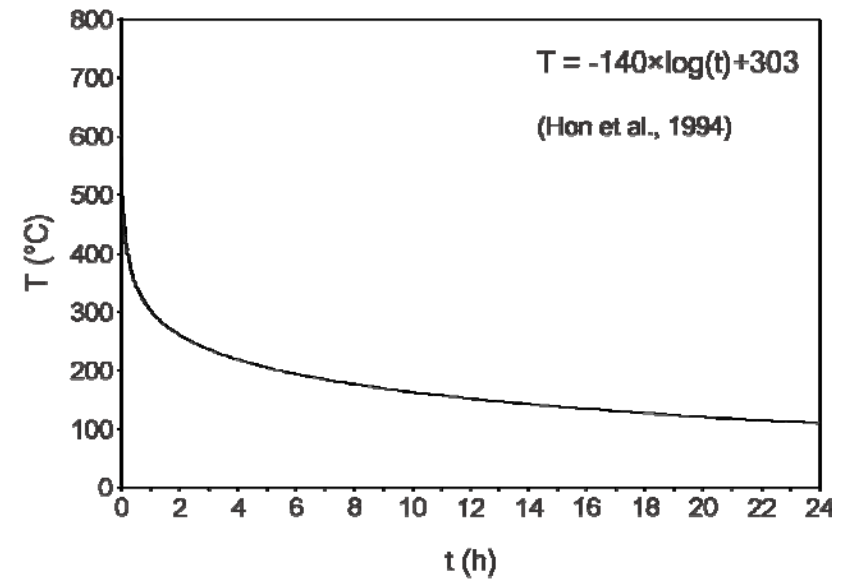
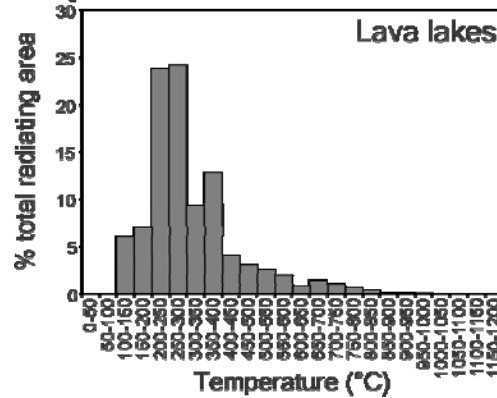
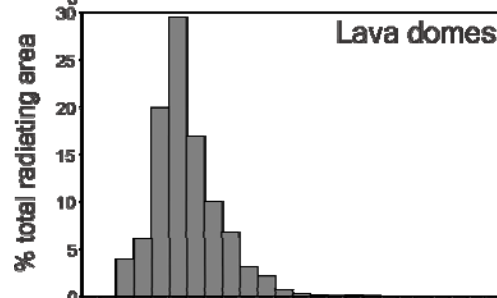
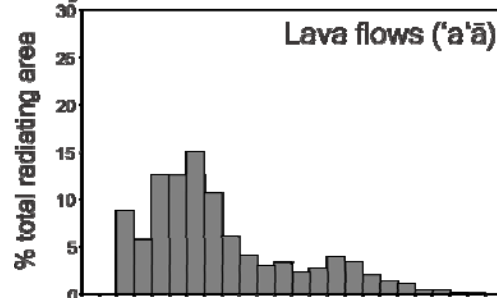
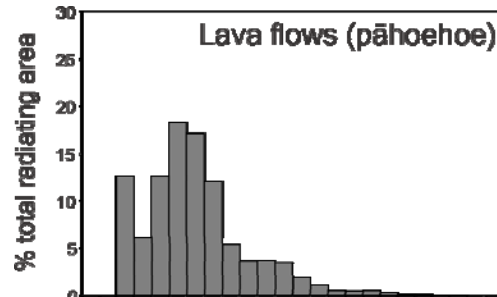
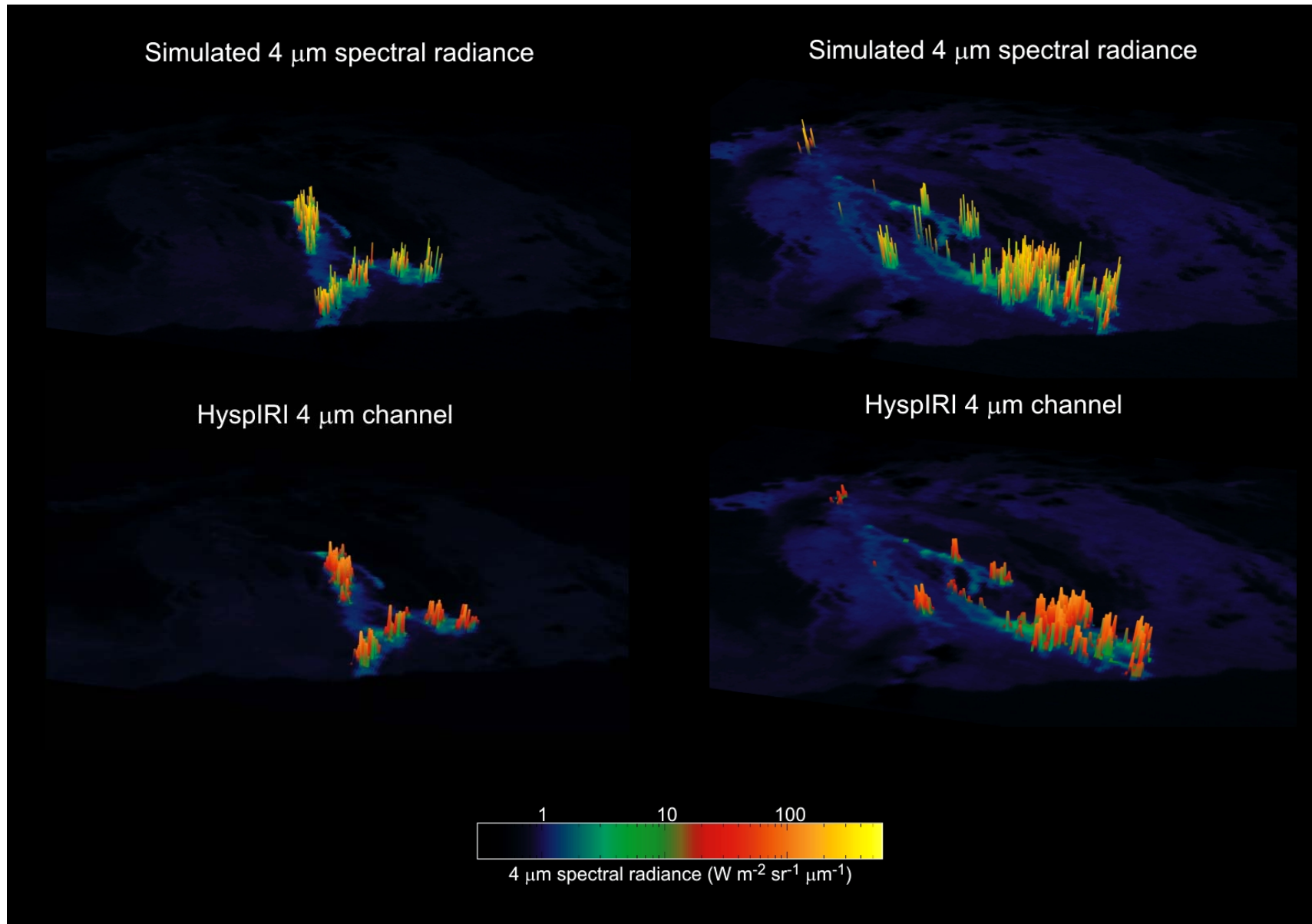


Photo: C. Oppenheimer



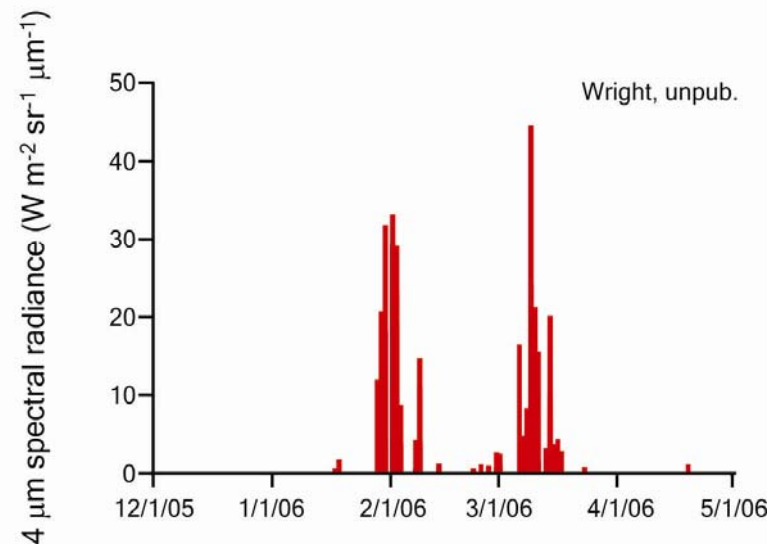
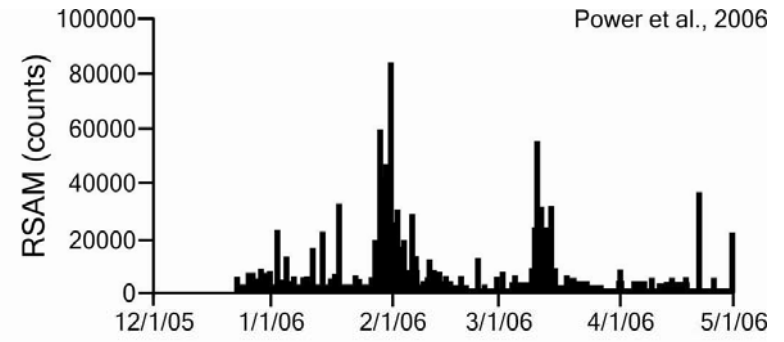
Simulating the performance of a detection algorithm



End

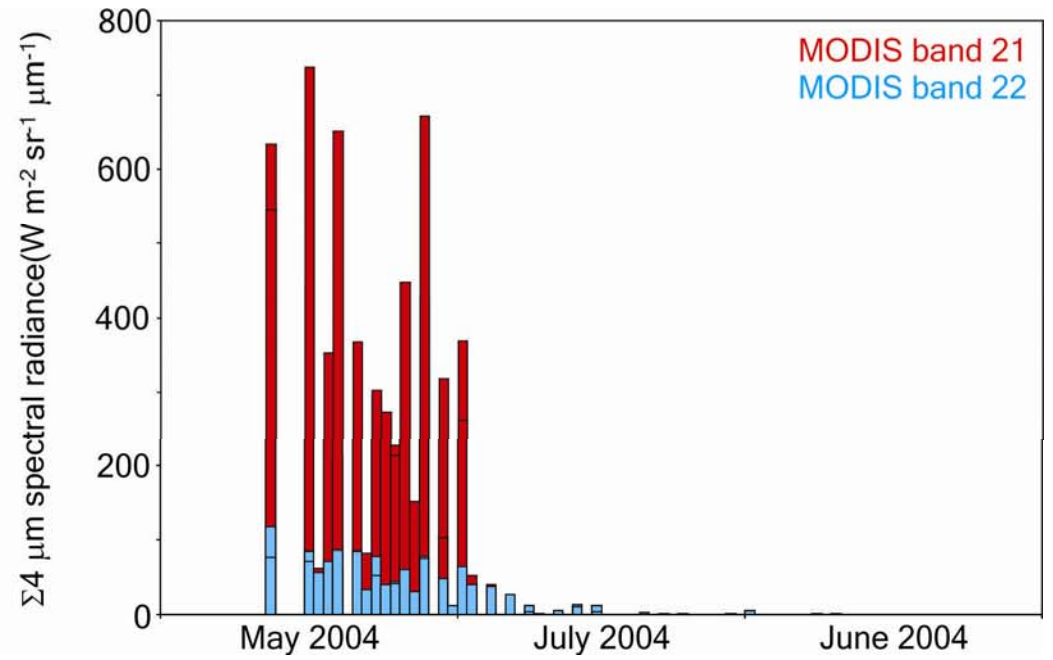


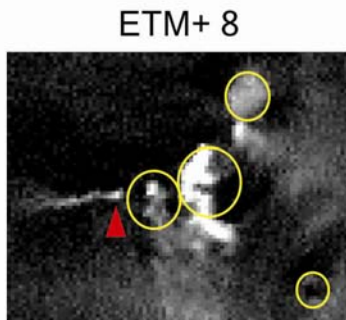
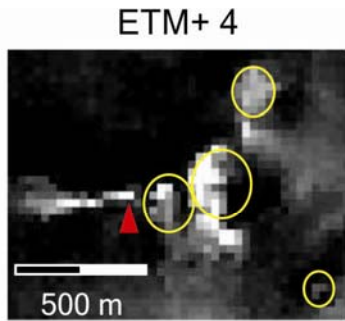
P. Izbekov (UAF/AVO)



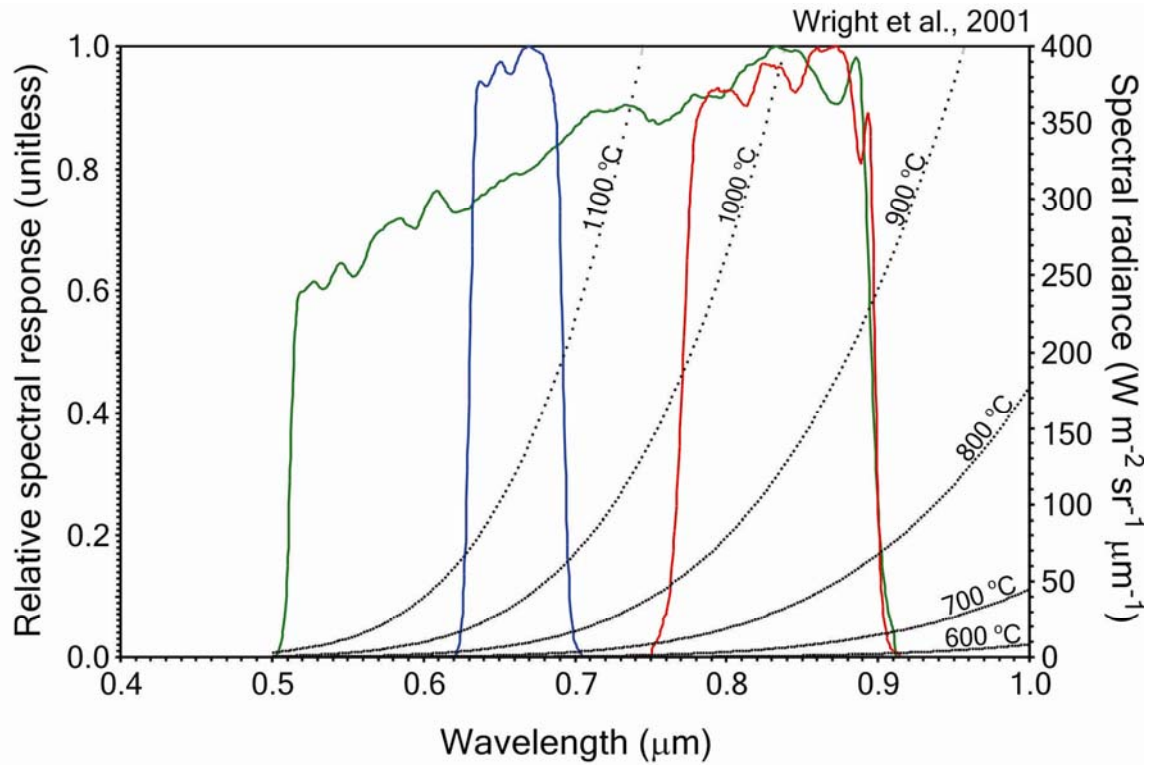


J. Durieux (©)





Mount Etna, 28 October 1999



$$\text{ETM+ 4} = 35.7 \text{ W m}^{-2} \text{ sr}^{-1} \mu\text{m}^{-1} = 819 \text{ }^{\circ}\text{C}$$

$$\text{ETM+ 8} = 28.3 \text{ W m}^{-2} \text{ sr}^{-1} \mu\text{m}^{-1} = 922 \text{ }^{\circ}\text{C}$$

A difference in whole pixel temperature of 100 °C

